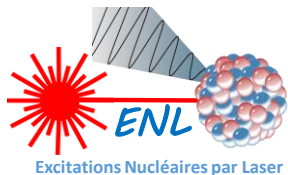


Introduction to Nuclear physics with high-power lasers

Medhi Tarisien



How a laser can have an effect on a nucleus?

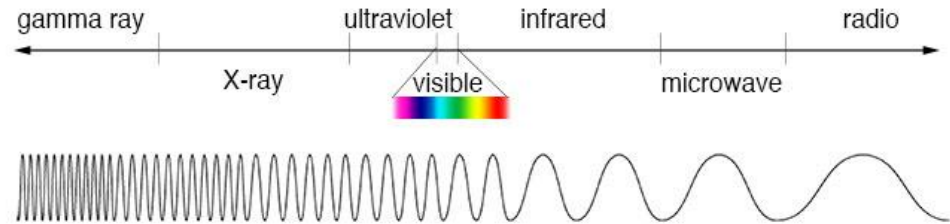
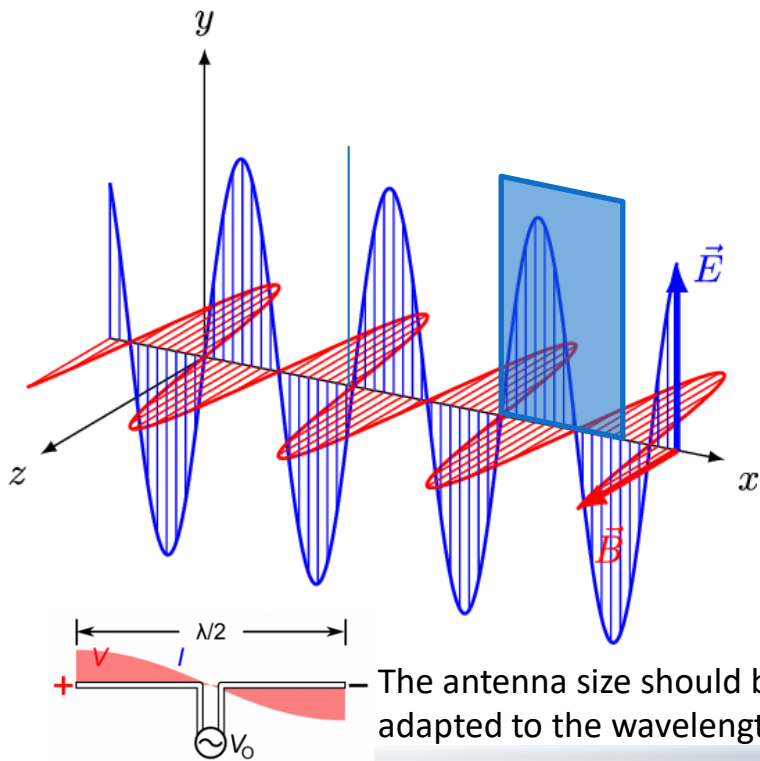
Electromagnetic radiation

λ_0 : Wavelength (m)

$c = 299\,792\,458$ m/s

ν : frequency (s^{-1})

T : period (s)

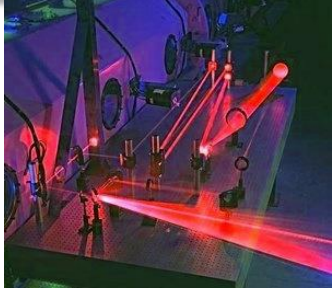


m	10^{-13}	10^{-11}	10^{-9}	10^{-7}	10^{-5}	10^{-3}	10^{-1}	10
cm (10^{-2} m)	10^{-11}	10^{-9}	10^{-7}	10^{-5}	10^{-3}	10^{-1}	10	10^3
nm (10^{-9} m)	10^{-4}	10^{-2}	1	10^2	10^4	10^6	10^8	10^{10}

Hz	10^{21}	10^{19}	10^{17}	10^{15}	10^{13}	10^{11}	10^9	10^7
----	-----------	-----------	-----------	-----------	-----------	-----------	--------	--------

eV	10^7	10^5	10^3	10	10^{-1}	10^{-3}	10^{-5}	10^{-7}
----	--------	--------	--------	----	-----------	-----------	-----------	-----------

How a laser can have an effect on a nucleus?



λ_0 : Wavelength (m)
 $c = 299\,792\,458$ m/s
 ν : frequency (s^{-1})
 T : period (s)

One photon

$$\lambda_0 = cT = \frac{c}{\nu}$$



The light we use : $\lambda_0 \approx 1\mu\text{m}$
 $T \approx 3$ fs

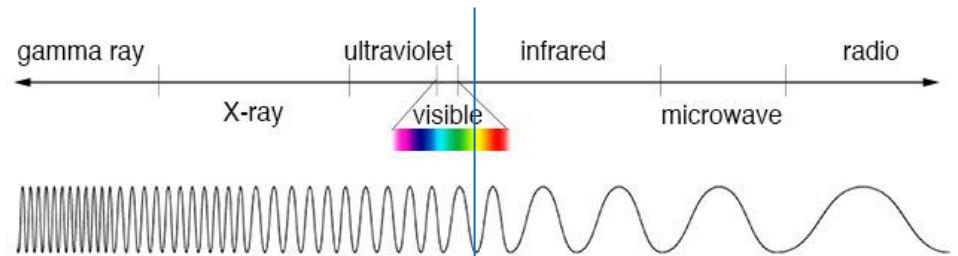
$$E = h \nu$$

$$h = 4.135667696 \times 10^{-15} \text{ eV}\cdot\text{s}$$

Planck constant

$$\Leftrightarrow 1.3 \text{ eV}$$

No direct effect of IR photon
on the nucleus !



m	10^{-13}	10^{-11}	10^{-9}	10^{-7}	10^{-5}	10^{-3}	10^{-1}	10
cm (10^{-2} m)	10^{-11}	10^{-9}	10^{-7}	10^{-5}	10^{-3}	10^{-1}	10	10^3
nm (10^{-9} m)	10^{-4}	10^{-2}	1	10^2	10^4	10^6	10^8	10^{10}

Wavelength

Hz	10^{21}	10^{19}	10^{17}	10^{15}	10^{13}	10^{11}	10^9	10^7
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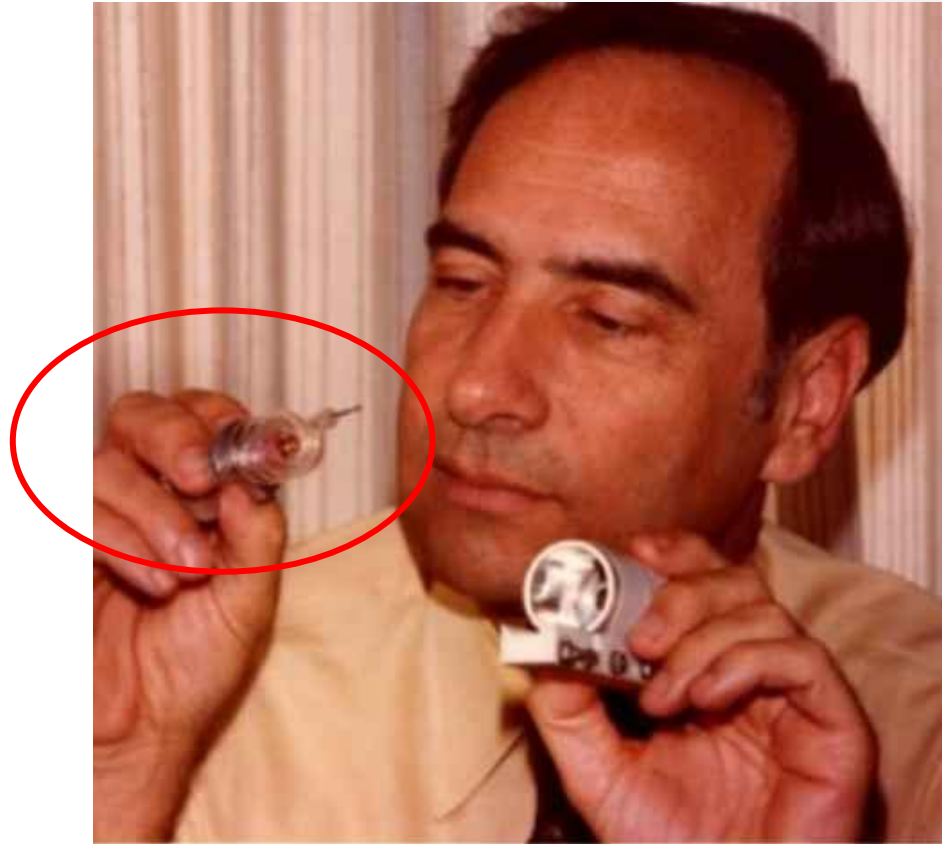
Frequency

eV	10^7	10^5	10^3	10	10^{-1}	10^{-3}	10^{-5}	10^{-7}
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Energy

How a laser can have an effect on a nucleus?

The laser



*Light
Amplification by
Stimulated
Emission of
Radiation*

Theodore Maiman : 1960 LASER rubis

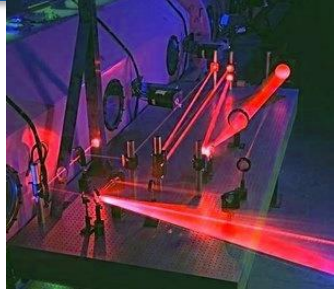
How a laser can have an effect on a nucleus?

The High power Laser



LULI
2000

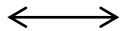
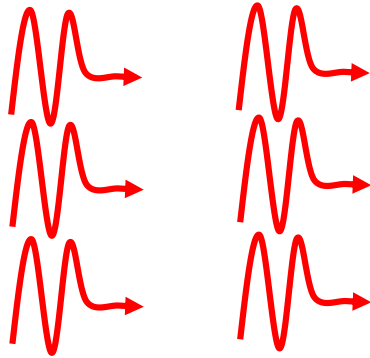
How a laser can have an effect on a nucleus?



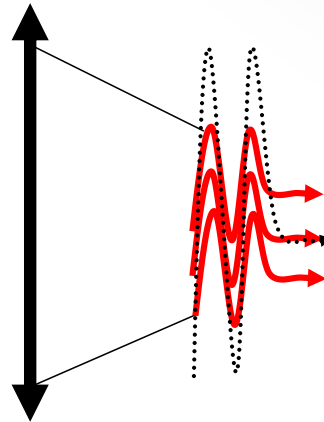
Plenty of photons



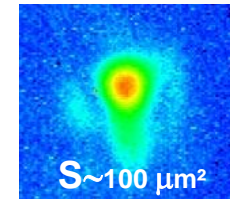
$\lambda \approx 1\mu\text{m}$
 $T \approx 3\text{ fs}$
 $\Leftrightarrow 1.3\text{ eV}$



Δt FWHM of a gaussian pulse



$$E_L = N_{\text{ph}} \times 1.3\text{ eV} \times 1.602 \cdot 10^{-19}\text{ J/eV}$$



$S \sim 100\ \mu\text{m}^2$
 Focal spot

The intensity

$$I = \frac{E_L}{\Delta t \cdot S}$$

Equivalent electric field :

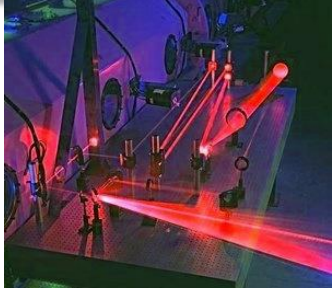
$$E(\text{V/cm}) = 27,446 \sqrt{I(\text{W/cm}^2)}$$

$$E_{\text{atom}} \approx 9 \cdot 10^9 \frac{1,6 \cdot 10^{-19}}{(10^{-10})^2} \approx 10^9 \text{ V/cm}$$

$$E_{\text{nucl}} \approx 9 \cdot 10^9 \frac{1,6 \cdot 10^{-19}}{(10^{-15})^2} \approx 10^{19} \text{ V/cm}$$

Which Intensities are required to have an effect on the atom? The nucleus?

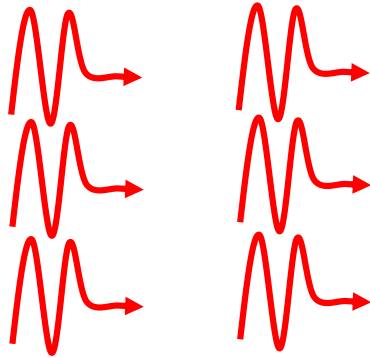
How a laser can have an effect on a nucleus?



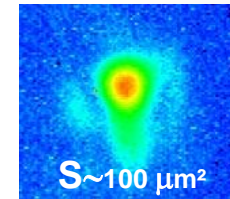
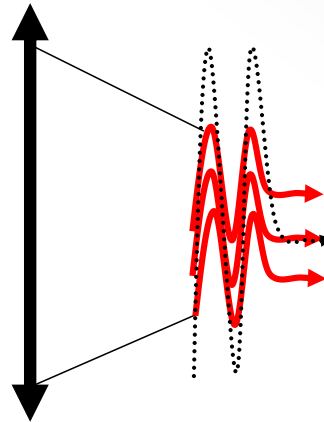
Plenty of photons



$\lambda \approx 1 \mu\text{m}$
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Δt FWHM of a gaussian pulse



Focal spot

The intensity

$$I = \frac{E_L}{\Delta t \cdot S}$$

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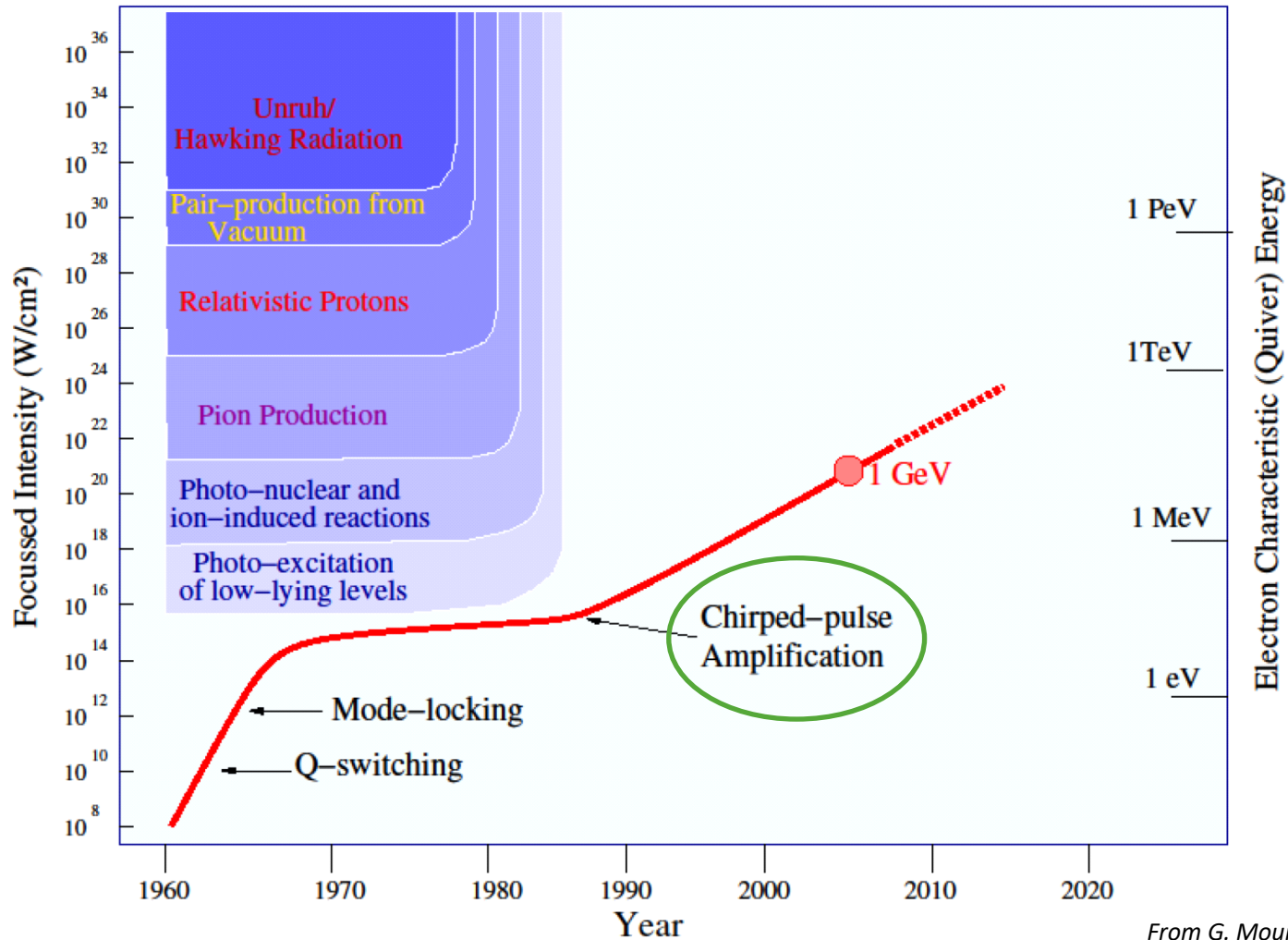
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$$E_{\text{nucl}} \approx 9 \cdot 10^9 \frac{1,6 \cdot 10^{-19}}{(10^{-15})^2} \approx 10^{19} \text{ V/cm}$$

$$\text{Atom : } I \approx 1.3 \times 10^{15} \text{ W/cm}^2$$

$$\text{Nucleus : } I \approx 1.3 \times 10^{35} \text{ W/cm}^2$$

High power lasers



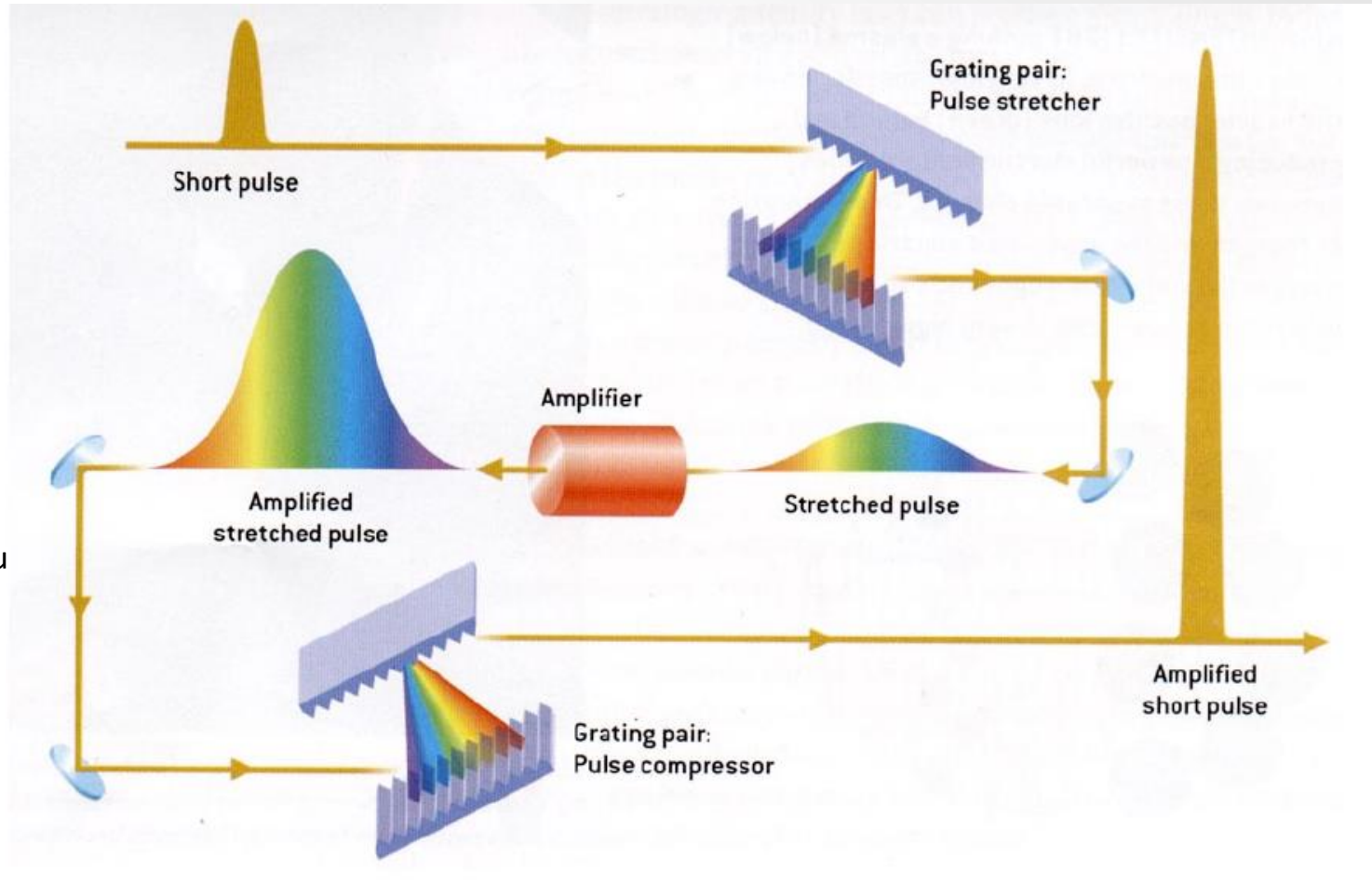
High power lasers

Chirped Pulse Amplification (CPA)

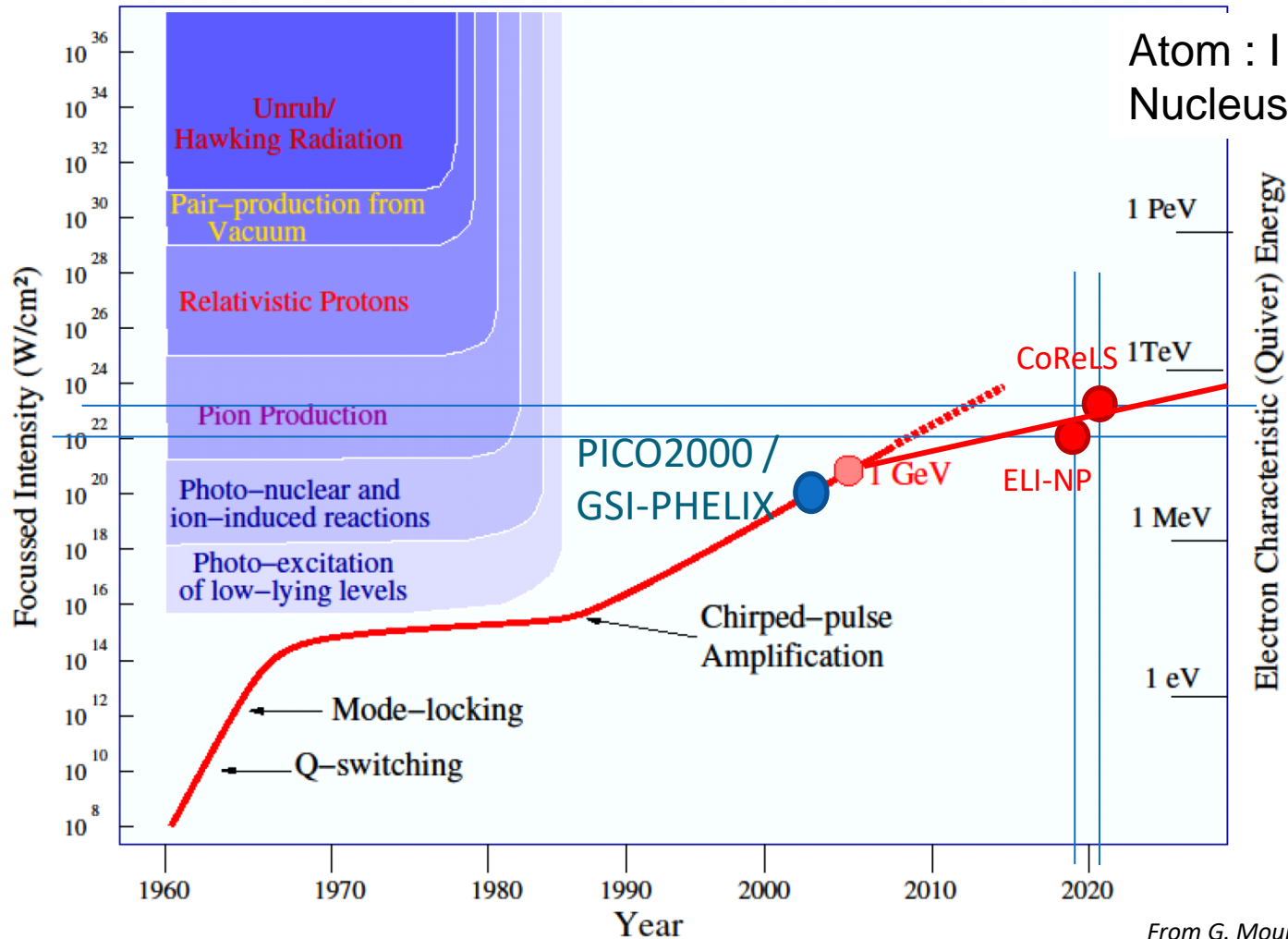


D. Strickland & G. Mourou

Nobel prize in
Physics 2018



High power lasers



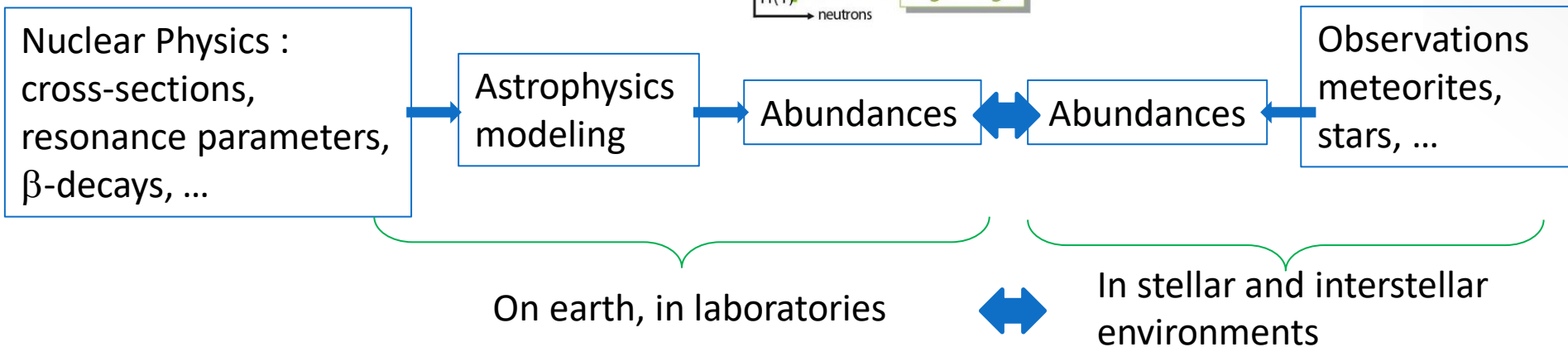
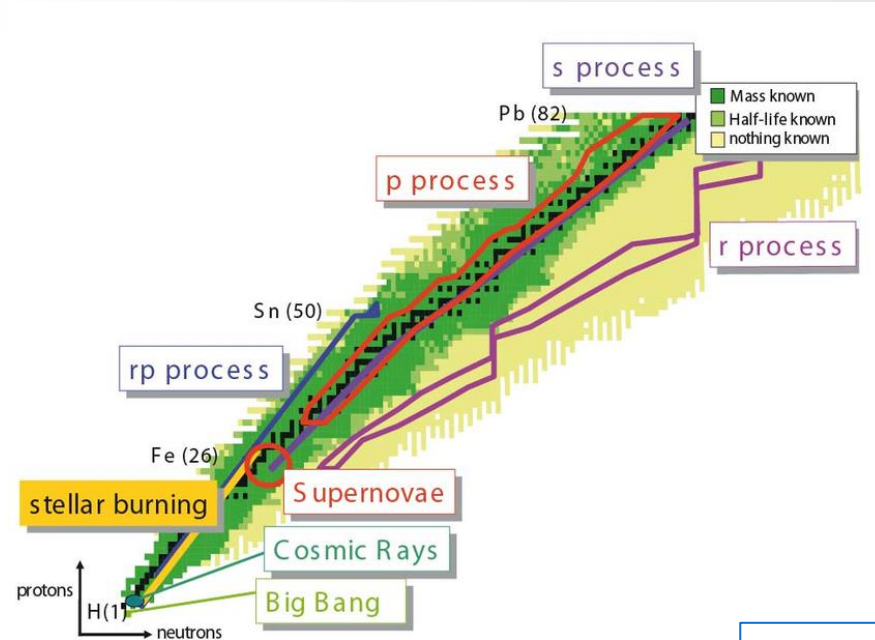
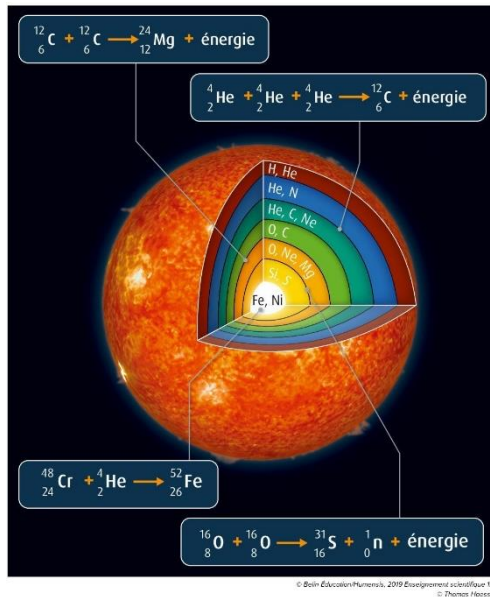
No direct effect of the electric field on the nucleus before a century !!!

BUT we can use the **interplay between atomic electrons and nucleus**



From G. Mourou

A contribution in Nucleosynthesis studies



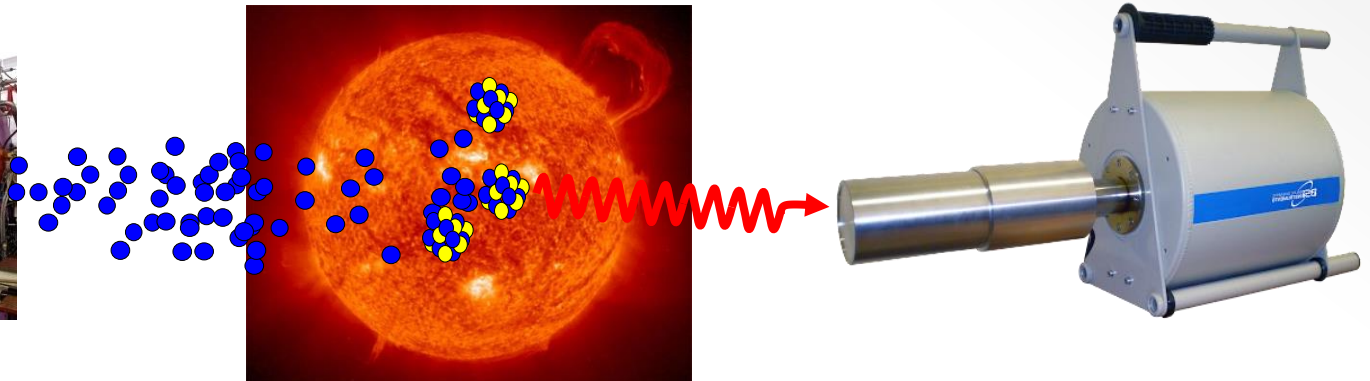
Are nuclear reactions the same in a star as in a solid?

Nuclear physics in stellar medium

Take a star in a laboratory as a target

Take an accelerator and send some particle projectiles on the star

Detect the nuclear reaction signatures



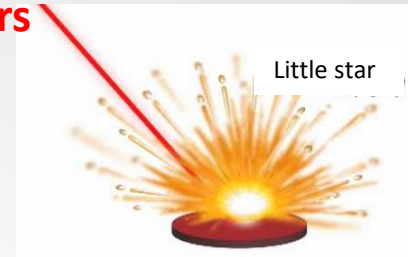
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Nuclear physics in stellar medium

High power lasers

Take a star in a laboratory as a target :

you can have one during ~ 1 ns only

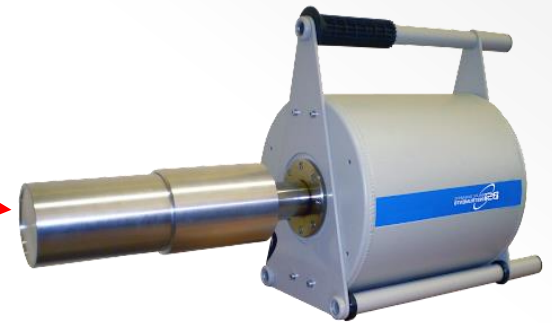
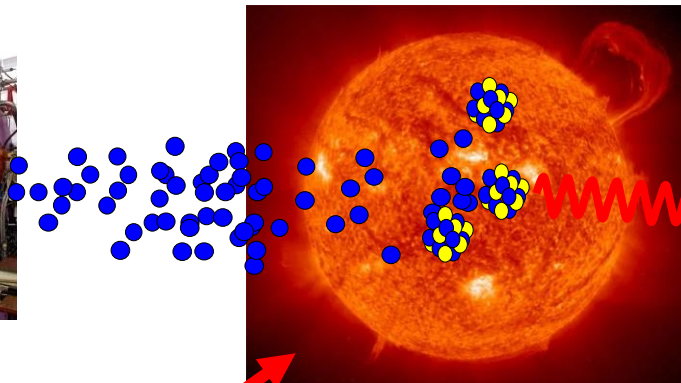


Take an accelerator and send some particle projectiles on the star :

What should be the intensity of a proton beam for 10^{13} protons passing through the plasma target? ($e = 1.6 \cdot 10^{-19}$ C)

$I = ?$

Detect the nuclear reaction signatures



High power lasers

Are nuclear reactions the same in a star as in a solid?

Nuclear physics in stellar medium

High power lasers

Take a star in a laboratory as a target :

you can have one during 1 ns only



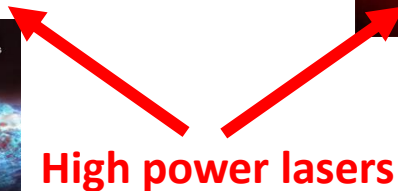
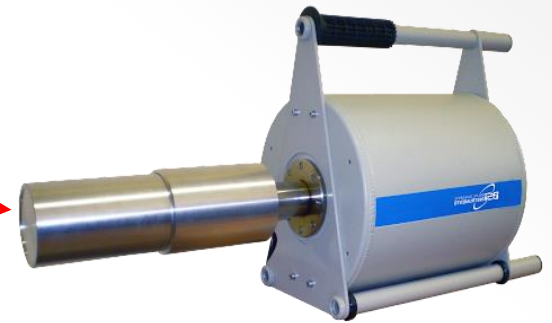
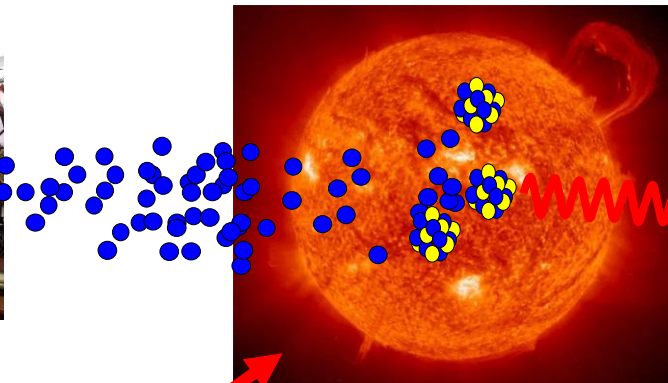
Take an accelerator and send some particle projectiles on the star :

What should be the intensity of a proton beam for 10^{13} protons passing through the plasma target? ($e = 1.6 \cdot 10^{-19} \text{ C}$)

$$I = 10^{13} \times 1.6 \cdot 10^{-19} \text{ C} / 10^{-9} \text{ s} = 1,6 \text{ kA}$$

accelerators Ultra-High Intensity : $\sim 100 \text{ mA}$

Detect the nuclear reaction signatures



High power lasers

Nuclear physics with high-power lasers

- High power lasers and their interaction with matter
- Laser-Plasma Acceleration
- Nuclear Physics in plasmas
- Challenges to take up

Part 1

HIGH POWER LASERS AND THEIR INTERACTION WITH MATTER

- High power laser characteristics
- What is a Plasma ?
- Laser/plasma interaction

High power laser characteristics

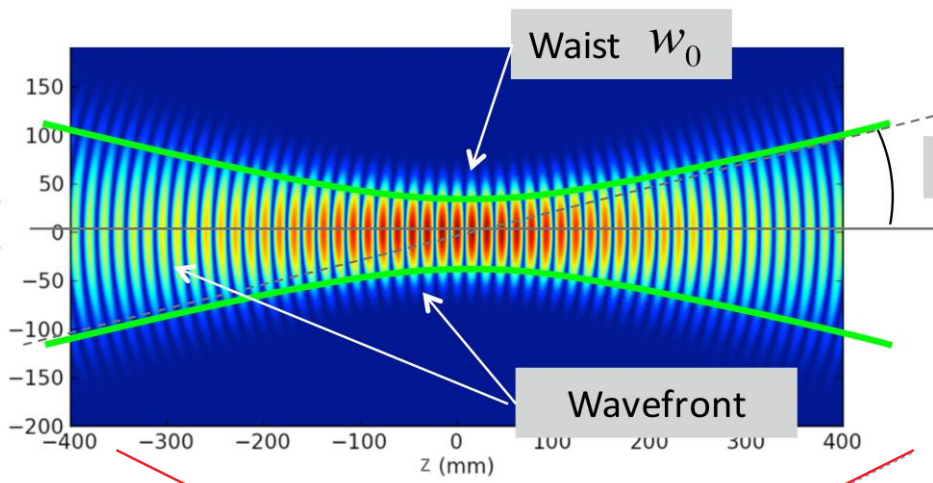
$$I = \frac{E_L}{\Delta t \cdot S}$$

Is it possible to decrease $S \rightarrow 0 \text{ cm}^2$ to increase $I \rightarrow \infty$?

High power laser characteristics

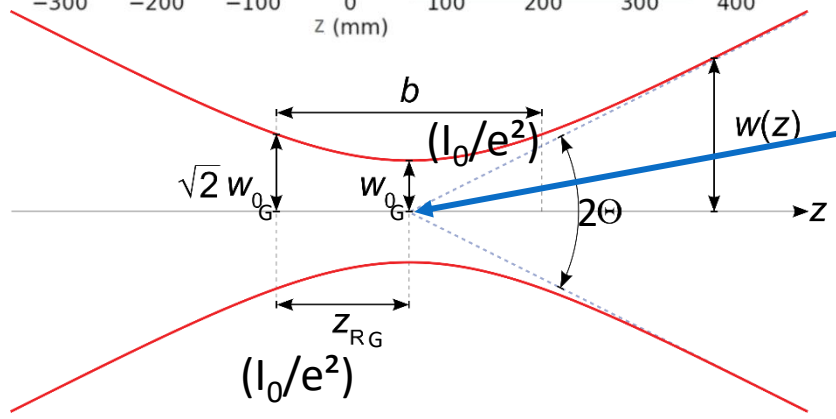
$$I = \frac{E_L}{\Delta t \cdot S}$$

Is it possible to decrease $S \rightarrow 0 \text{ cm}^2$ to increase $I \rightarrow \infty$?



Example of a focused **Gaussian** pulse
(30 000 times a laser size)

- $\lambda = 30 \text{ mm}$
- $\Delta t = 2.6 \text{ ns} \leftrightarrow 780 \text{ mm}$
- $\theta = 0.24 \text{ rad} = 13.7^\circ$
- $w_{0G} = 40 \text{ mm}$



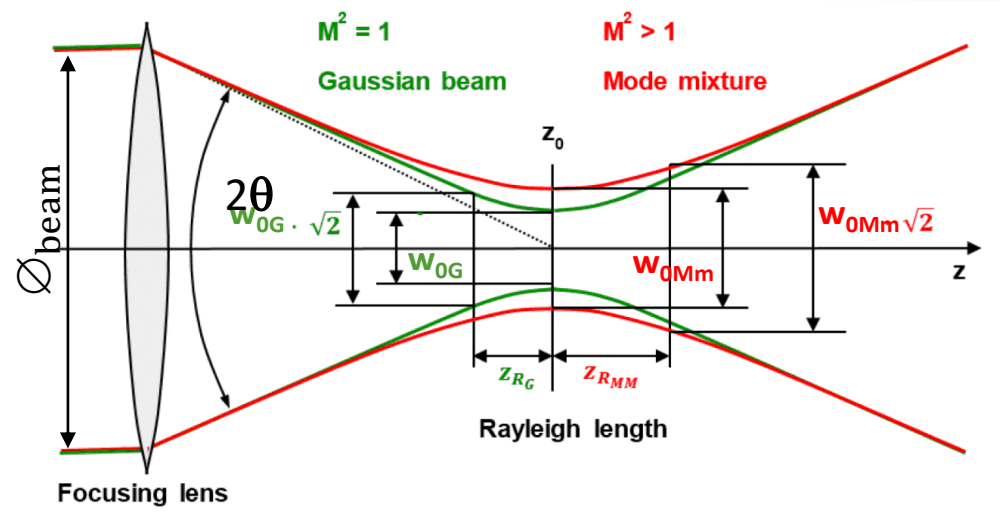
$$I_0 = \frac{2E_L}{\Delta t \times \pi w_0^2}$$

Field depth $b \leftrightarrow$ Rayleigh length

$$b = 2 \times Z_{RG} = 2 \times \pi \frac{w_{0G}^2}{\lambda_L}$$

Here $b = 334 \text{ mm}$

High power laser characteristics



$$\tan\theta = \frac{\varnothing_{\text{beam}}/2}{f}$$

$$w_{0MM} = \frac{M^2\lambda_L}{\pi\theta}$$

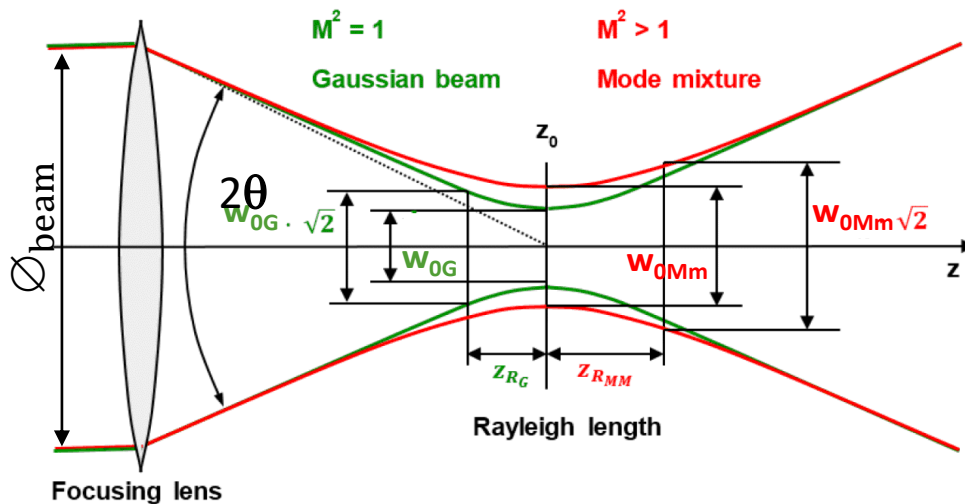
$$z_{RMM} = \pi \frac{w_0^2}{M^2\lambda_L}$$

$$I_0 = \frac{2E_L}{\Delta t \times \pi w_0^2}$$

- LULI-PICO2000 has the following characteristics :
 - $\lambda_L = 1.053 \mu\text{m}$, $M^2 = \sim 2$, $\varnothing_{\text{beam}} = 18 \text{ cm}$, $E_L = 90 \text{ J}$, $\Delta t = 1 \text{ ps}$
- It is focused with a parabola mirror of 800 mm focal distance

What is the ultimate focal spot diameter we can obtain? Which focal depth sensitivity? Which maximum intensity?

High power laser characteristics



$$\tan\theta = \frac{\varnothing_{\text{beam}}/2}{f}$$

$$w_{0MM} = \frac{M^2\lambda_L}{\pi\theta}$$

$$z_{RMM} = \pi \frac{w_0^2}{M^2\lambda_L}$$

$$I_0 = \frac{2E_L}{\Delta t \times \pi w_0^2}$$

- LULI-PICO2000 has the following characteristics :

$$\lambda_L = 1.053 \mu\text{m}, M^2 = \sim 2, \varnothing_{\text{beam}} = 18 \text{ cm}, E_L = 90 \text{ J}, \Delta t = 1 \text{ ps}$$

- It is focused with a parabola mirror of 800 mm focal distance

$$\tan\theta = \frac{180/2}{800} \quad \rightarrow \theta = 0.112 \text{ rad}$$

$$\pm z_{RMM} \approx \pi \frac{(6\mu\text{m})^2}{2 \times 1.053\mu\text{m}} \approx \pm 54 \mu\text{m}$$

$$w_{0MM} = \frac{2}{\pi \times 0.122} \times 1.053 \mu\text{m} \approx 6 \mu\text{m}$$

$$\varnothing_{\text{beam}} \approx 12 \mu\text{m}$$

$$I_0 \approx \frac{2 \times 90 \text{ J}}{10^{-12} \times \pi \times (6.10^{-4} \text{ cm})^2} \approx 1.6 \cdot 10^{20} \text{ W/cm}^2$$

High power laser characteristics

a_0 : the normalized laser vector potential

Is the ratio of classical speed of electron accelerated by the electric field of the laser in the middle of the waist over light celerity

$$a_0 = \frac{v}{c}$$

$0 \leq a_0 < 1$ classical electron, linear regime

$a_0 \gg 1$ relativistic electron, non linear regime

$a_0 \approx 1$: quasi-linear regime

$$a_0 = \sqrt{\frac{e^2 I \lambda_L^2}{2\pi^2 \epsilon_0 m_e^2 c^5}} \approx 0.85 \left(\frac{I \lambda_{L\mu m}^2}{10^{18} \text{Wcm}^{-2} \mu m^2} \right)^{1/2}$$

→ L. Gremillet presentation

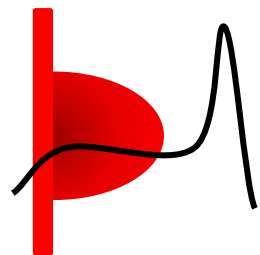
High power laser characteristics

- A laser pulse is preceded by a pedestal : Amplified Spontaneous Emission (ASE)
- The ration between the pre-pulse intensity and the main pulse one is called Contrast

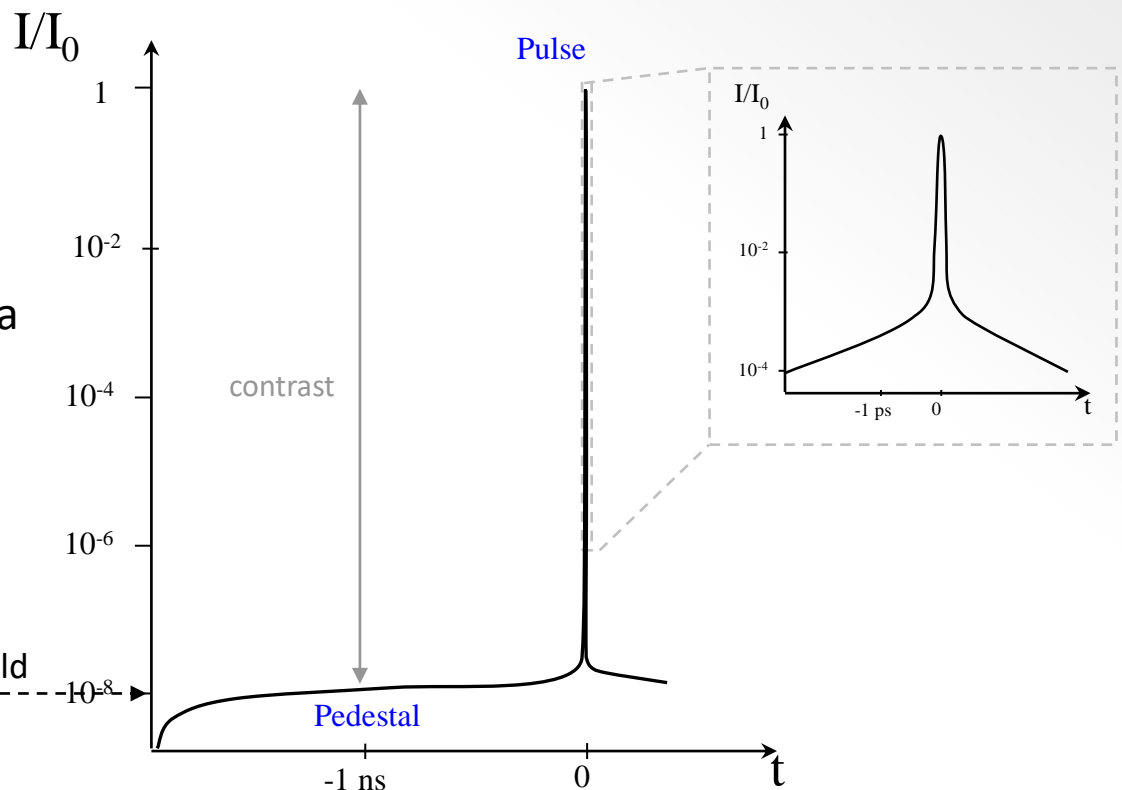
$$I_0 = \frac{2E_L}{\Delta t \times \pi w_0^2}$$

$$I_0 \sim 10^{20} \text{ W/cm}^2$$

The pedestal create a pre-plasma

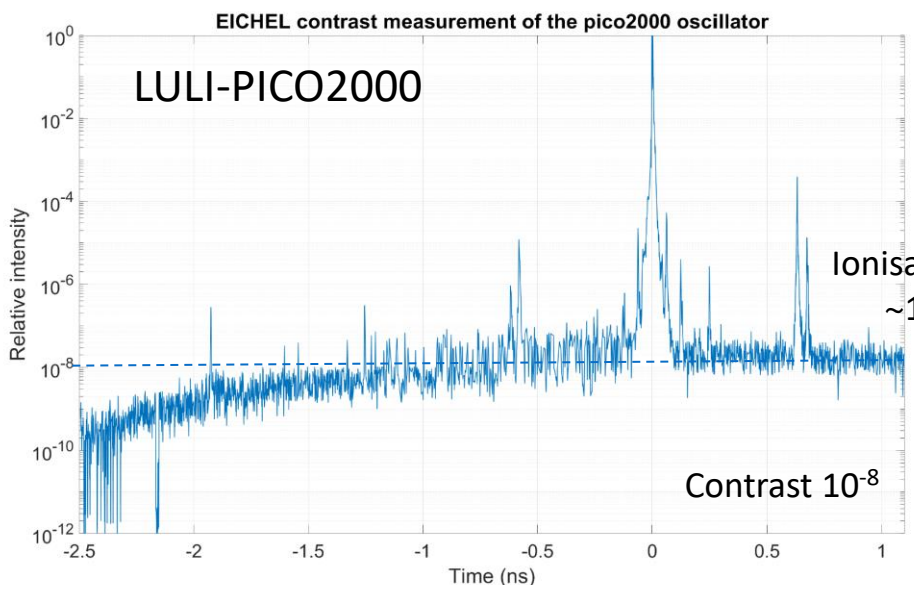


Ionisation threshold
 $\sim 10^{12} \text{ W.cm}^{-2}$

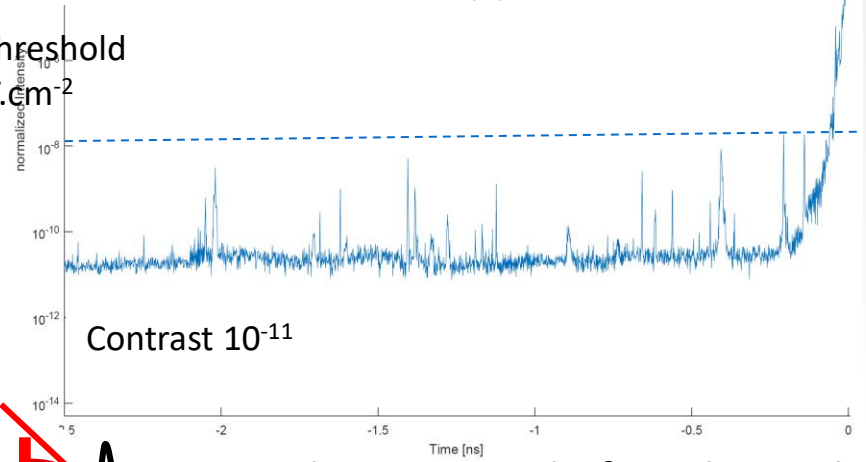
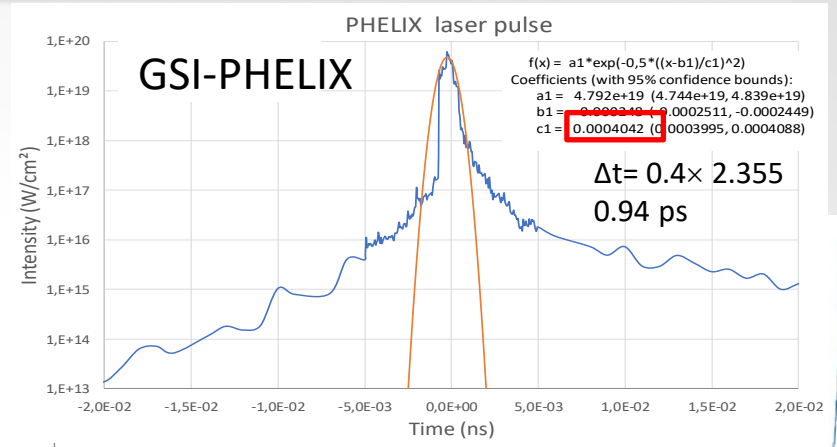



High power laser characteristics


$$I_{\max} = 10^{20} \text{ W/cm}^2 ; E \sim 50 \text{ J} ; \Delta t = 1 \text{ ps}$$



Ionisation threshold
 $\sim 10^{12} \text{ W.cm}^{-2}$

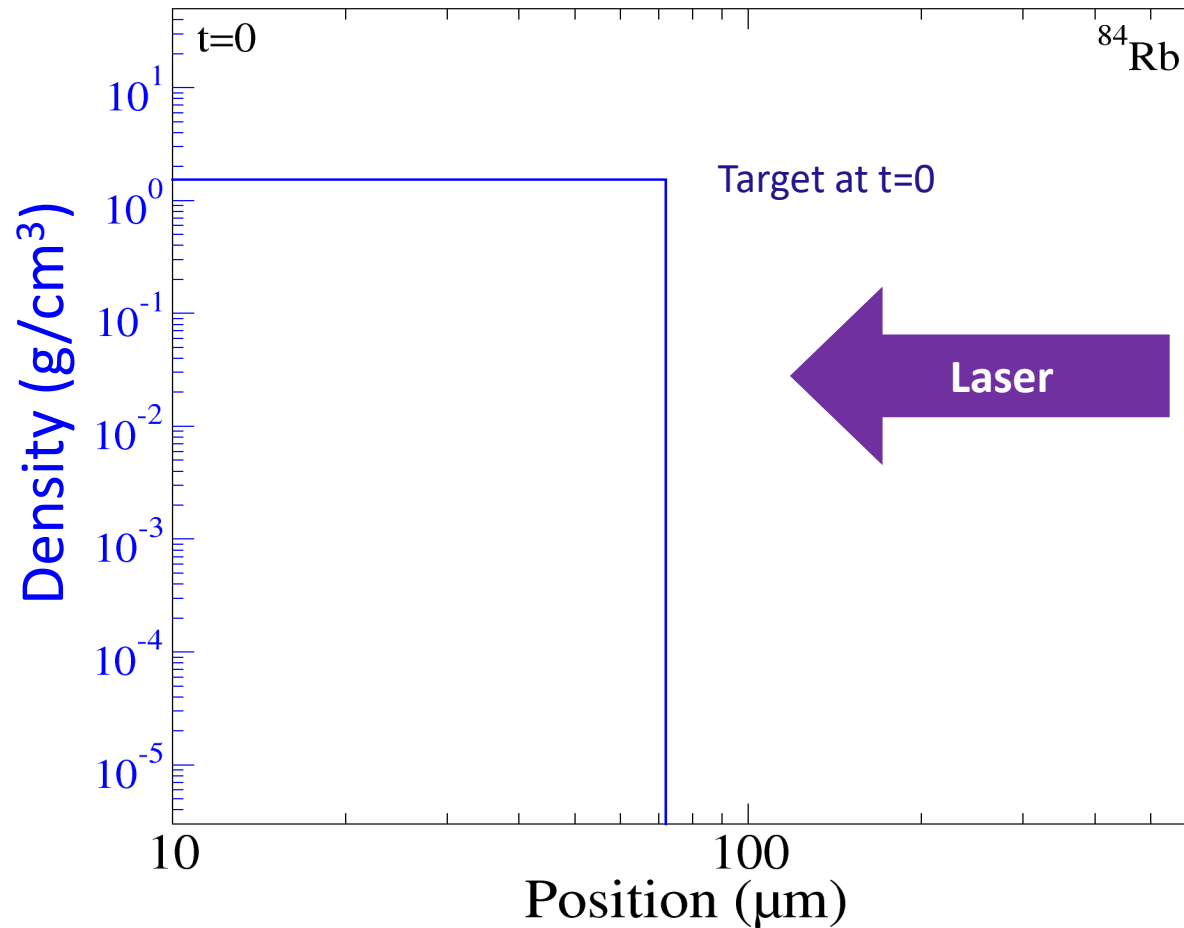


 Pre-plasma >1 ns before the pulse

~~~~ No pre-plasma >1 ns before the pulse

Every research laser facility is unique and has its own characteristics

Laser / solid target Interaction

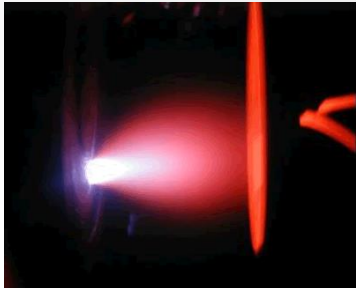
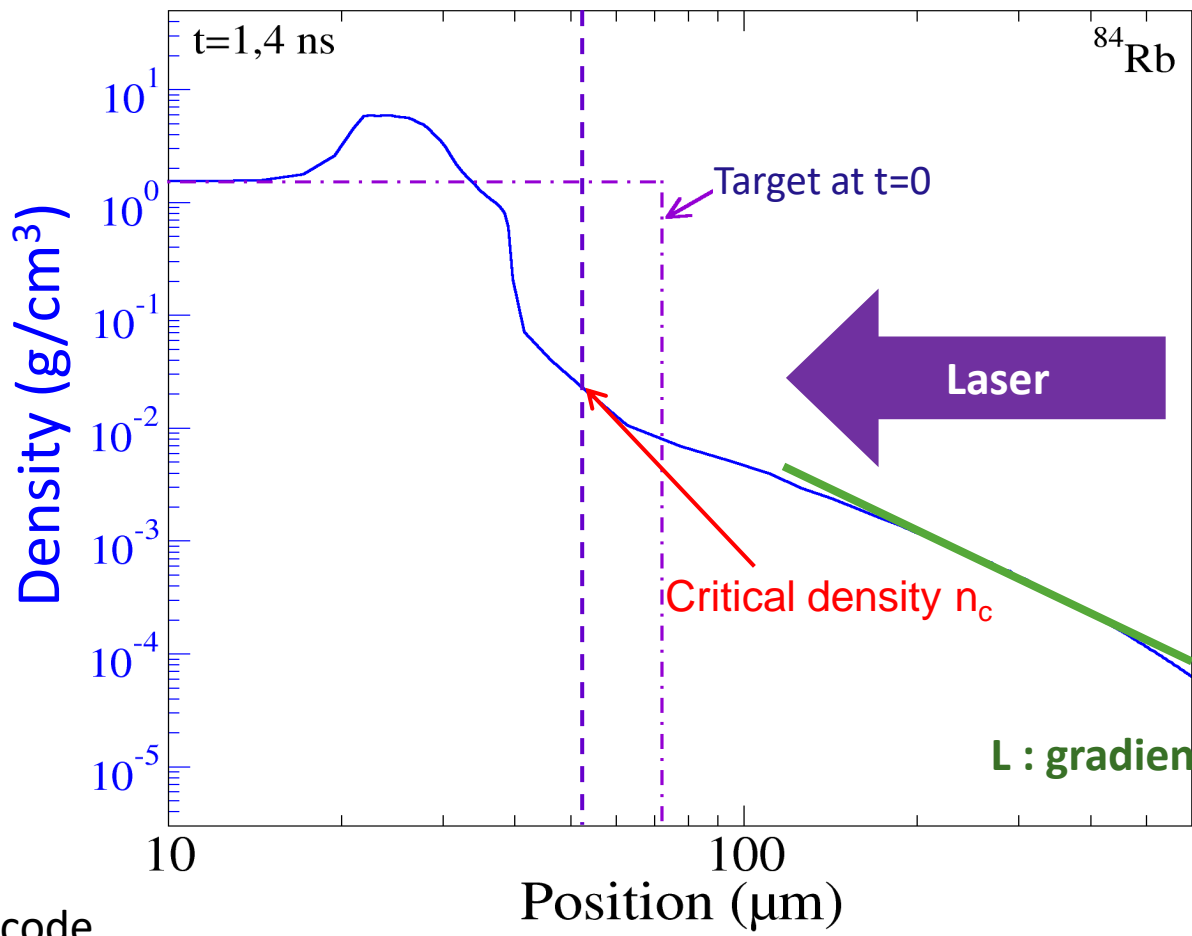


$$I = 10^{14} \text{ W/cm}^2$$
$$a_0 = 9 \cdot 10^{-3}$$

Laser / solid target Interaction

$$I = 10^{14} \text{ W/cm}^2$$

$$a_0 = 9 \cdot 10^{-3}$$

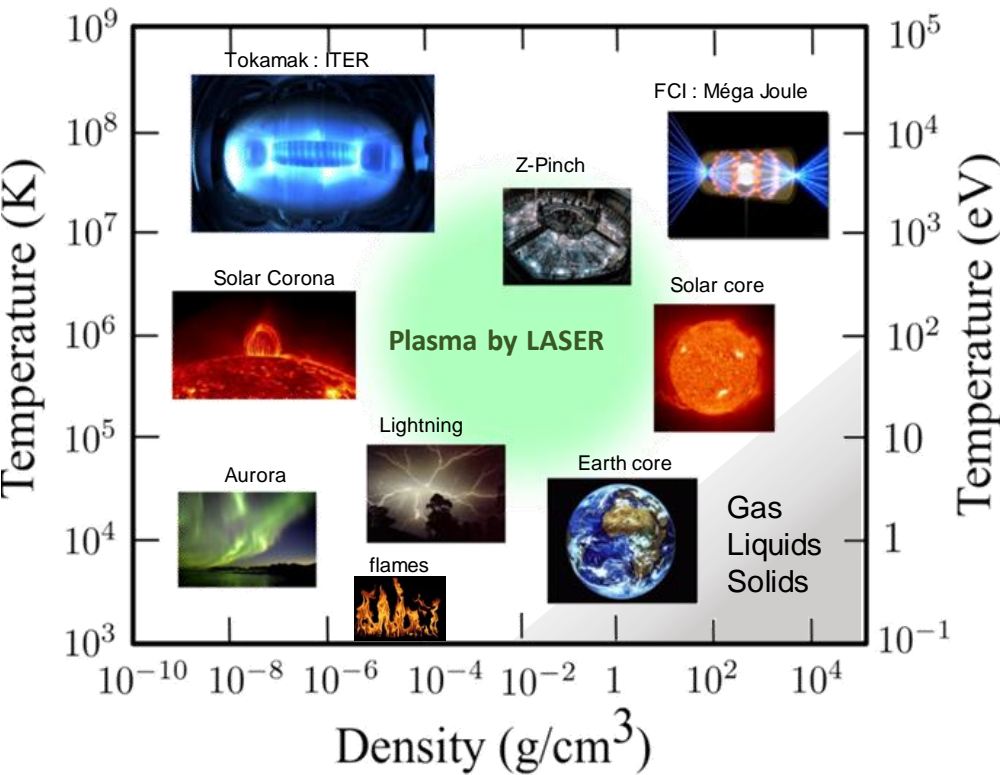


Chivas code

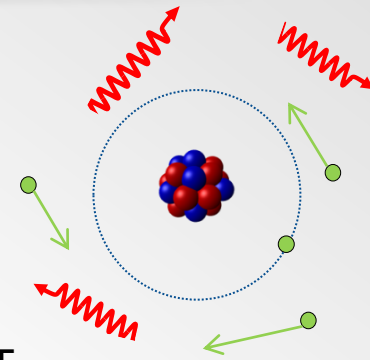
The high power laser pulses always interact with a plasma!

What is a plasma?

The 4th state of matter (99.9% of the visible matter)



- Multicharged ions
- Free electrons
- Photons



3 temperatures $T_e ; T_{ion} ; T_{ph}$

Different models to describe the plasma :

- **Thermodynamic Equilibrium:**
 $T_e = T_{ion} = T_{ph}$ (never reached)
- **Local Thermodynamic Equilibrium :**
 $T_e = T_{ion} \neq T_{ph}$ (Locally reached in dense plasmas)
- **Non Local Thermodynamic Equilibrium**
 $T_e \neq T_{ion} \neq T_{ph}$

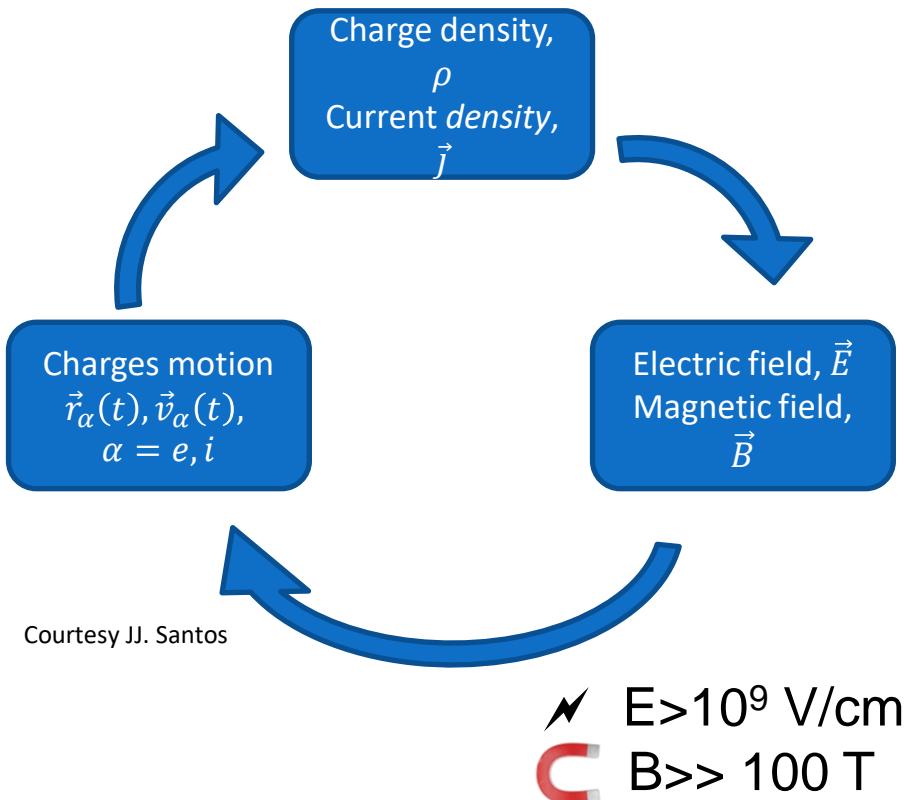
$$\overline{Q_{NETL}}(\rho, T_e) = \overline{Q_{ETL}}(\rho, T_Z)$$

Ionization Temperature T_Z

What is a plasma?

General behavior of a plasma

Plasma: a **quasi-neutral** gas of charged (electrons + ions, $n_i \approx n_e$) and neutral particles (atom less than 1% + photons) which exhibit collective behavior

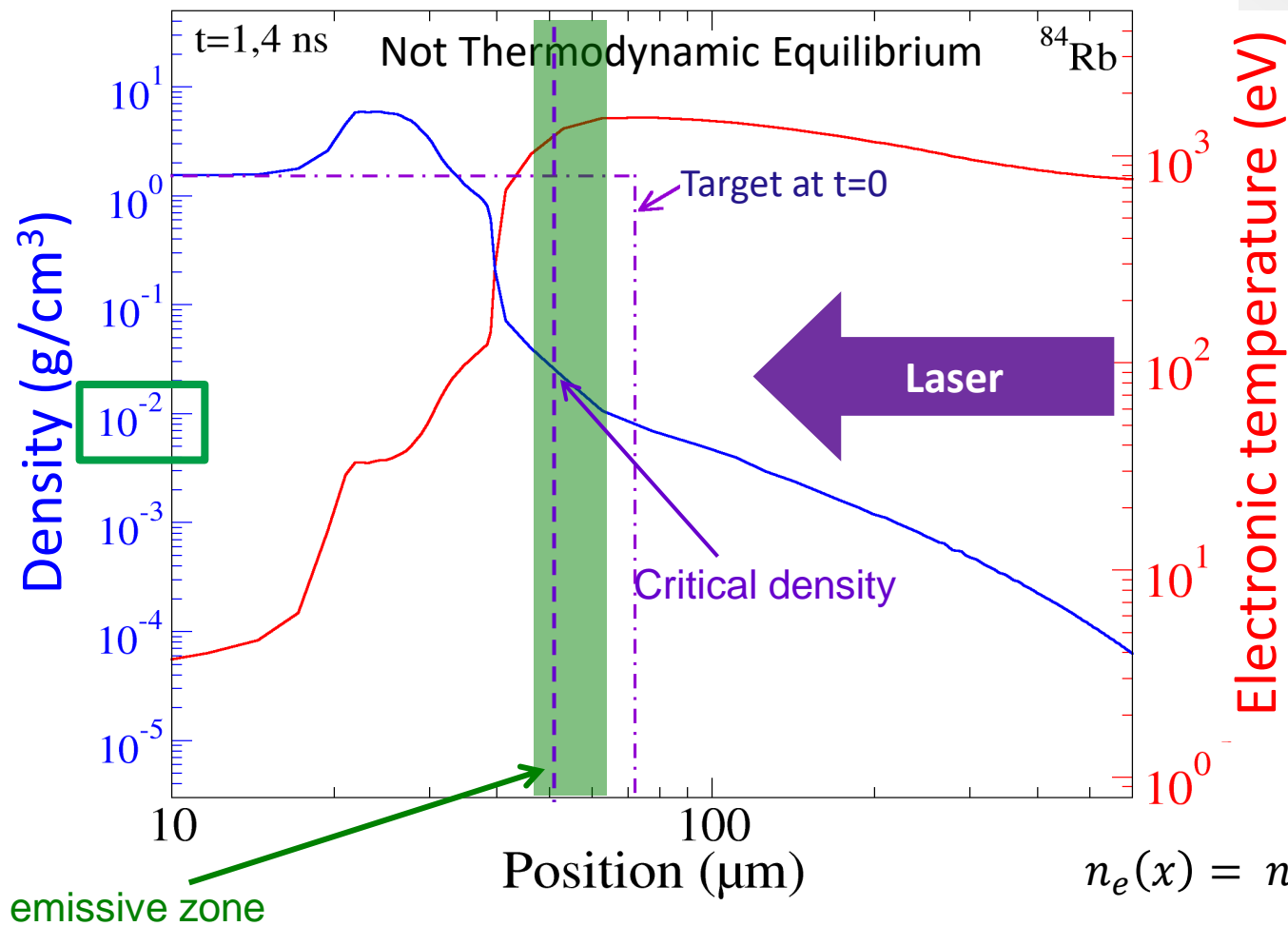


Courtesy JJ. Santos

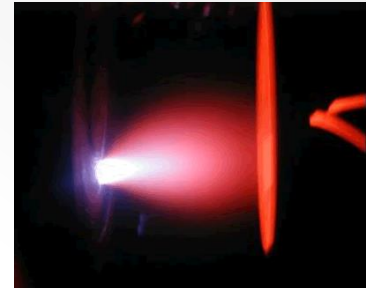
- Charges move and generate local concentration of + or - charges → **E-fields**.
Also, motion of charges generates currents → **B-fields**.
- Fields affect motion of other particles far away
→ **collective behavior**
- Particles (charges) exchange momentum and energy through collisions. System tends to converge to equilibrium (**Maxwell distributions**, thermalisation).
- $m_e \ll m_i$ → much shorter time-scale for electron dynamics.
Ions follow by electrostatic effect on **slower time scale**.
- According to temperature, density and the time-scale, plasma dynamics are described **kinetically** (Particles-In-Cell - PIC) or as **fluids** (MagnetoHydroDynamics).

What is a plasma?

Fluid (MagnetoHydroDynamics).



$I = 10^{14} \text{ W/cm}^2$
 $a_0 = 9 \cdot 10^{-3}$

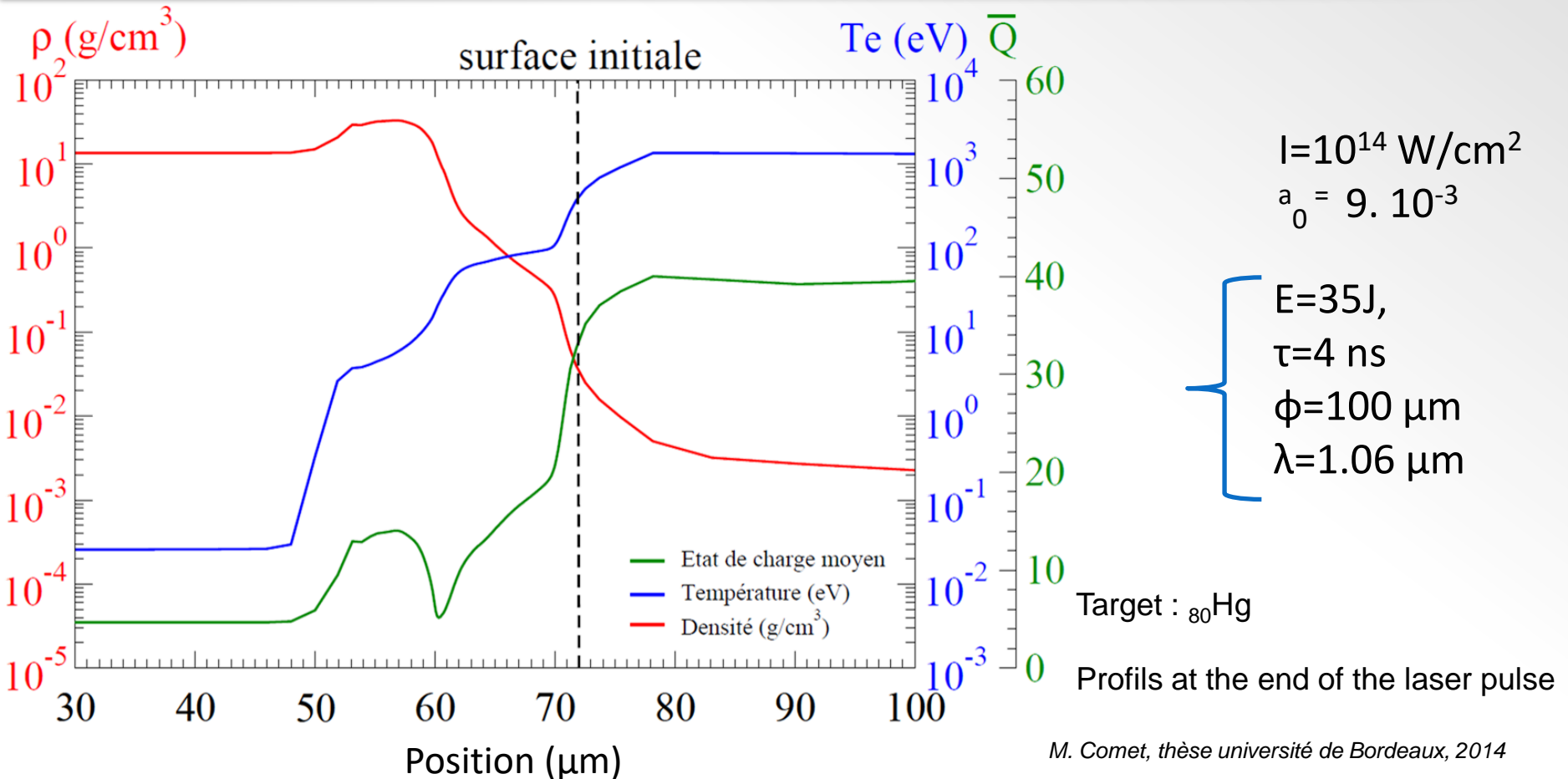


Most emissive zone

Requires ns to create a dense and hot plasma

What is a plasma?

Fluid (MagnetoHydroDynamics).



- Highly mean charge states reached → lot of atomic transitions
- Multiple collisions $\Delta t \Delta E \geq h/4\pi$ → resonant phenomena are enhanced

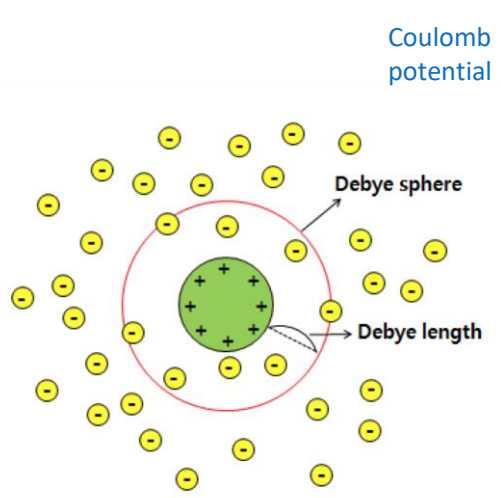
What is a plasma?

Particles

Spatial and time scales

→ Debye length:

Spatial-scale of deviation from electric neutrality around a multicharged ion



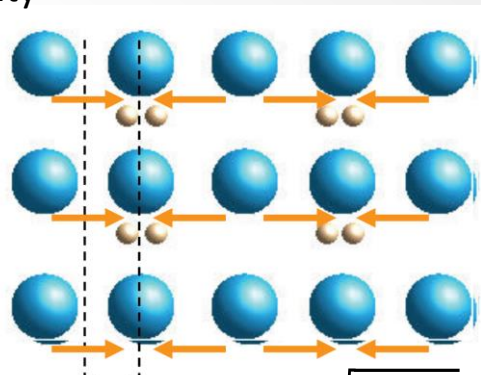
$$V(r) = \frac{Qe}{4\pi\epsilon_0} e^{-\frac{r}{\lambda_D}}$$

Plasma screening

$$\lambda_D = \sqrt{\frac{\epsilon_0 k_B T_e}{e^2 n_e}}$$

→ Plasma pulsation frequency:

(inverse of) time-scale of deviation from electric neutrality



$$\omega_{pe} = \sqrt{\frac{e^2 n_e}{\epsilon_0 m_e}}$$

→ Critical density n_c :

If electrons oscillate faster than the laser, it can not propagate inside : $\omega_{pe} > \omega_L$

$$n_c = \epsilon_0 m_e \frac{\omega_L^2}{e^2}$$

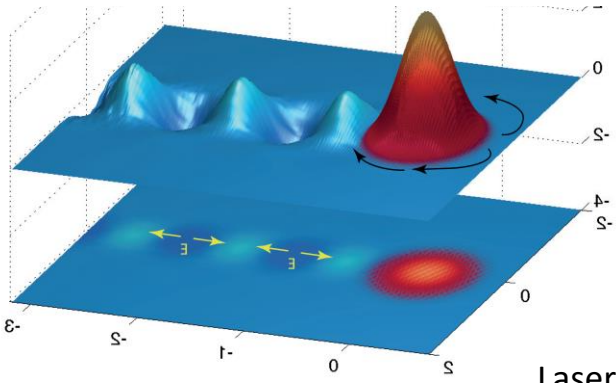
In relativist regime : $n_{c-rel} = \gamma n_c$

For PICO2000 $\lambda_L = 1052 \text{ nm}$,
 $n_c = \gamma 10^{21} \text{ e/cm}^3$

Laser / plasma interaction

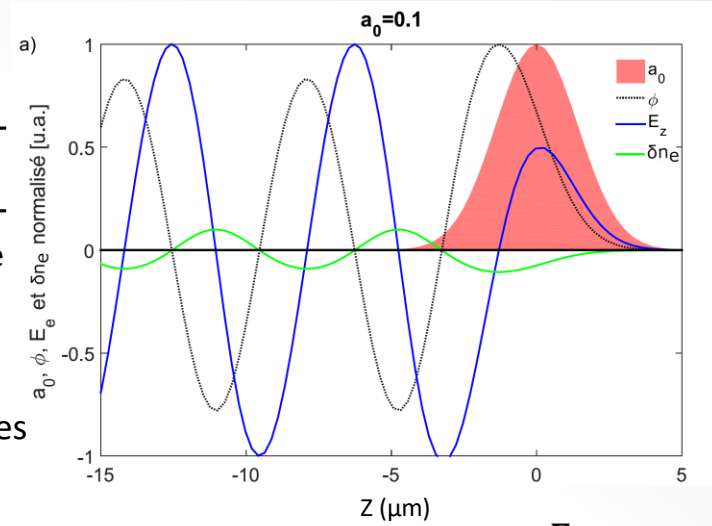
Particles In Cell

Lasers do not move ions significantly, but they move electrons (ions too heavy)



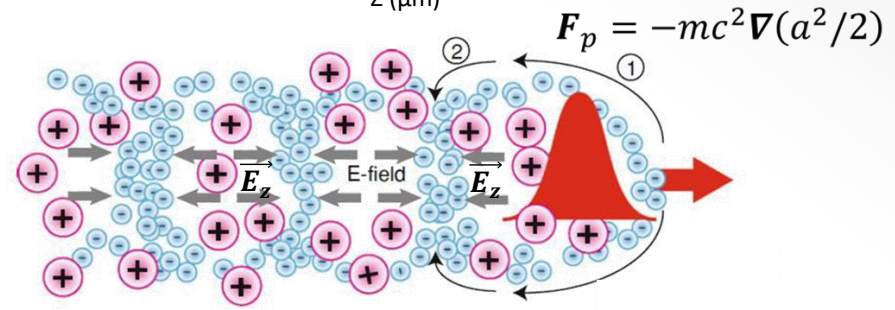
$$\omega_{pe} = \sqrt{\frac{e^2 n_e}{\epsilon_0 m_e}}$$

Laser-induced plasma waves



Phase velocity : $v_{pe} = \sqrt{1 - \frac{\omega_{pe}^2}{\omega_L^2}} c$

Plasma wavelength: $\lambda_{pe} = v_{pe} \frac{2\pi}{\omega_{pe}}$



PICO2000 : $\lambda_L = 1052 \text{ nm}$ in a plasma density of 10^{19} e-/cm^3 .

→ $v_{pe} = 99.5\% \times c$

→ $\lambda_{pe} \approx 10.50 \mu\text{m}$

Part 2

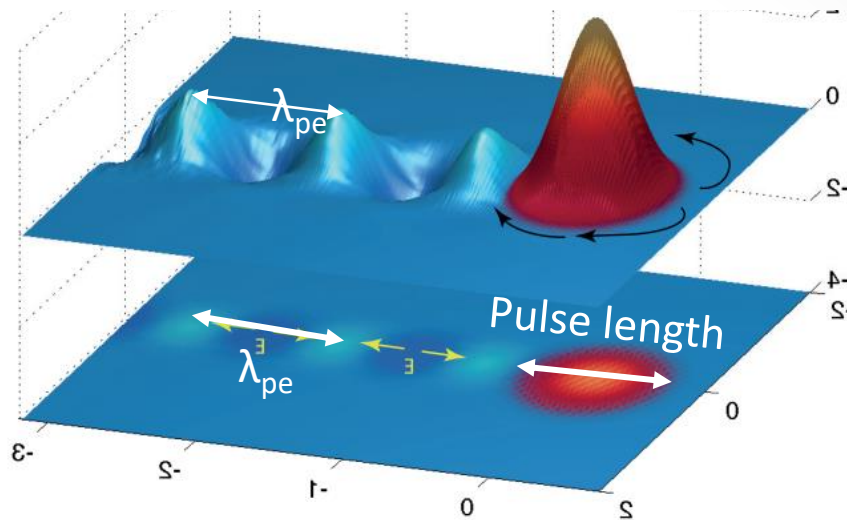
LASER-PLASMA ACCELERATION

- Electrons
- Photons
- Ions
- Neutrons

Electron source

Particles In Cell

Lasers do not move ions significantly, but they move electrons (ions too heavy)



$$\omega_{pe} = \sqrt{\frac{e^2 n_e}{\epsilon_0 m_e}}$$

$$v_{pe} = \sqrt{1 - \frac{\omega_{pe}^2}{\omega_L^2}} c$$

$$\lambda_{pe} = v_{pe} \frac{2\pi}{\omega_{pe}}$$

LULI-PICO2000 : $\lambda_L = 1052 \text{ nm}$; $\Delta t = 1 \text{ ps}$ in a plasma density of $10^{19} \text{ e}^-/\text{cm}^3$.

→ $v_{pe} = 99.50\% \times c$

→ $\lambda_{pe} \approx 10.5 \text{ } \mu\text{m}$

→ Pulse length = $298 \text{ } \mu\text{m}$

LULI-APOLLON : $\lambda_L = 840 \text{ nm}$; $\Delta t = 30 \text{ fs}$ in a plasma density of $10^{19} \text{ e}^-/\text{cm}^3$.

→ $v_{pe} = 99.68\% \times c$

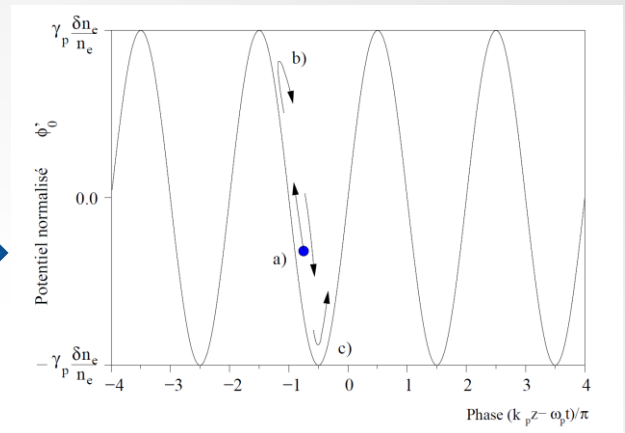
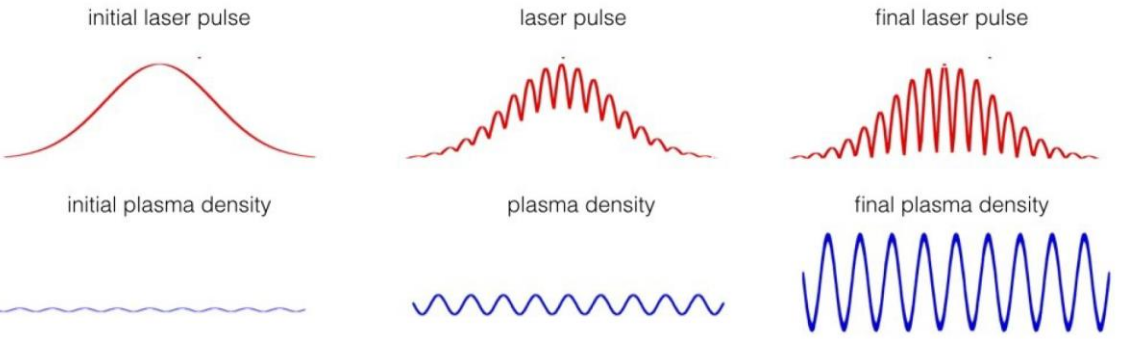
→ $\lambda_{pe} \approx 11 \text{ } \mu\text{m}$

→ Pulse length = $9 \text{ } \mu\text{m}$

Electron source

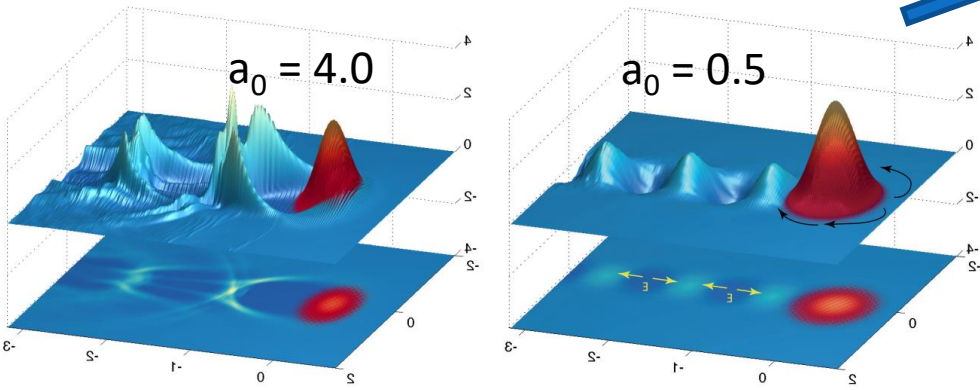
Long pulse compared to λ_{pe}

Self Modulated Laser Wake Field Acceleration process – SMLWFA

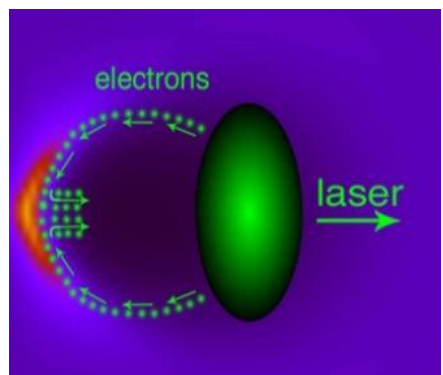


Maxwellian energy distribution

short pulse compared to λ_{pe}

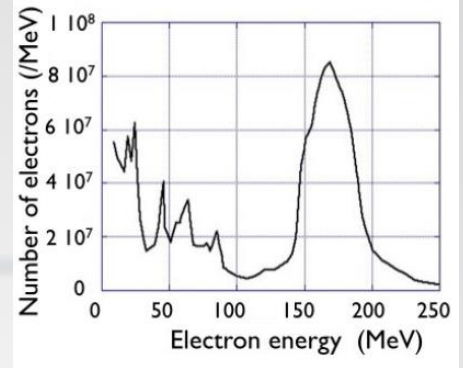


Bubble regime



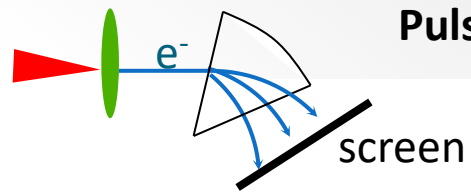
peaked energy distribution

Electron source



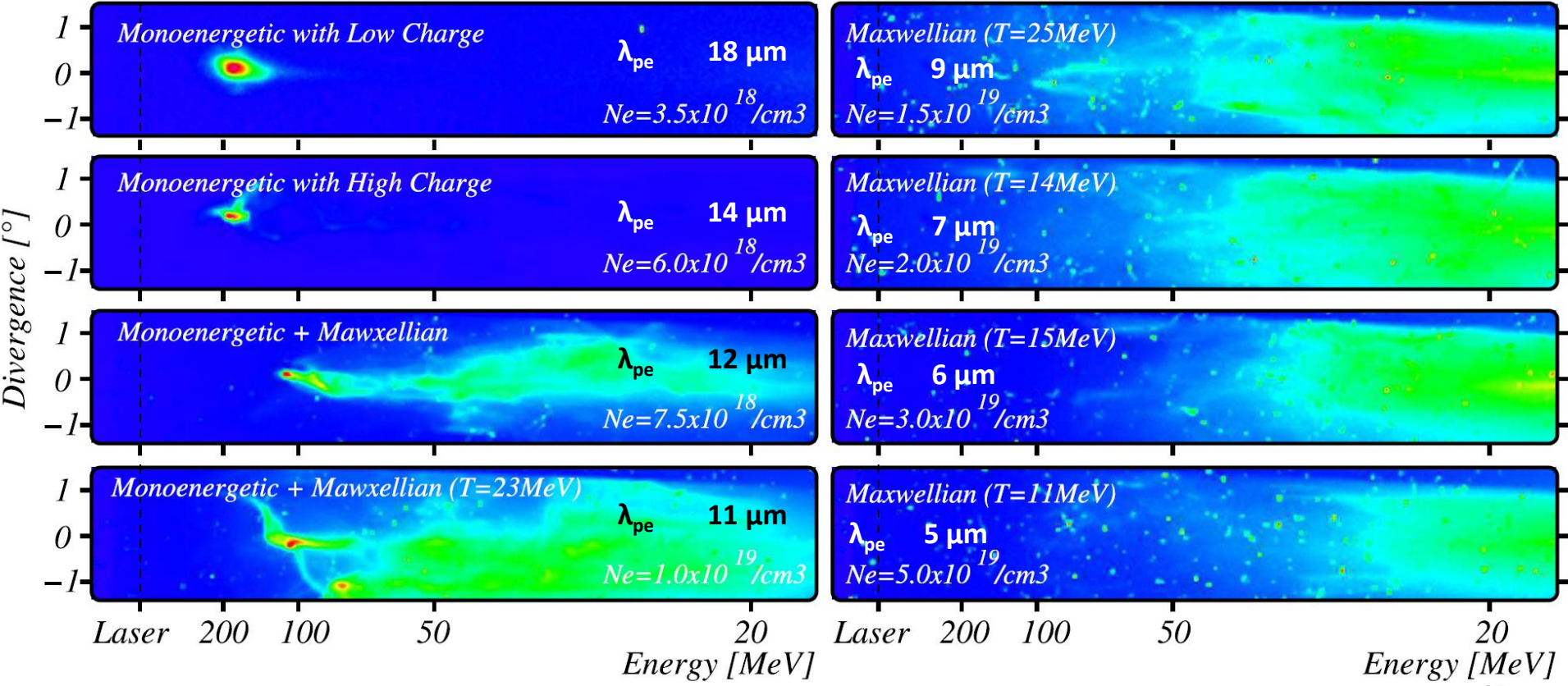
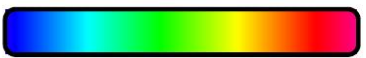
LOA « Salle Jaune » : Ti :Saphir $\lambda_L = 820 \text{ nm}$, $\Delta t = 30 \text{ fs}$ $a_0 = 1.3$

V. Malka et al, *Physics of Plasmas* 12, 056702 (2005)



Pulse length : 9 μm

Arbitrary Unit



Electron source

PHYSICAL REVIEW LETTERS 122, 084801 (2019)

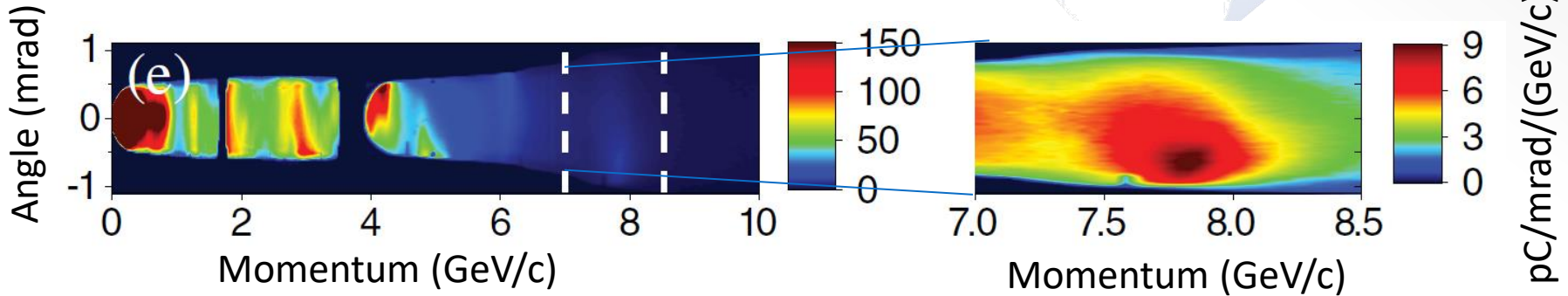
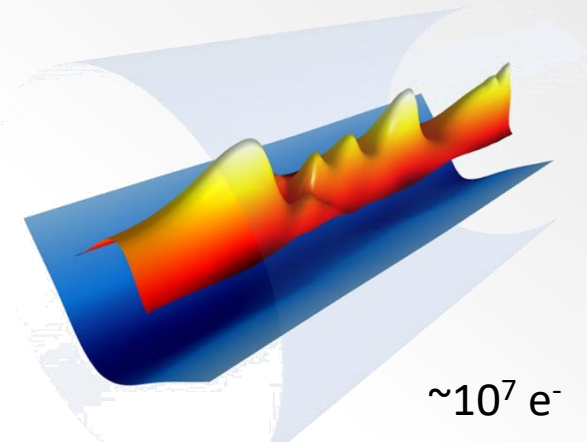
Editors' Suggestion Featured in Physics

Petawatt Laser Guiding and Electron Beam Acceleration to 8 GeV in a Laser-Heated Capillary Discharge Waveguide

A. J. Gonsalves,^{1,*} K. Nakamura,¹ J. Daniels,¹ C. Benedetti,¹ C. Pieronek,^{1,2} T. C. H. de Raadt,¹ S. Steinke,¹ J. H. Bin,¹ S. S. Bulanov,¹ J. van Tilborg,¹ C. G. R. Geddes,¹ C. B. Schroeder,^{1,2} Cs. Tóth,¹ E. Esarey,¹ K. Swanson,^{1,2} L. Fan-Chiang,^{1,2} G. Bagdasarov,^{3,4} N. Bobrova,^{3,5} V. Gasilov,^{3,4} G. Korn,⁶ P. Sasorov,^{3,6} and W. P. Leemans^{1,2,†}

¹Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA
²University of California, Berkeley, California 94720, USA
³Keldysh Institute of Applied Mathematics RAS, Moscow 125047, Russia
⁴National Research Nuclear University MEPhI (Moscow Engineering Physics Institute), Moscow 115409, Russia
⁵Faculty of Nuclear Science and Physical Engineering, CTU in Prague, Brehova 7, Prague 1, Czech Republic
⁶Institute of Physics ASCR, v.v.i. (FZU), ELI-Beamlines Project, 182 21 Prague, Czech Republic

Bella PW laser : $E_L = 31 \text{ J}$; $\Delta t = 35 \text{ fs}$
In 20 cm capillary $n_e = 3 \cdot 10^{17} / \text{cm}^3$

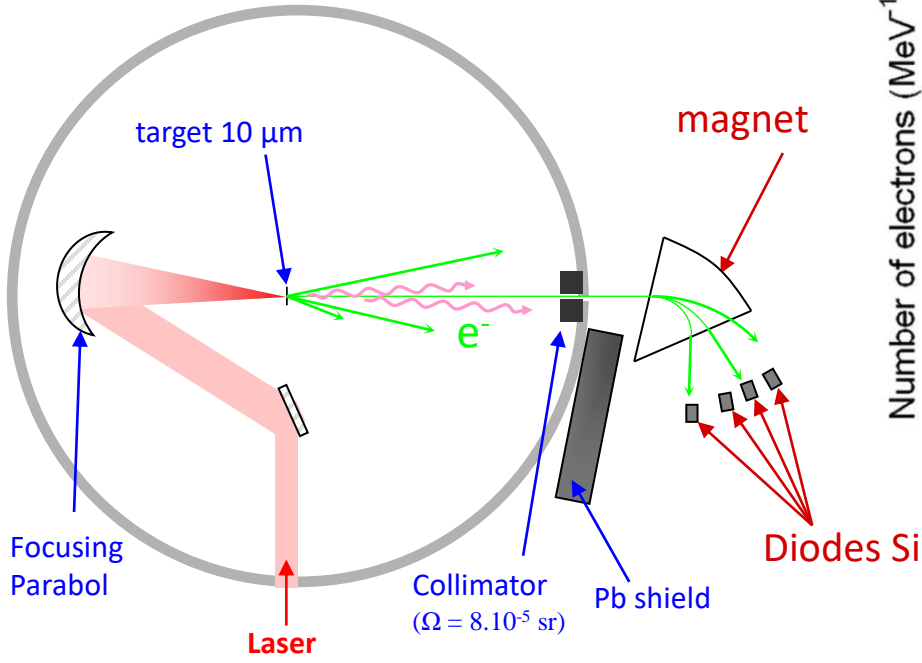


For nuclear reactions, we need tens of MeV electrons in huge quantity for bremsstrahlung

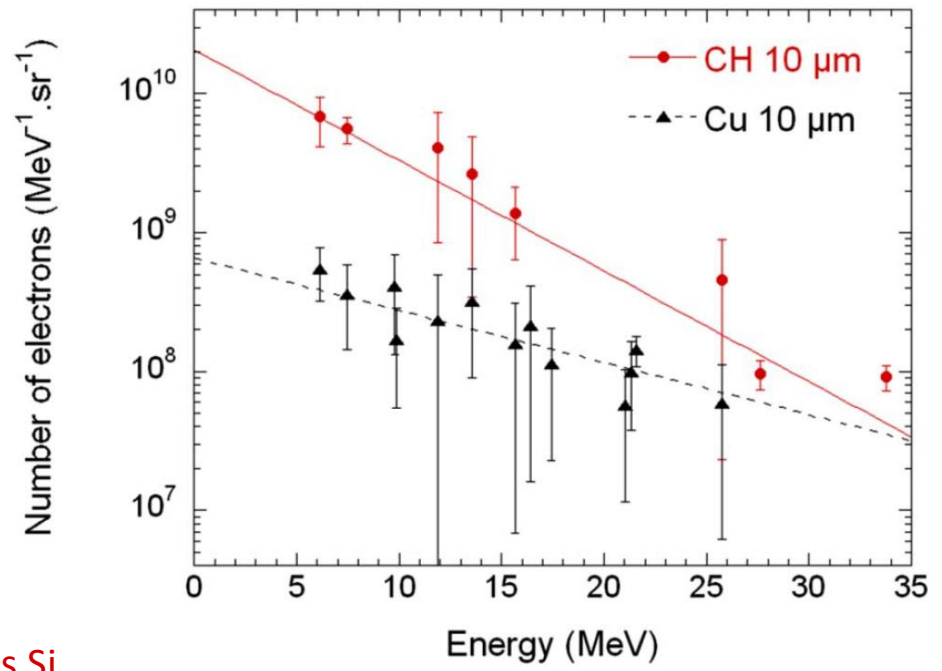
Electron source from solid targets



Laser	Energy (Joules)	pulse duration (fs)	power (TW)	focal (μm)	rate (Hz)	Intensity (W/cm^2)
LOA salle jaune	1	30	33	25	10	7,E+18



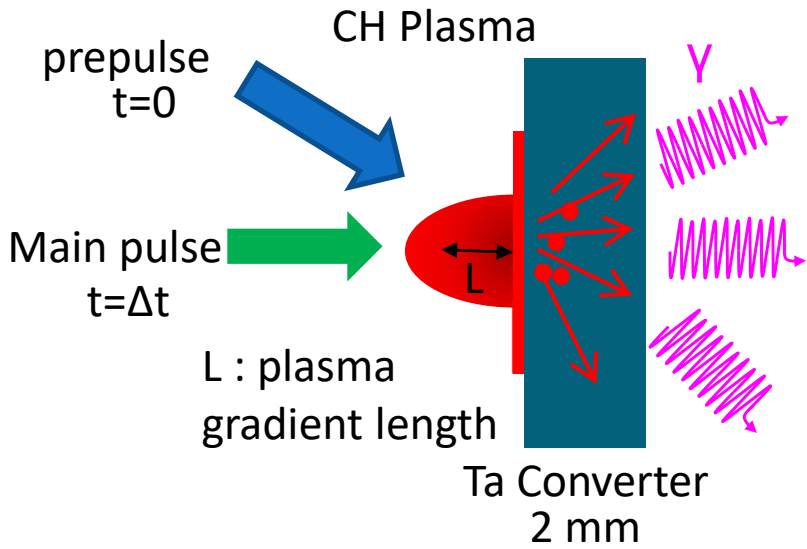
$\Omega = 8.10^{-5} \text{ sr}$



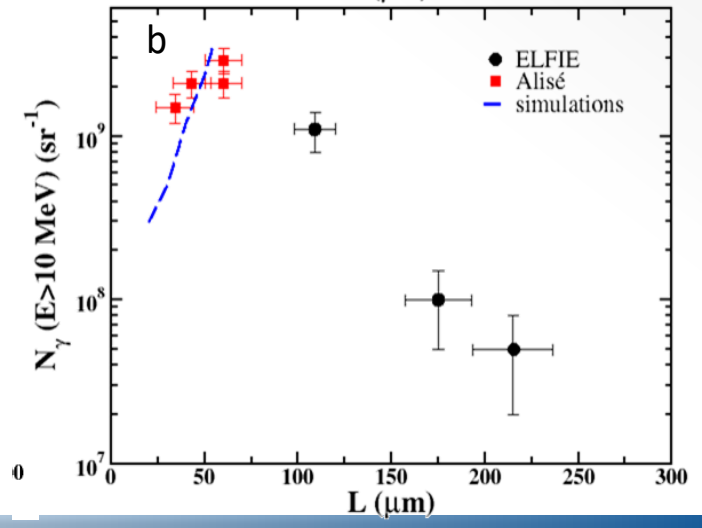
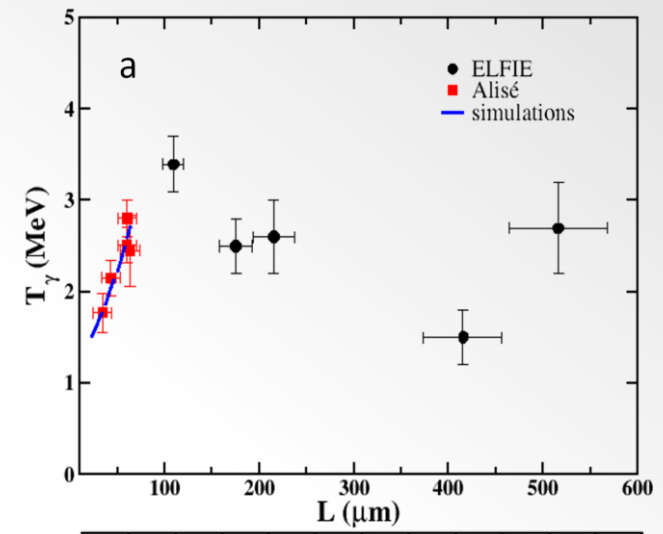
10 μm target	Temperature (MeV)
CH	5.9 ± 0.8
Cu	10 ± 2

Bremsstrahlung source

Laser	Energy (Joules)	Pulse duration(ps)	power (TW)	focal (μm)	rate (1/min)	Intensity (W/cm^2)
Pre-pulse	40	600	0,07	160	1/20	3,3E+14
Main pulse	7	0,4	17,5	17	1/20	8.E+18

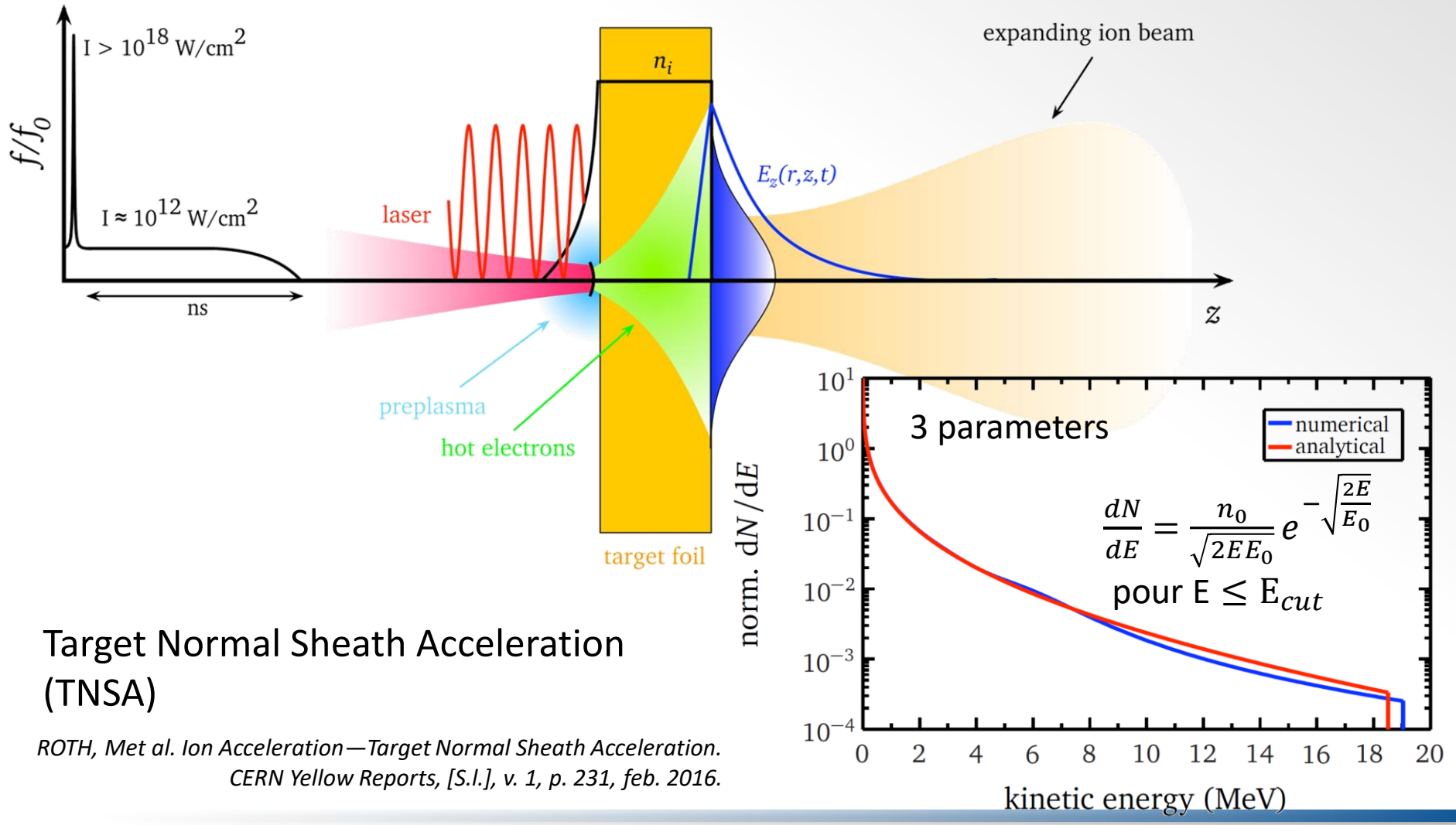


Photon source optimal (number and temperature)
 $L \approx 100 \mu\text{m}$



Laser driven ion source

with solid targets



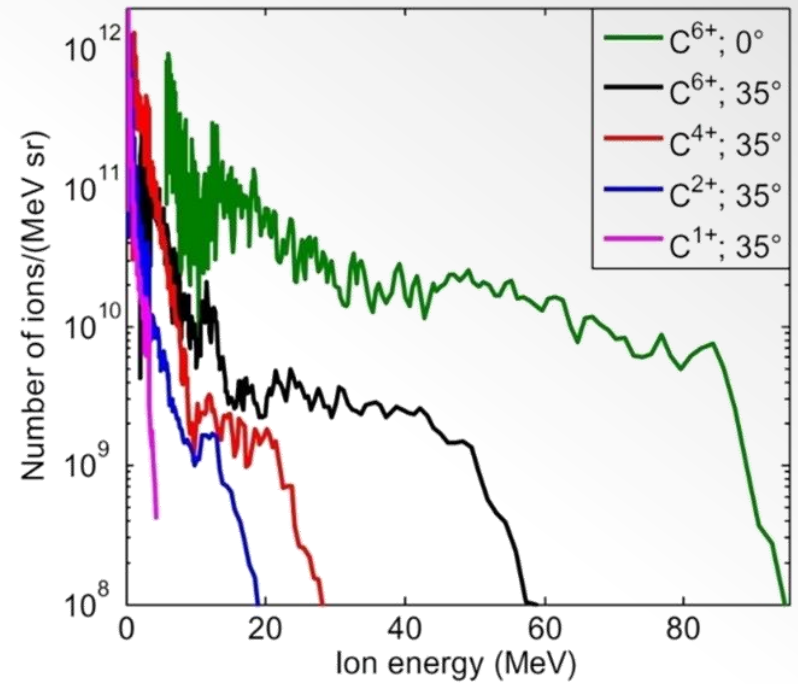
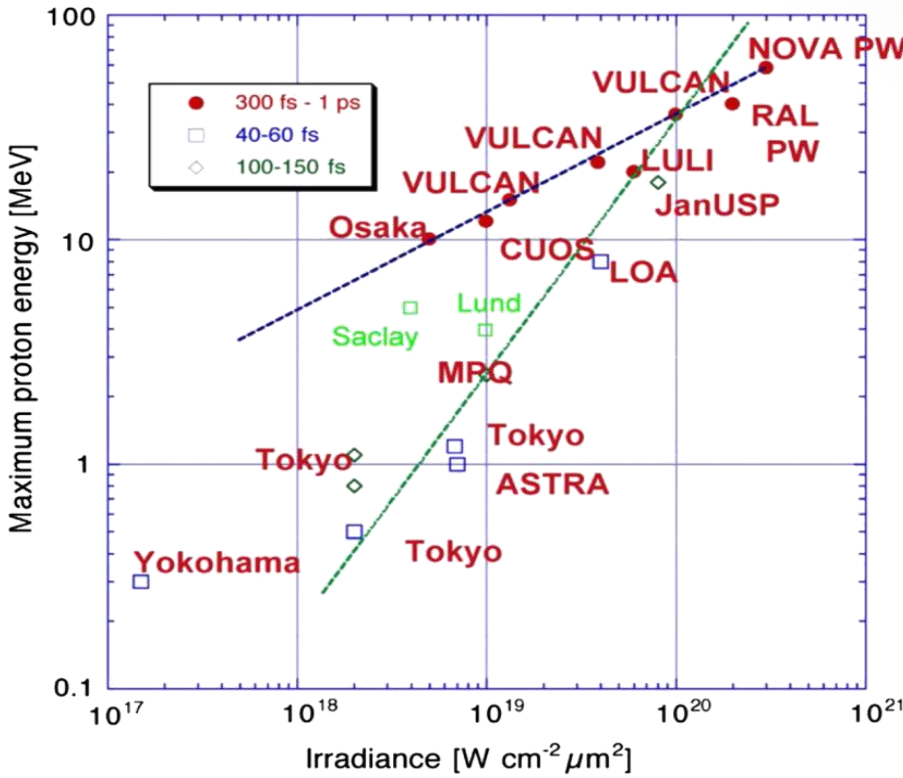
Target Normal Sheath Acceleration (TNSA)

ROTH, Met al. Ion Acceleration—Target Normal Sheath Acceleration. CERN Yellow Reports, [S.I.], v. 1, p. 231, feb. 2016.

Laser driven ion source

with solid targets

Laser accelerated protons overview



Carroll, D. C et al., *New Journal of Physics* 12 (2010) 045020 (15pp)

Astra-Gemini Laser : 115 TW ; 6 J ; 50 fs
=> $7 \times 10^{20} \text{ W cm}^{-2}$

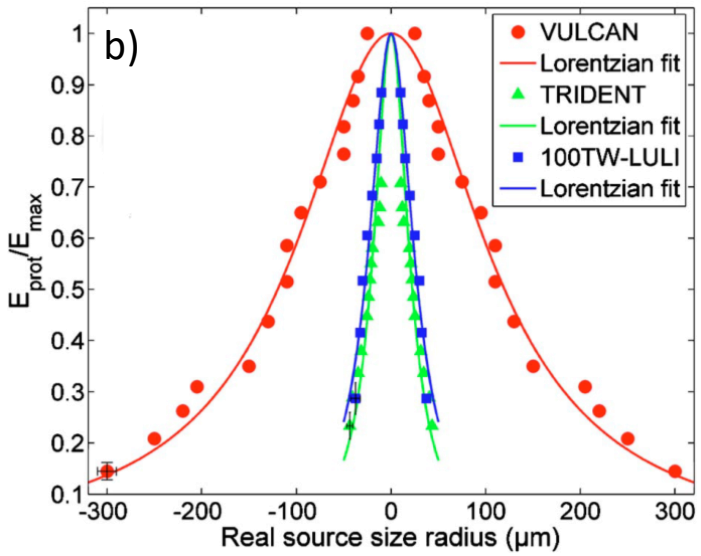
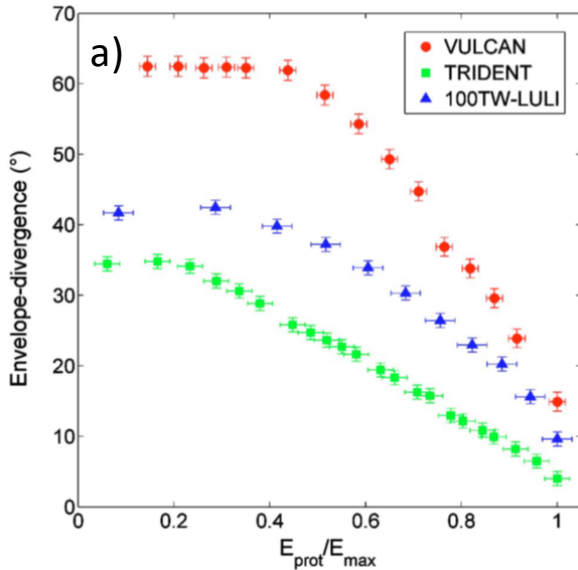
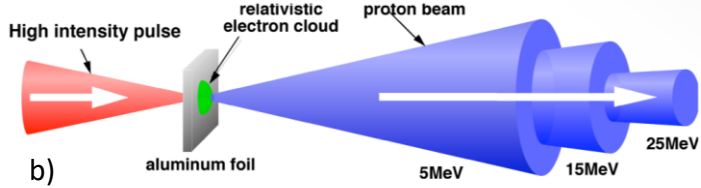
Macchi et al. *Rev. Mod. Phys.*, 85, 751 (2013)

TNSA

Laser driven ion source

with solid targets

Laser	Energy (J)	duration (fs)	Focal spot (μm)	Gold thickness (μm)	Target	E_{max} (MeV) (cut-off)	ϕ (μm) source diameter @ E_{max}
Vulcan	125	1000	5	25		17.4	80
Trident	19	600	14	10		13.5	20
100 TW-Luli	15	8	50		13.3	20	



Good emittance for injectors in conventional accelerators

F. Nürnberg et al., Rev.Sci.Instrum, 80, 033301 (2009)

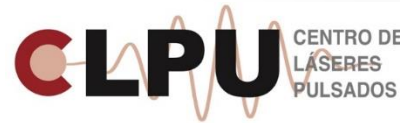
TNSA well known, reliable, suitable energies BUT not compatible with new generation of lasers

Laser driven ion source with gas jet targets

Because of high-repetition rate lasers (0,01-10 Hz)



10 PW every minute



1 PW every second

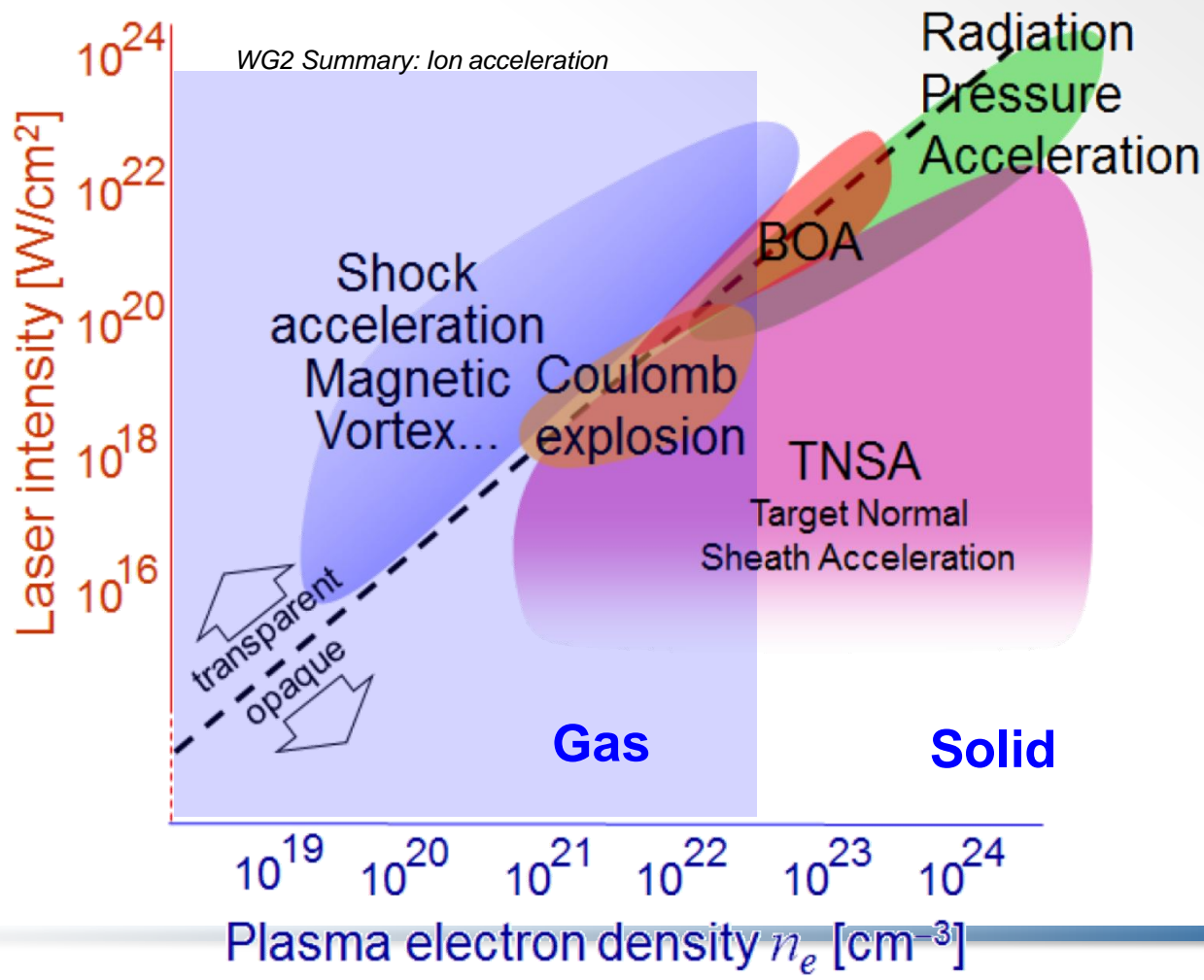


1 PW every 0.1 second

- ✓ Target regeneration and alignment
- ✓ Less debris production (25 weeks of pico 2000 \approx 2 min ELI-BL HAPLS@10Hz)
- ✓ Repeatability shot to shot
- ✓ Easy access to different ions
- ✓ Promote acceleration processes for high energies, high flux of ions and no maxwell distribution of energy

Laser driven ion source with gas jet targets

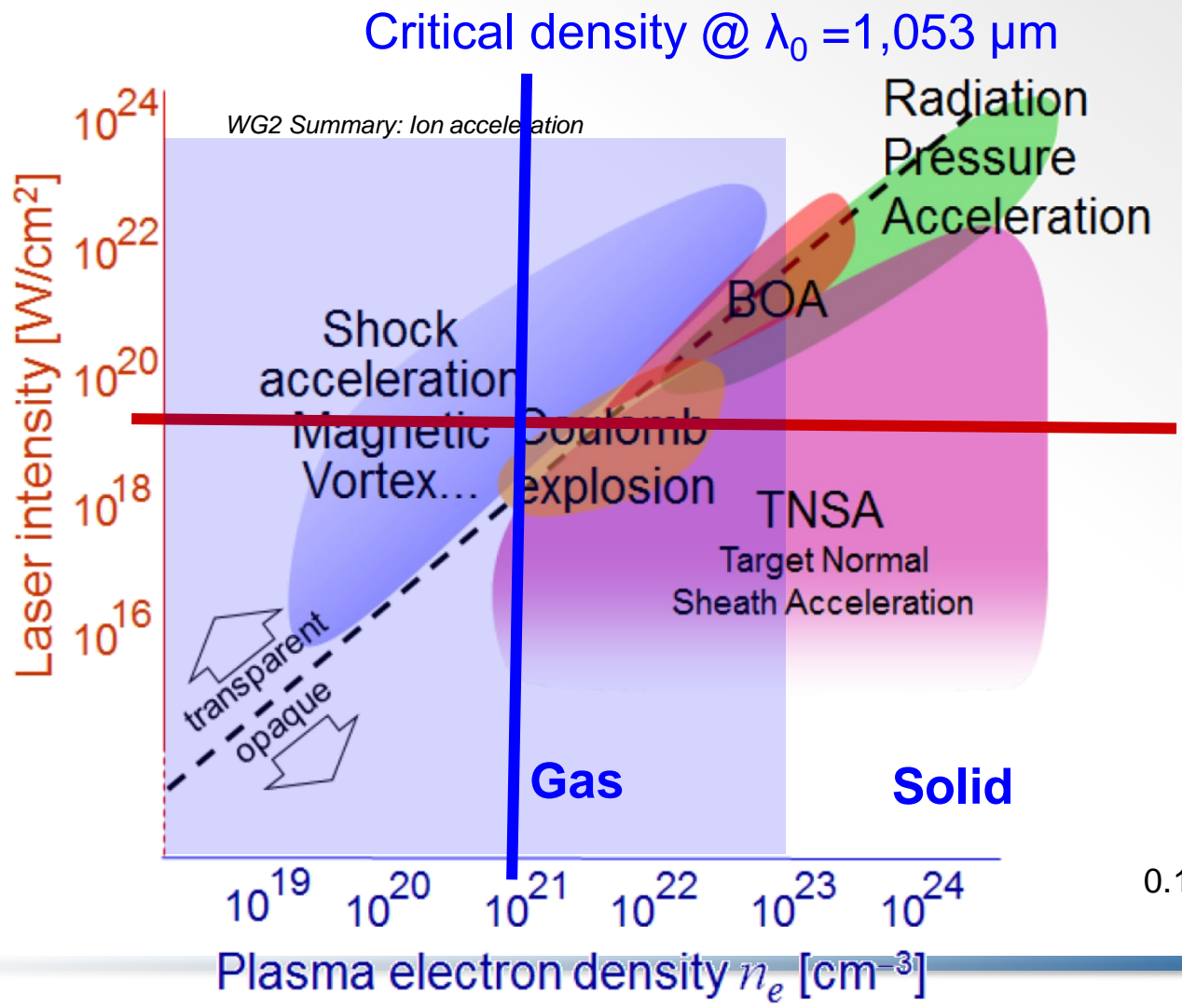
Ion acceleration processes



Laser driven ion source with gas jet targets

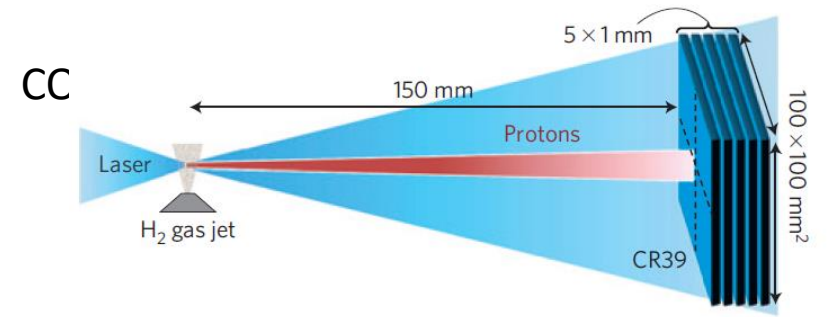
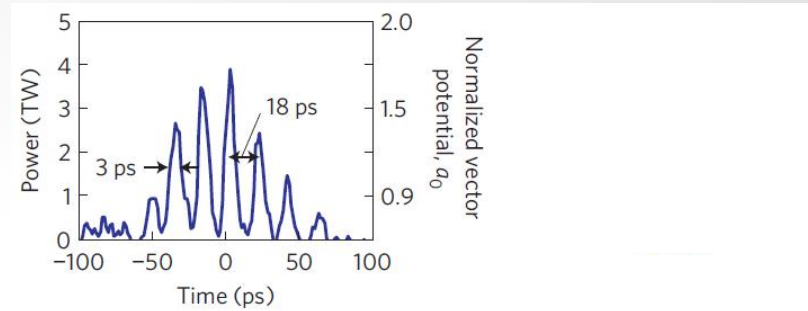
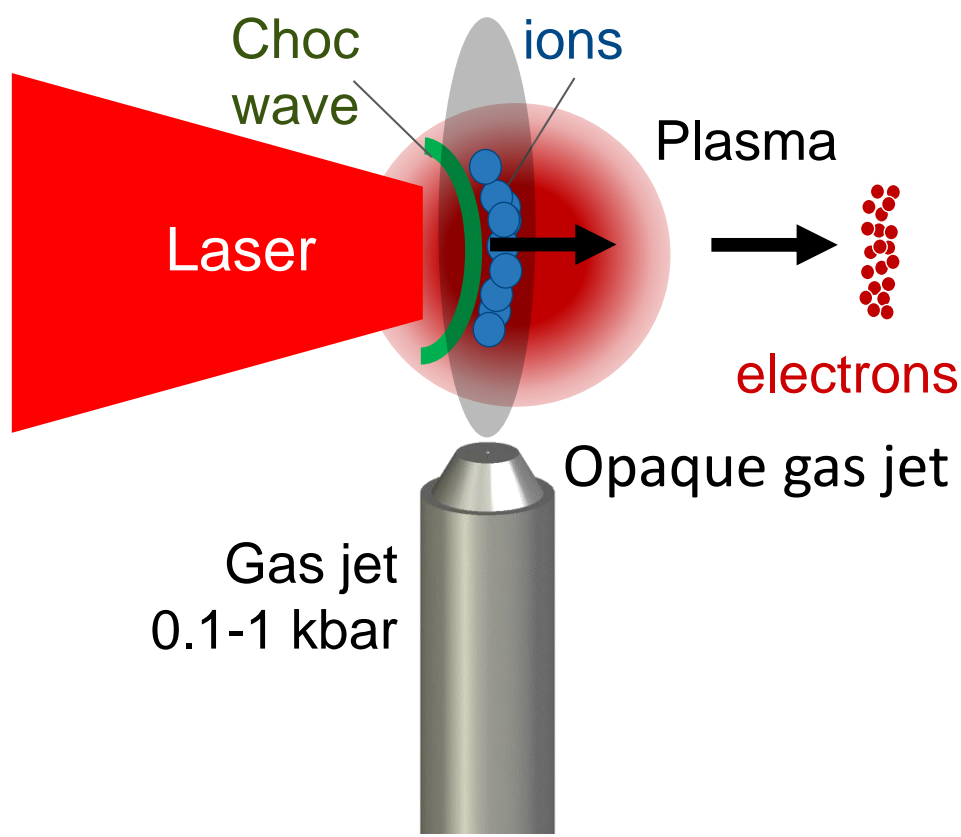


Laser
PICO2000 :
50J ; 1ps ;
 $5 \cdot 10^{19} \text{ W/cm}^2$
 $\lambda_0 = 1,053 \mu\text{m}$

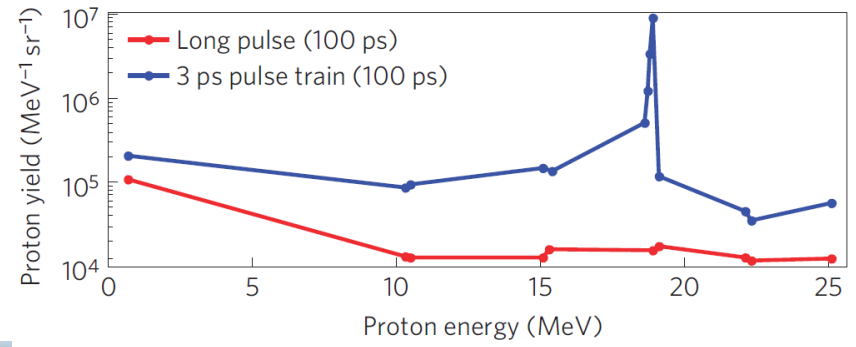


Laser driven ion source with gas jet targets

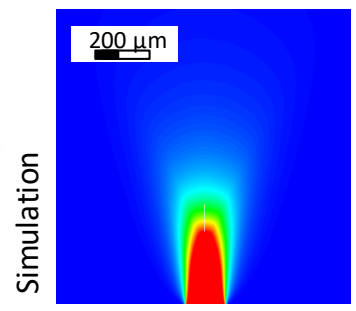
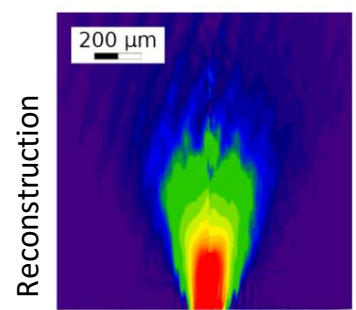
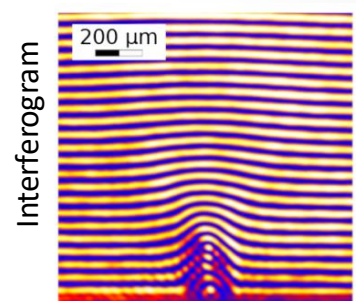
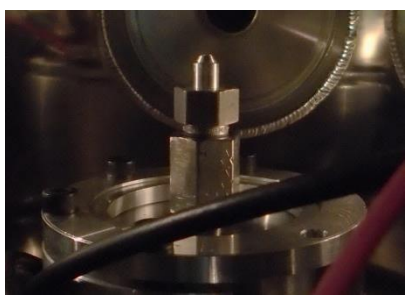
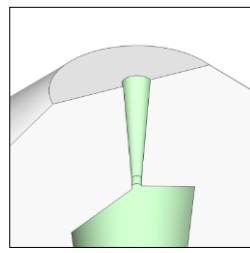
Collisionless Shock Acceleration (CSA)



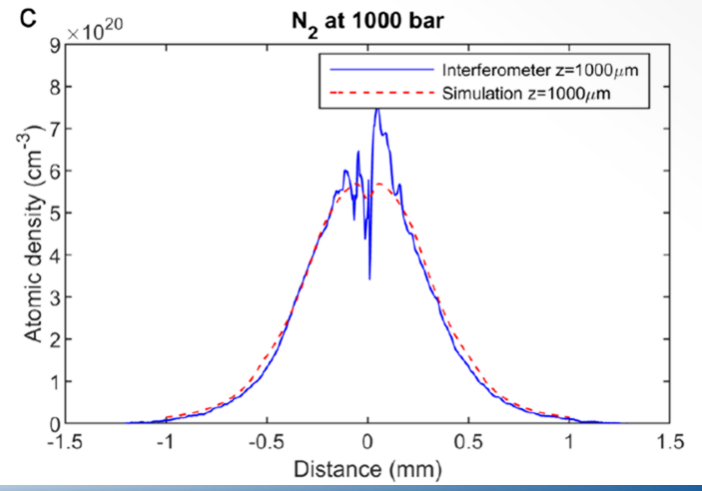
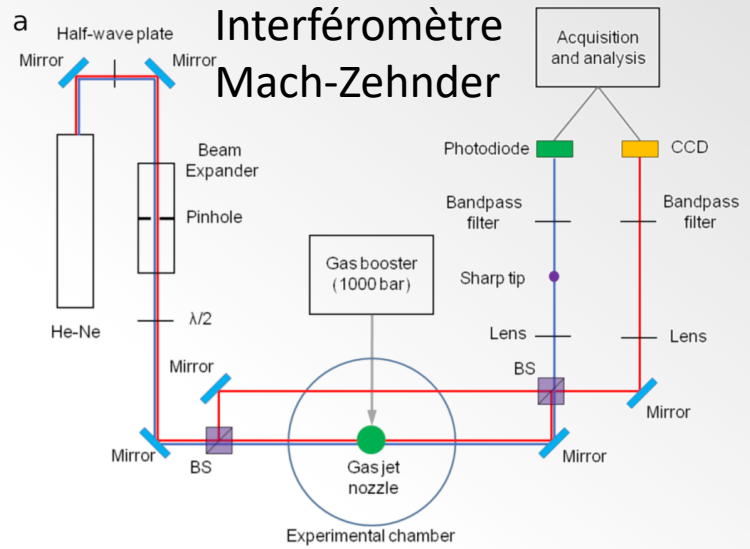
Van der Geer et al., Nature Physics, 2016, 12, 95-99 (2016)



Laser driven ion source with gas jet targets



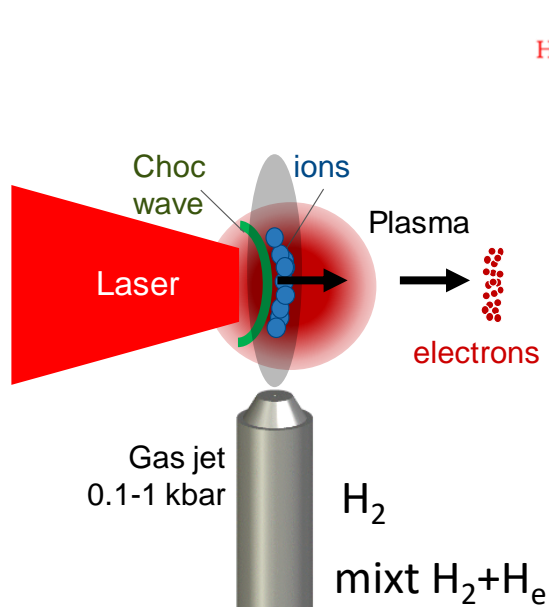
ANSYS FLUENT



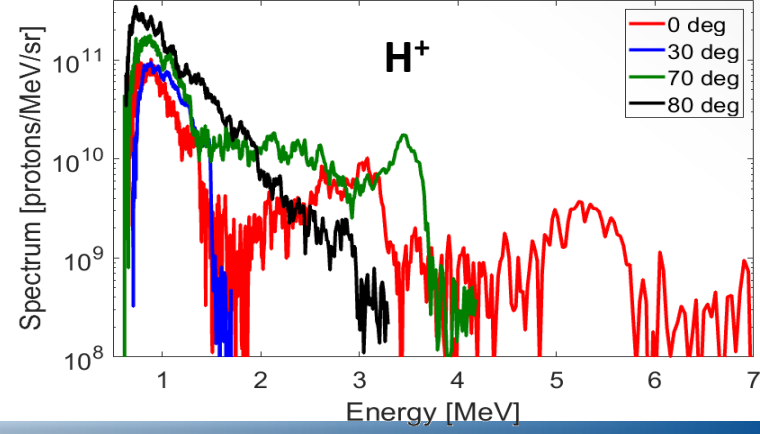
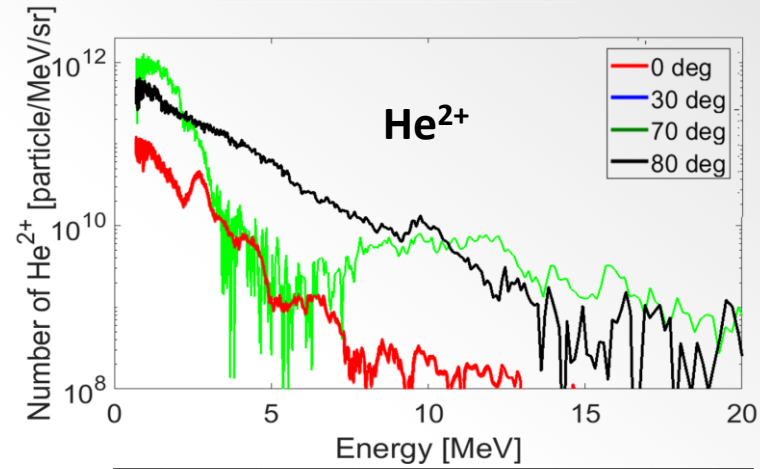
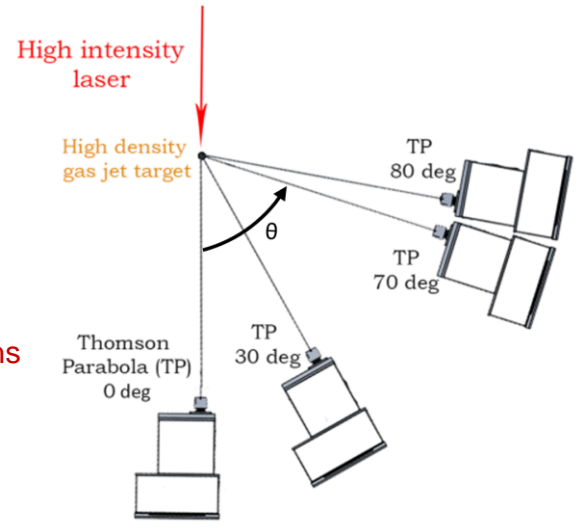
Laser driven ion source with gas jet targets



Laser	Energie (Joules)	Durée du pulse (ps)	Puissance (TW)	focal (μm)	Intensité (W/cm^2)
PICO 2000	60	1	60	13	5E+19

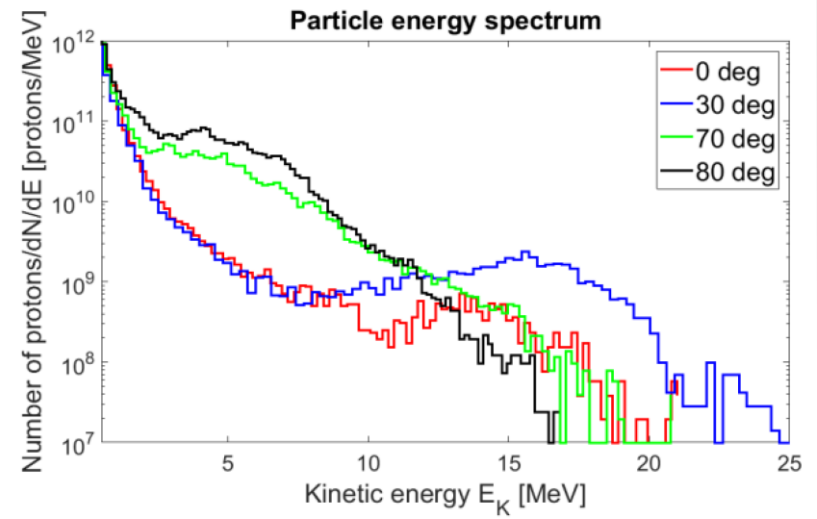
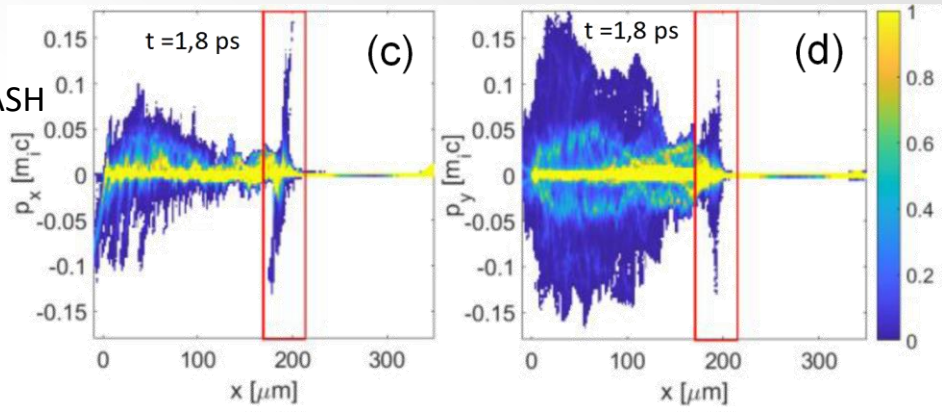
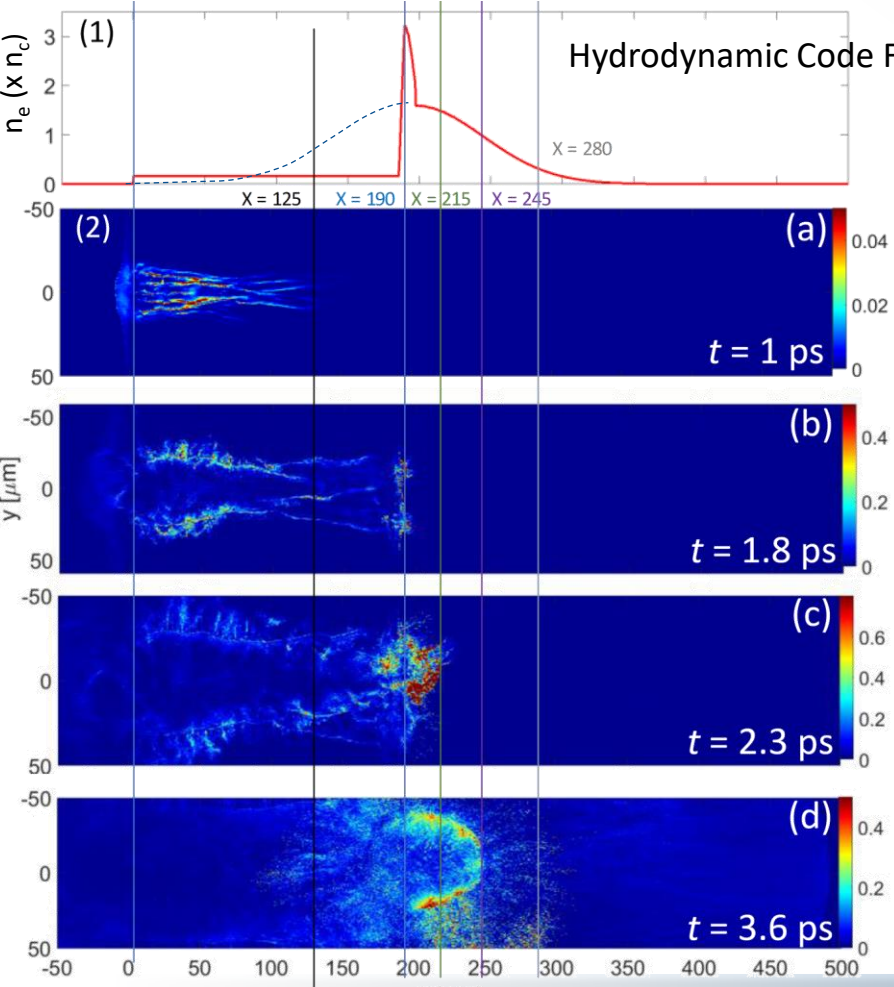


CSA :
Collisionless Shock Acceleration



Laser driven ion source with gas jet targets

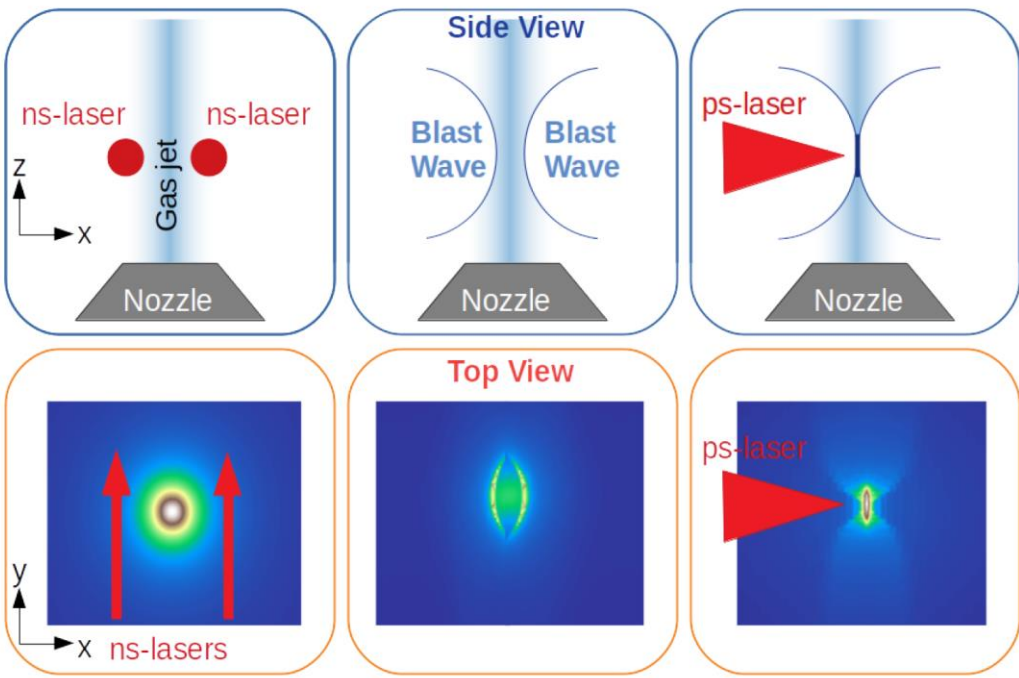
Code Particle In Cell : PICLS



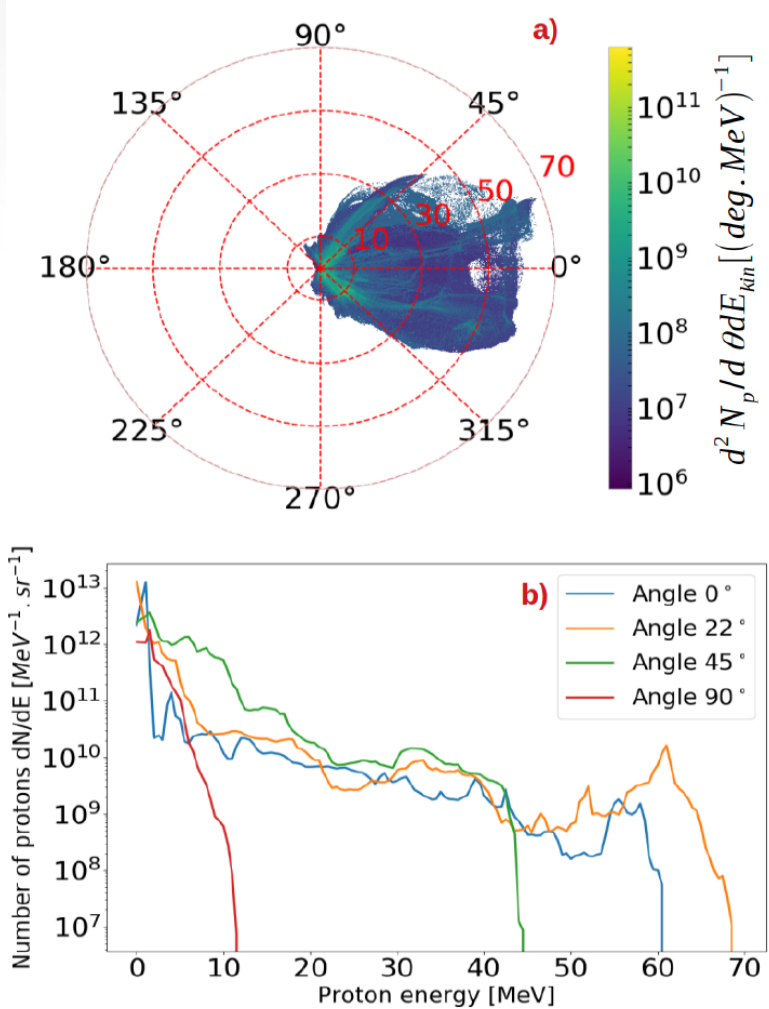
Hole Boring involved but no CSA

Laser driven ion source with gas jet targets

To promote the CSA process ... the **Plasma Tailoring**



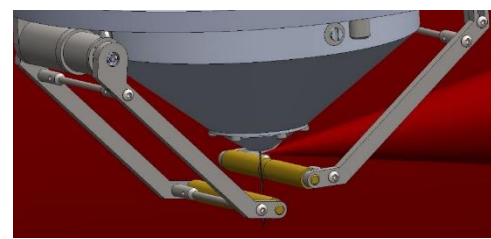
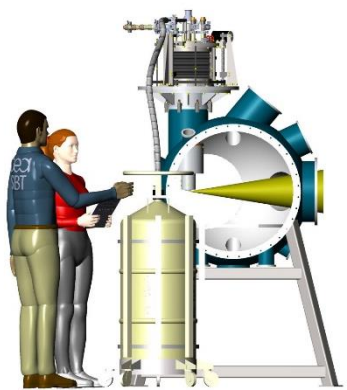
- Experiment at GSI in November 2020 :
structures of energy distribution confirmed, but
energy around 1 – 2 MeV



Laser driven ion source with other targets

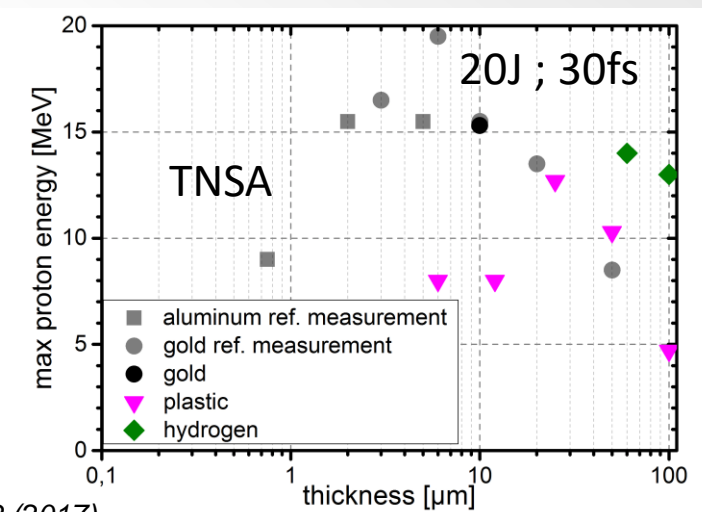
- TNSA on cryogenic target

Ribbon 1mm x 100μm



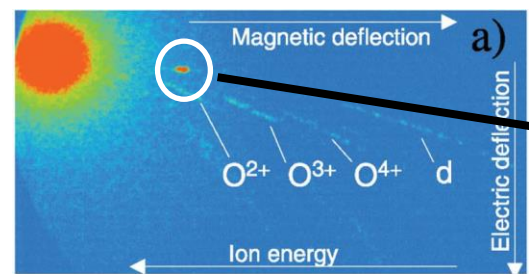
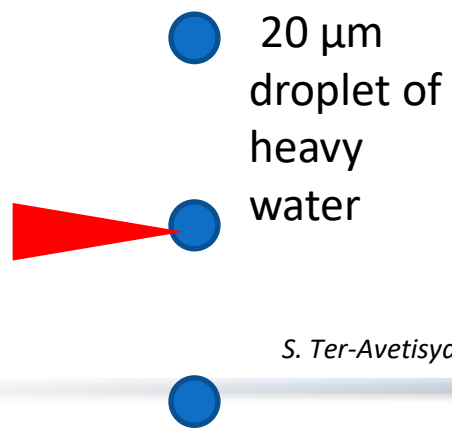
Courtesy F.Souris; D-SBT - CEA/Université Grenoble Alpes

“First demonstration of multi-MeV proton acceleration from a cryogenic hydrogen ribbon target » *Plasma Phys. Control. Fusion* 60 (2018) 044010 (6pp)

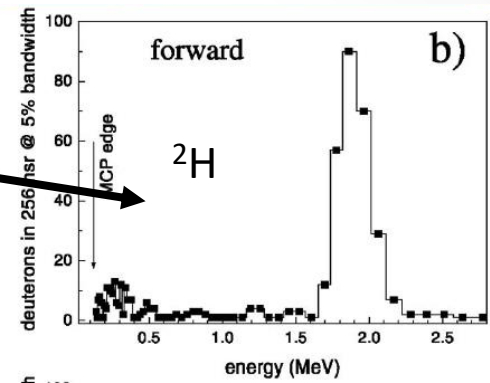


- Liquid targets *M. Gauthier, et al. Appl. Phys. Lett.* 111, 114102 (2017)

750 mJ ; 40 ;
 10^{19} W/cm² ;
 10 Hz



S. Ter-Avetisyan et al, Phys. Rev. Lett. 96(14) 145006 (2006)



Indirect Laser driven neutron source

Secondary source, the same as conventional accelerators

- **Laser accelerated ions (protons, deuterons) impinging a converter (Li; ^2H ; ^3H ...)**
- Bramsstrahlung radiation impinging a converter
- 2 GeV impinging a spallation target (not yet relevant)

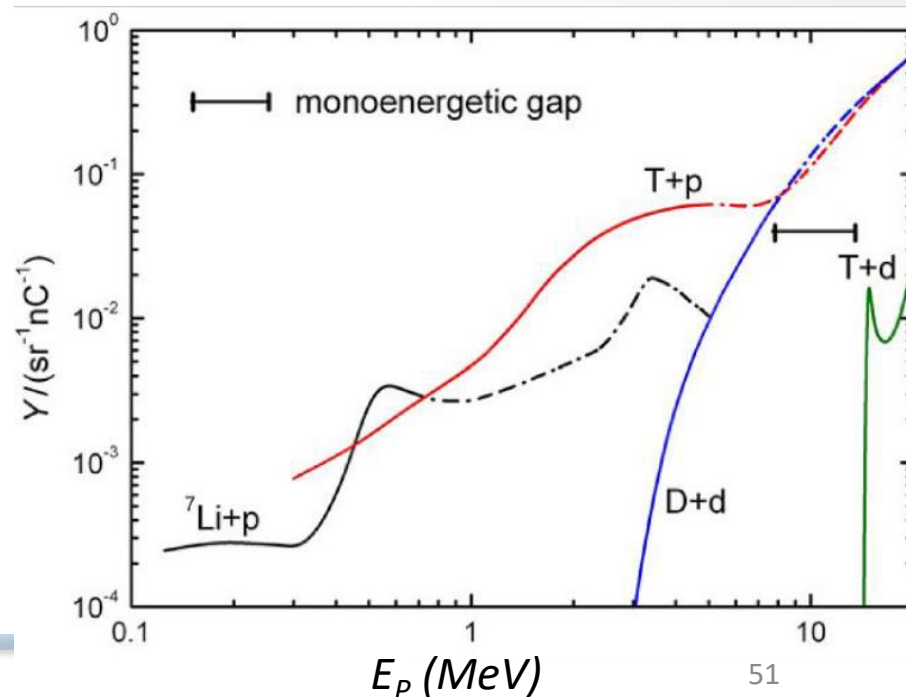
A.R. Junghans, EJC “Neutrons and Nuclei », <https://ejc2014.sciencesconf.org> (2014)

Reaction	Neutron Energy (MeV)
$p+^7\text{Li} \rightarrow n+^7\text{Be}$	[0.121 - 0.649]
$p+T \rightarrow n+^3\text{He}$	[0.6 - 2.6]
$d+D \rightarrow n+^3\text{He}$	$2.45 + f(E_{\text{projectile}})$
$d+T \rightarrow n+^4\text{He}$	$\sim 14.1 + g(E_{\text{projectile}})$

Neutron energy depends on the emission angle

Require :

- a well define projectile beam axis
- A mono energetic projectile



Indirect Laser driven neutron source

Secondary source, the same as conventional accelerators

- **Laser accelerated ions (protons, deuterons) impinging a converter (Li; ^2H ; ^3H ...)**
- Bremsstrahlung radiation impinging a converter
- 2 GeV impinging a spallation target (not yet relevant)

	Laser	Acceleration process	Neutron production reaction	Number of neutrons/shot	Neutron energies
A	Vulcan : 200J, 3×10^{20} W/cm ²	TNSA : CD 10 μm	CD 2 mm : D(p,n+p) ¹ H ; D(d,n) ³ He	10^9 /sr	[0-25] MeV
B	Trident : 60 J, 600 fs ; 10^{21} W/cm ²	BOA : CD 400 nm	Be \sim mm : Deuteron break-up, $^9\text{Be}(p,n)^9\text{B}$, $^9\text{Be}(d,n)^{10}\text{B}$	1.2×10^{10} /sr	[10-150] MeV
C	Elfie : 10 J, 350 fs ; 1.1×10^{19} W/cm ²	TNSA : CH 50 μm on 14 nm Al	LiF 2mm : $^7\text{Li}(p,n)^7\text{Be}$; $^6\text{Li}(p,n)^6\text{Be}$; $^{19}\text{F}(p,n)^{19}\text{Ne}$;	$\sim 1 \times 10^4$ /sr	[0.1-4] MeV

A) S. Kar et al, *New J. Phys.* 18 053002, (2016)

B) M. Roth et al, *Journal of Physics: Conference Series* 688, 012094, (2016)

C) D.P. Higginson et al, *PRL* 115(5), [054802], (2015)

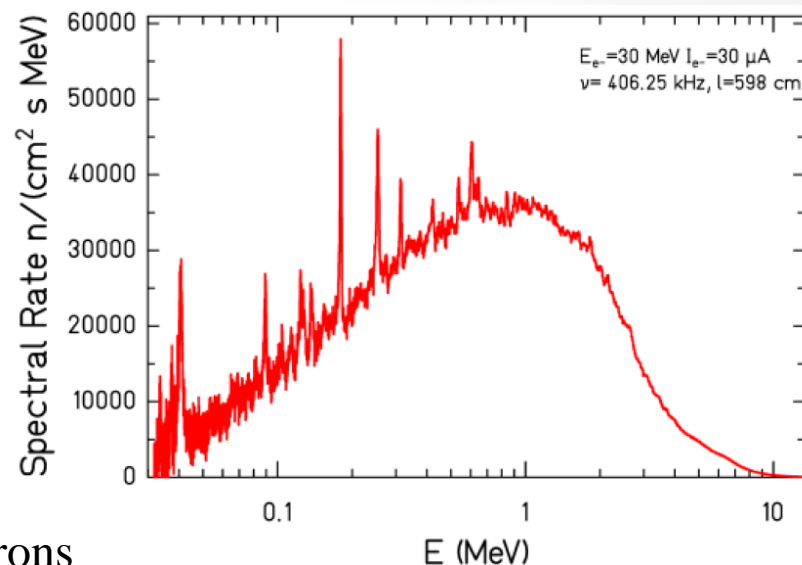
Indirect Laser driven neutron source

Secondary source, the same as conventional accelerators

- Laser accelerated ions (protons, deuterons) impinging a converter (Li; ^2H ; ^3H ...)
- **Bramsstrahlung radiation impinging a converter**
- 2 GeV impinging a spallation target (not yet relevant)

2 Facilities in Europe : ELBE at Dresden, Gemany and GELINA at Geel, Belgium.

	Target	Source Strength (s^{-1})	E_{e^-} (MeV)	I_{e^-} (μA)	f (Hz)
nELBE	Pb, liquid	$3 \cdot 10^{11}$	30	15	$2 \cdot 10^5$
GELINA	U, Hg cooled	$3 \cdot 10^{13}$	100	96	800



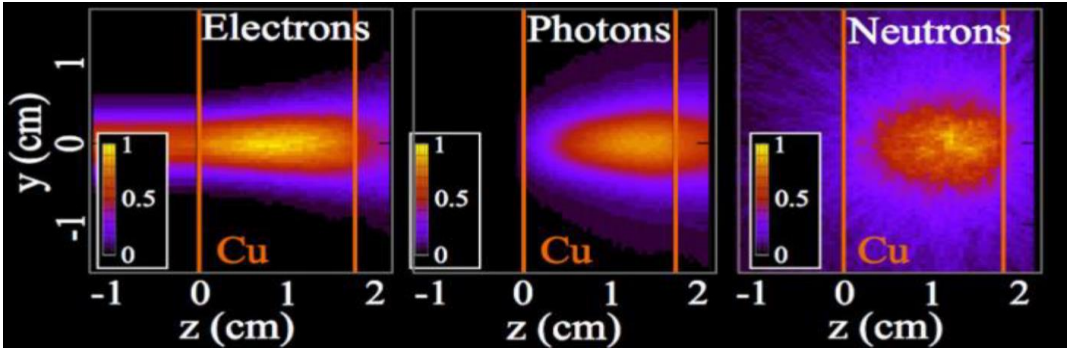
n-ELBE : 10^6 neutrons in 4π from 5×10^8 electrons

Indirect Laser driven neutron source

Secondary source, the same as conventional accelerators

- Laser accelerated ions (protons, deuterons) impinging a converter (Li; ^2H ; ^3H ...)
- **Bramsstrahlung radiation impinging a converter**
- 2 GeV impinging a spallation target (not yet relevant)

Laser 90J, 150fs



I. Pomerantz et al, PRL 113, 184801 (2014)

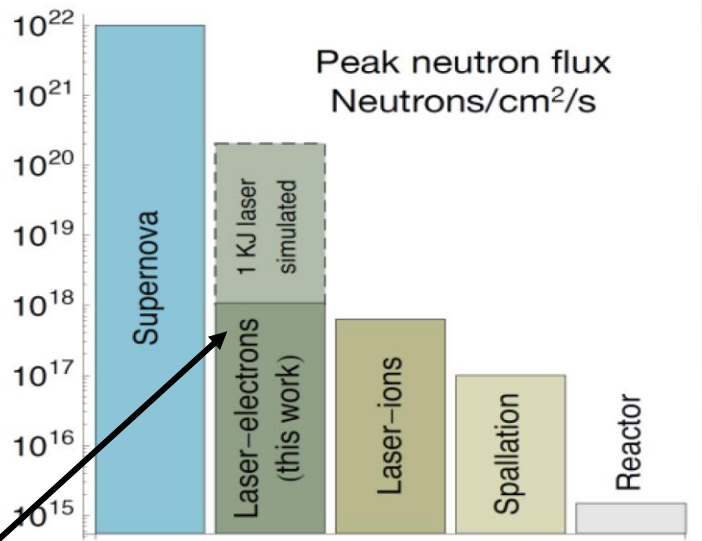
neutrons pulse duration : 50 ps

5×10^8 n/shot

$\sim 4 \times 10^7$ n/sr / shot

$\Leftrightarrow \sim 10^8$ n/cm² / shot @ 50 cm from source

\Leftrightarrow A peak of $\sim 10^{18}$ n/cm² / s



But not a continuous source but @10Hz :

$\sim 10^9$ n/cm² /s @ 50 cm

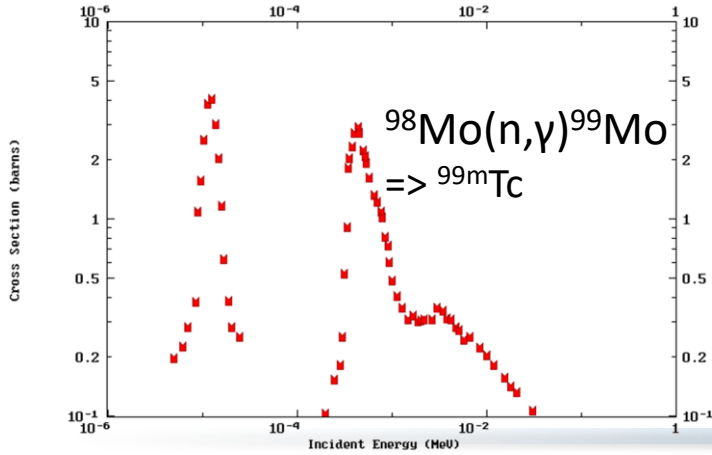
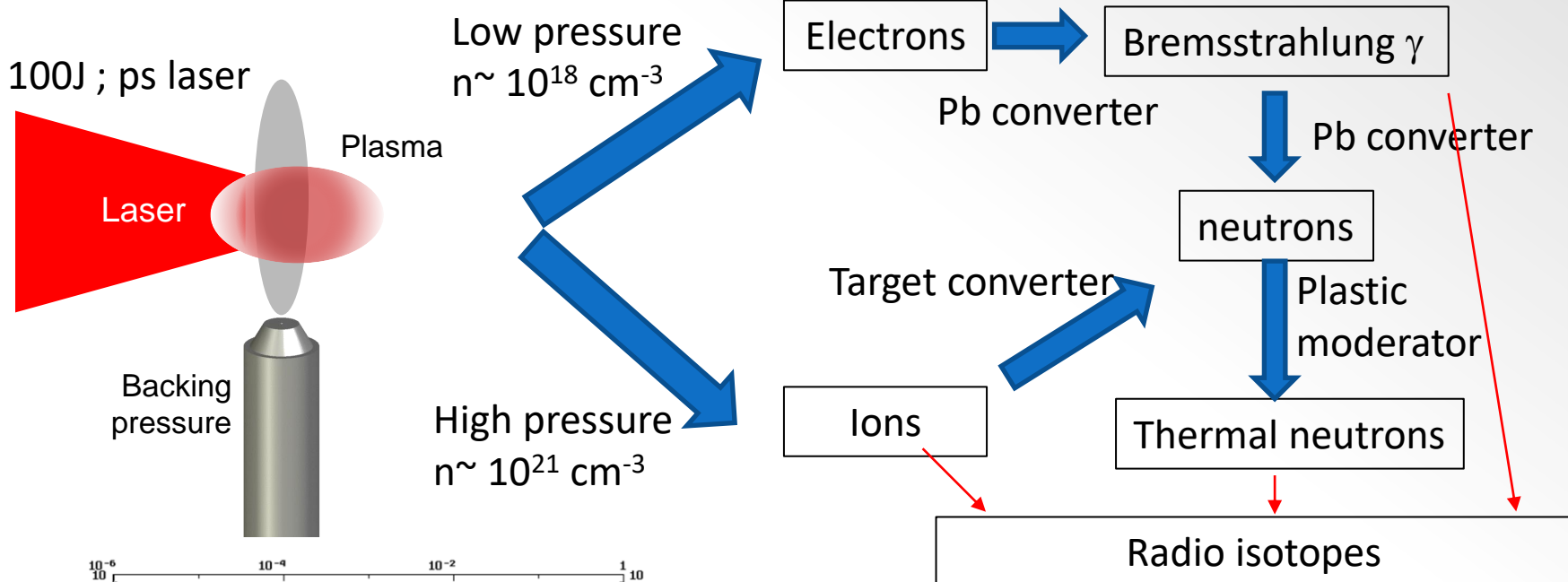
$\sim 10^{13}$ n/cm² /s @ 5 mm

The dreaming part...

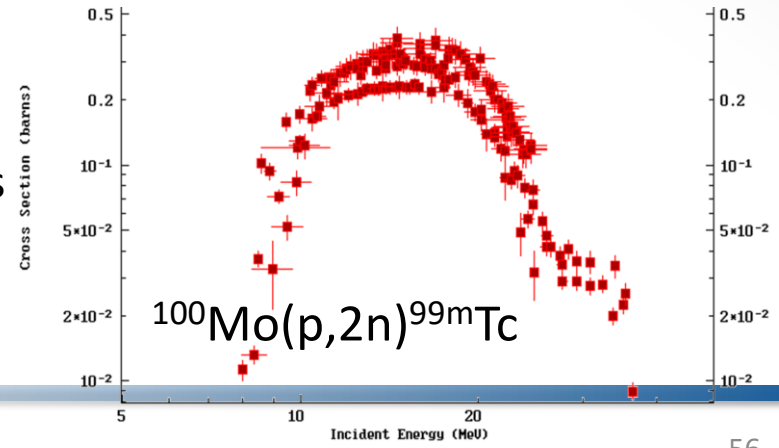
Open your mind,.....
breathe, ...
take time to dream...



A versatile source



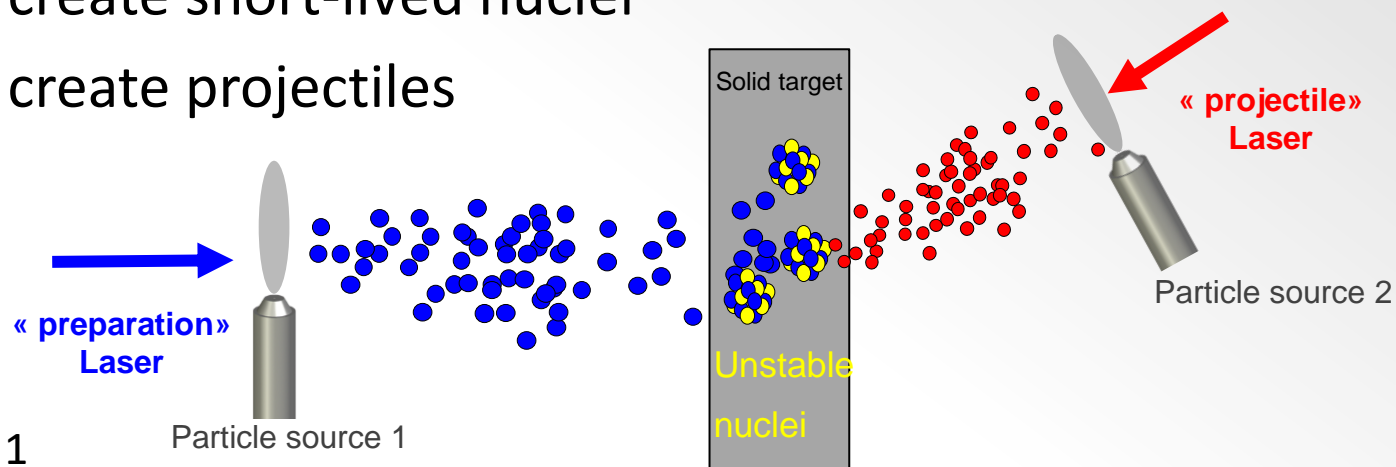
Medical radioisotopes



New type of nuclear physics experiments

- ✓ A beam to create short-lived nuclei
- ✓ A beam to create projectiles

Split the same laser



Particle from source 1

Projectiles : 10^{13} particles on \varnothing **100 μ m** spot ; cross section: **0,1 barn**,
Primary target: 10^{21} nuclei/cm² (ex : 100 μ m thick carbon)
→ 10^9 unstable nuclei on \varnothing 100 μ m ↔ 10^{13} nuclei / cm²

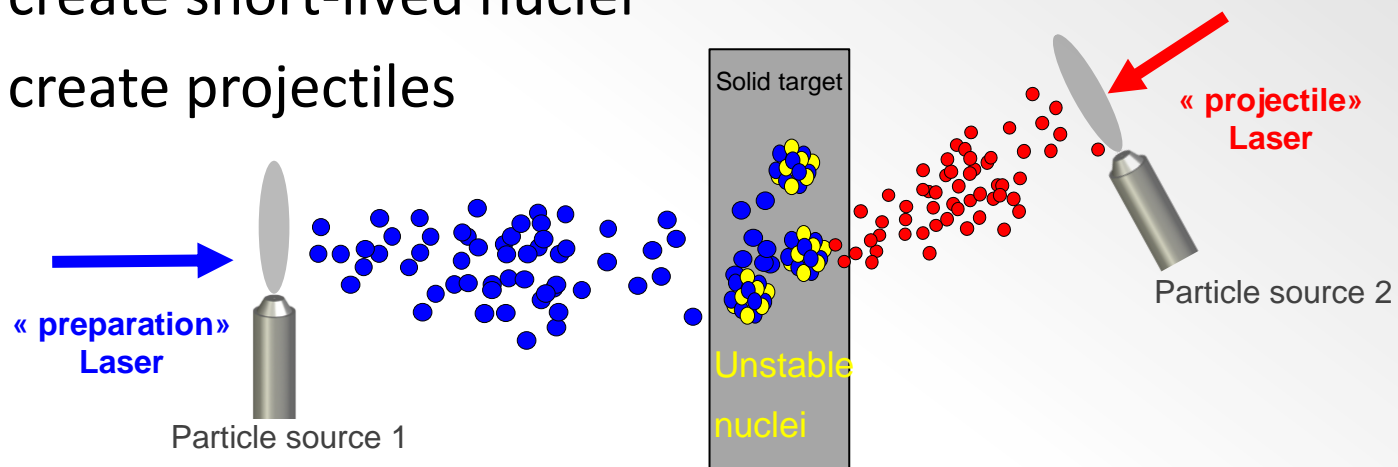
Wait Δt then

Particle from source 2

Projectiles : 10^{13} particles on \varnothing **100 μ m** spot ; cross section: **0,1 barn**,
secondary target: 10^{13} nuclei / cm²
→ 10 reactions/shot ; 1 shot / min → ~14 400 reactions/day

New type of nuclear physics experiments

- ✓ A beam to create short-lived nuclei
- ✓ A beam to create projectiles

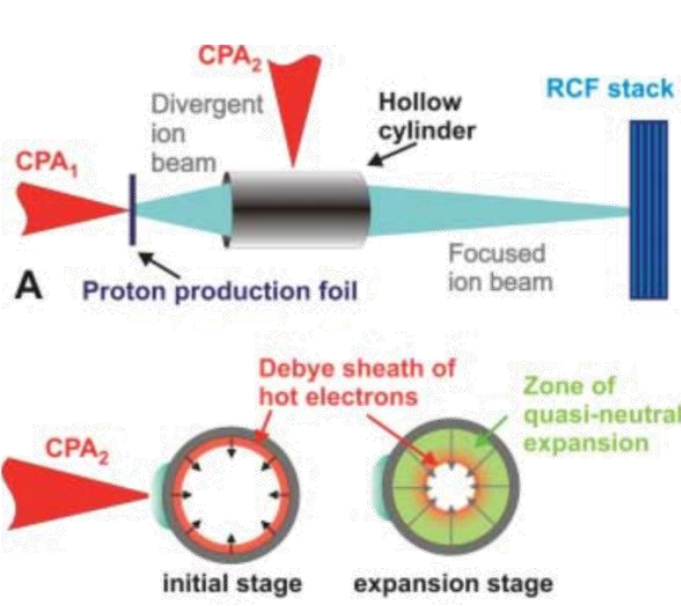


~10 000 reactions/day (0.1 barn)

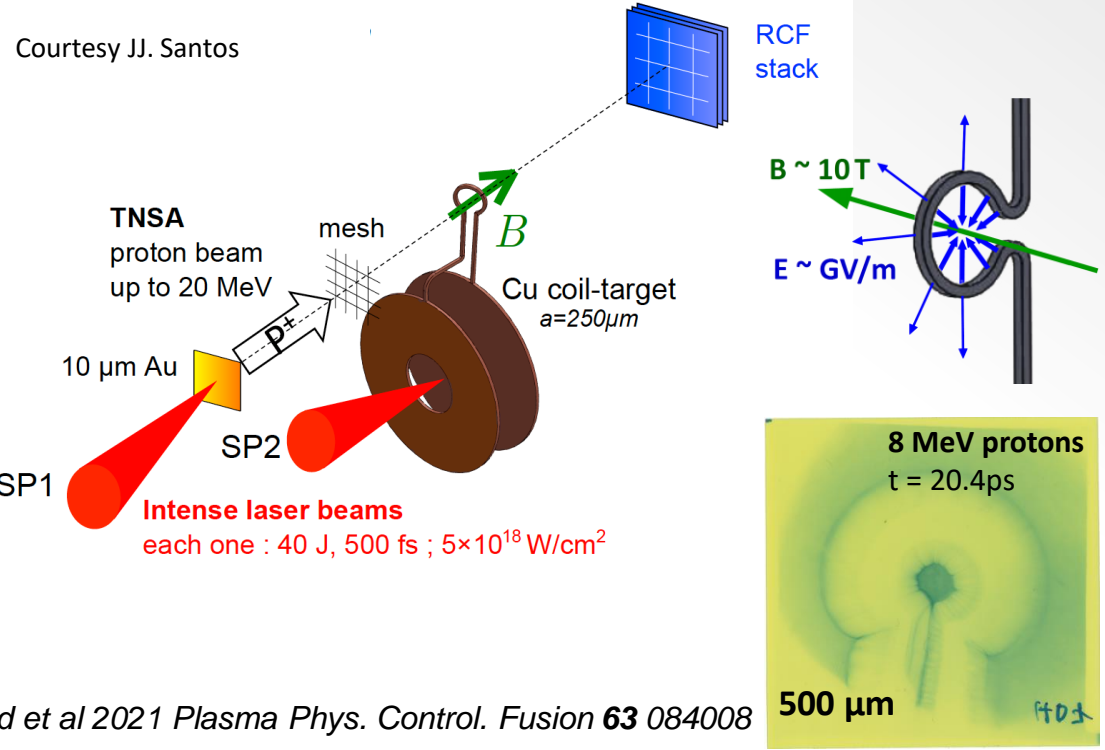
- ✓ Nuclear reactions on very short-lived radioactive nuclei (down to few ns)
- ✓ Nuclear reactions on excited nuclei
- ✓ BUT need to detect nuclear signal ☹️

New type of nuclear physics experiments

- ✓ Beam transport difficult because charge space :
 - ✓ 10^{13} protons @10 MeV in 1 mm diameter cylinder, flying in 100 ps bunch duration \leftrightarrow 1,6 kA in cylinder ; $5 \cdot 10^{-4}$ C/cm³



T. Toncian, et al. Science 312, 410-413 (2006)



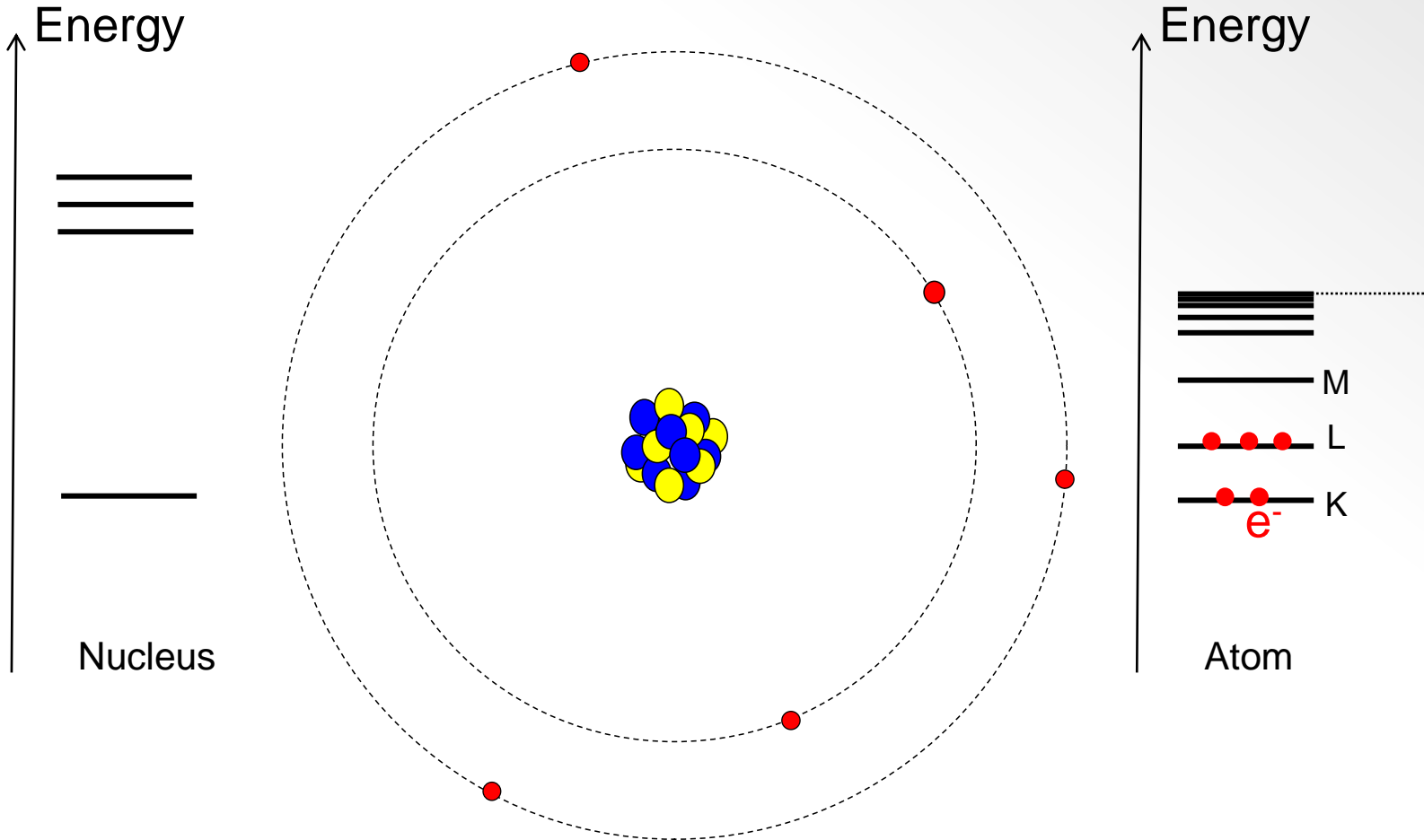
P Bradford et al 2021 Plasma Phys. Control. Fusion 63 084008

Part 3

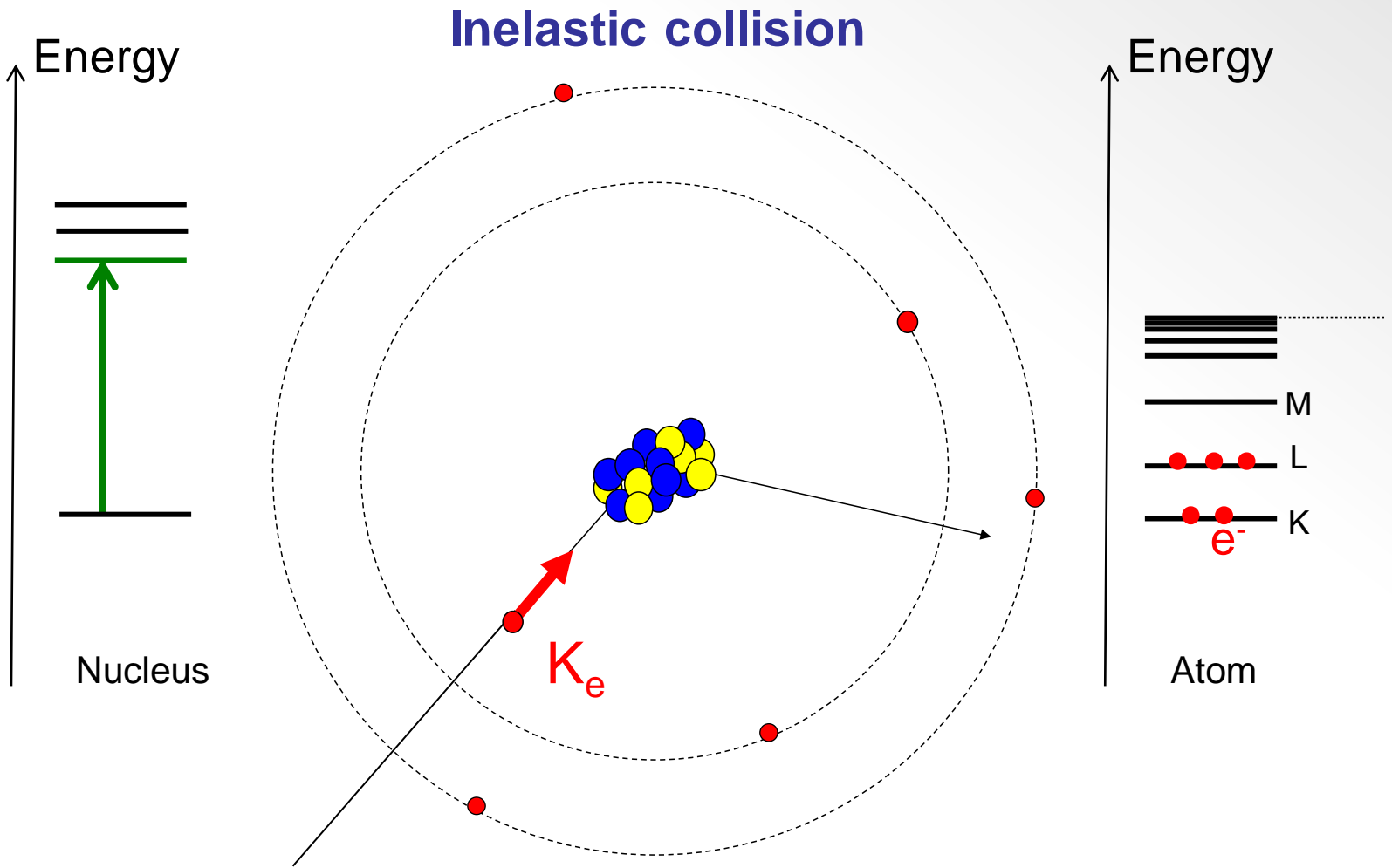
NUCLEAR PHYSICS IN PLASMAS

- The interplay between atomic electrons and nucleus
- Cross section modifications
- Half life modification : case of ^{84}Rb

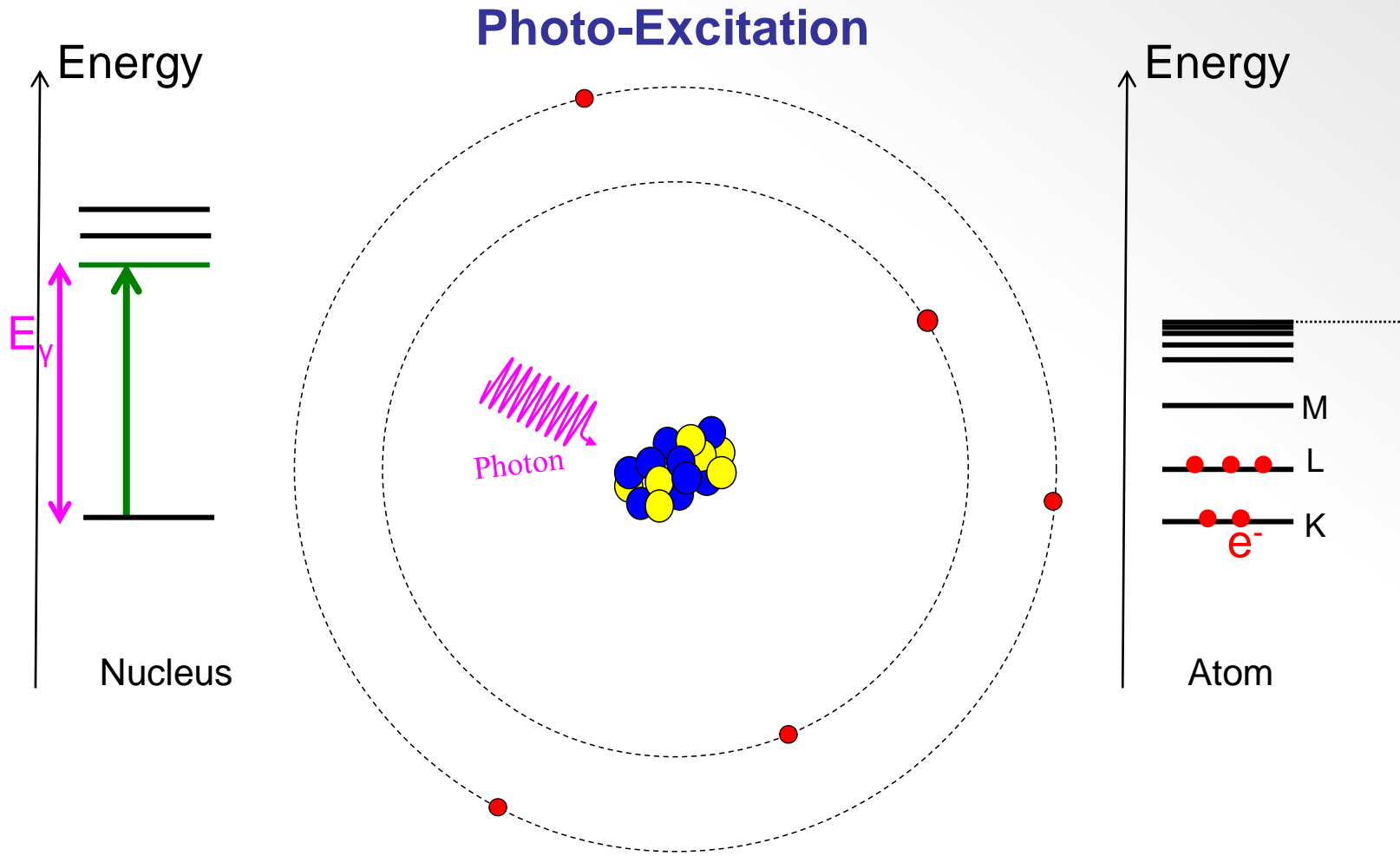
Nuclear excitation / de-excitation



Nuclear excitation / de-excitation

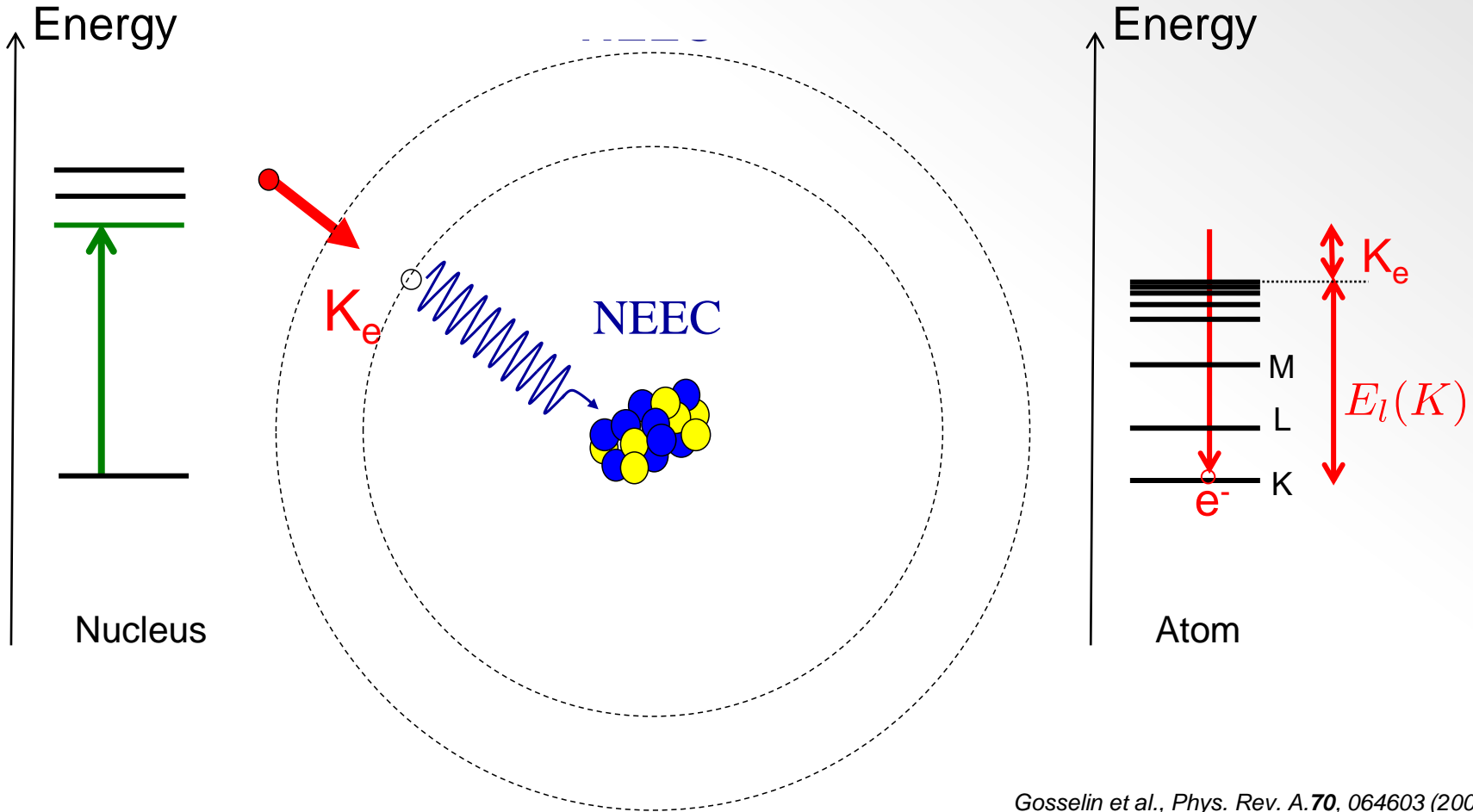


Nuclear excitation / de-excitation



Nuclear excitation / de-excitation

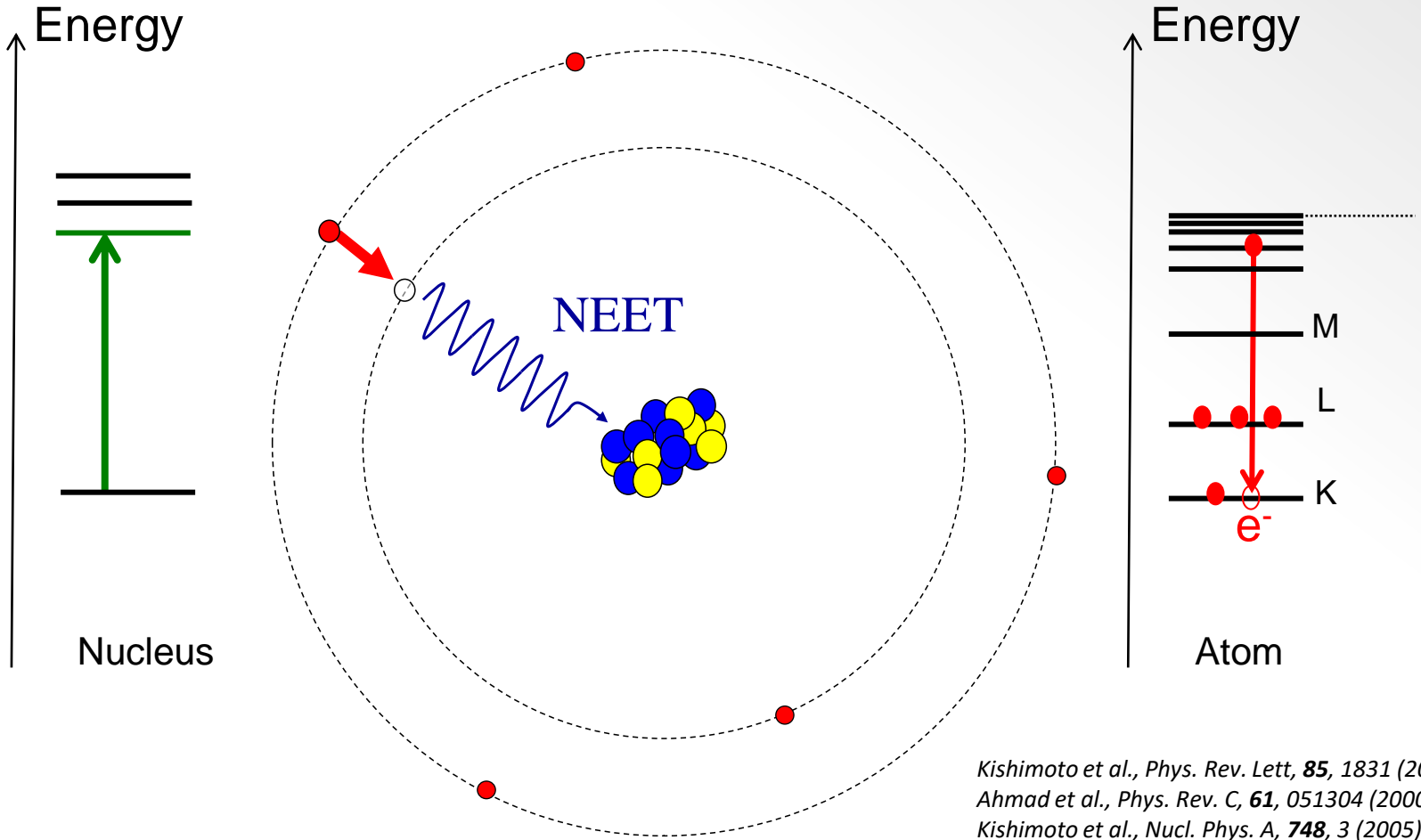
Nuclear Excitation by Electron Capture



Gosselin et al., Phys. Rev. A.70, 064603 (2004)

Nuclear excitation / de-excitation

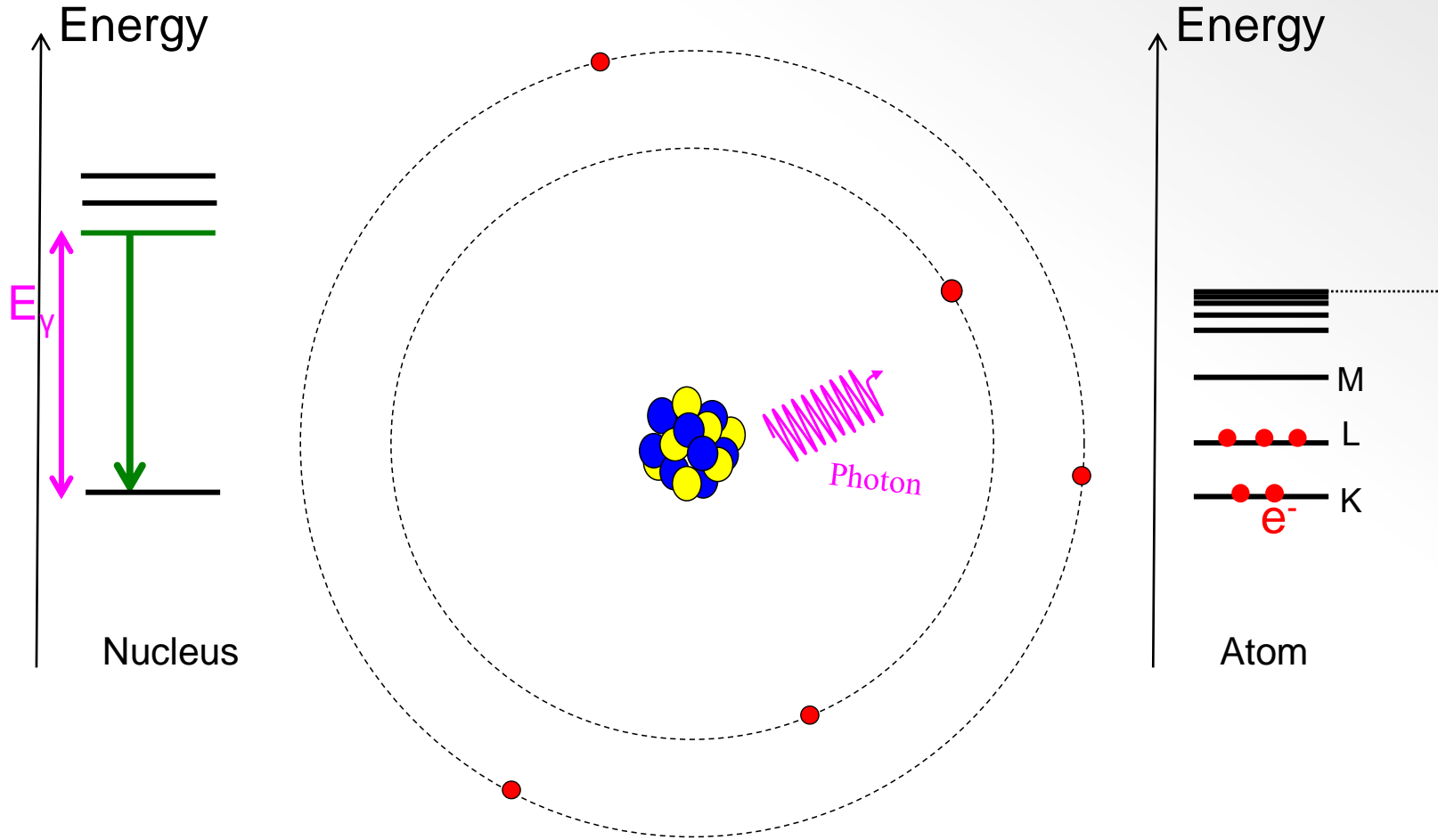
Nuclear Excitation by Electronic Transition



Kishimoto et al., *Phys. Rev. Lett*, **85**, 1831 (2000)
Ahmad et al., *Phys. Rev. C*, **61**, 051304 (2000)
Kishimoto et al., *Nucl. Phys. A*, **748**, 3 (2005)

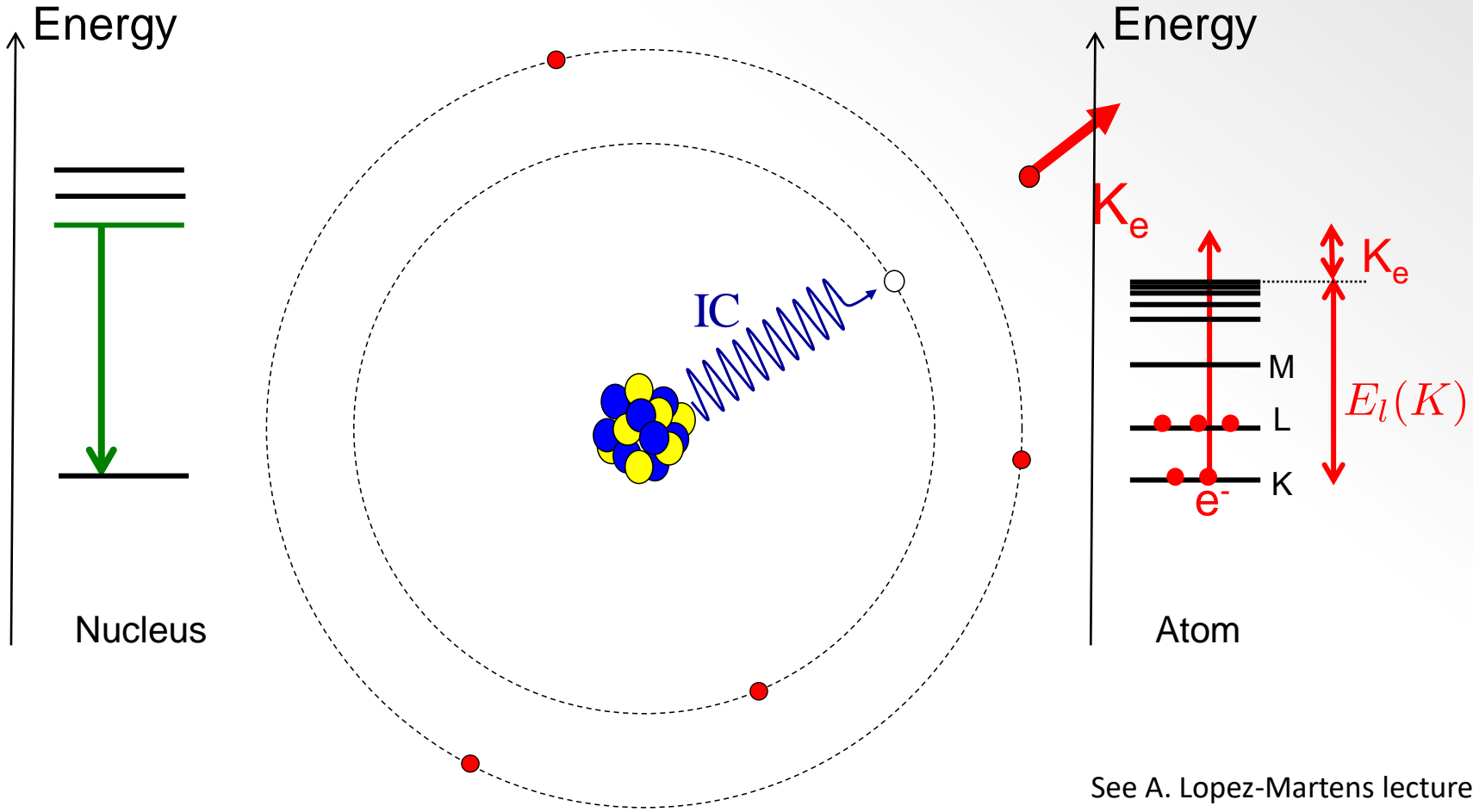
Nuclear excitation / de-excitation

Gamma emission



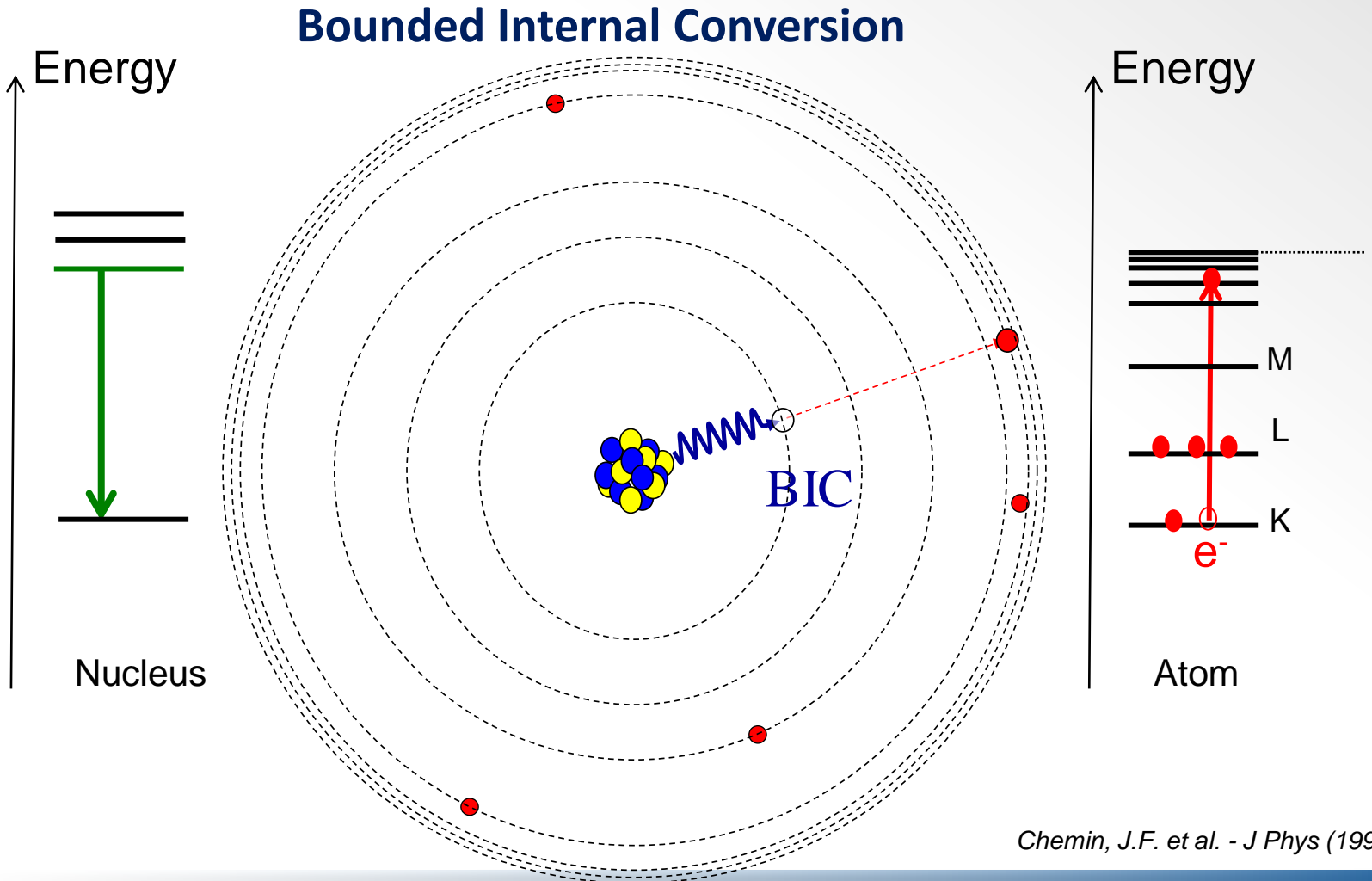
Nuclear excitation / de-excitation

Internal Conversion



See A. Lopez-Martens lecture

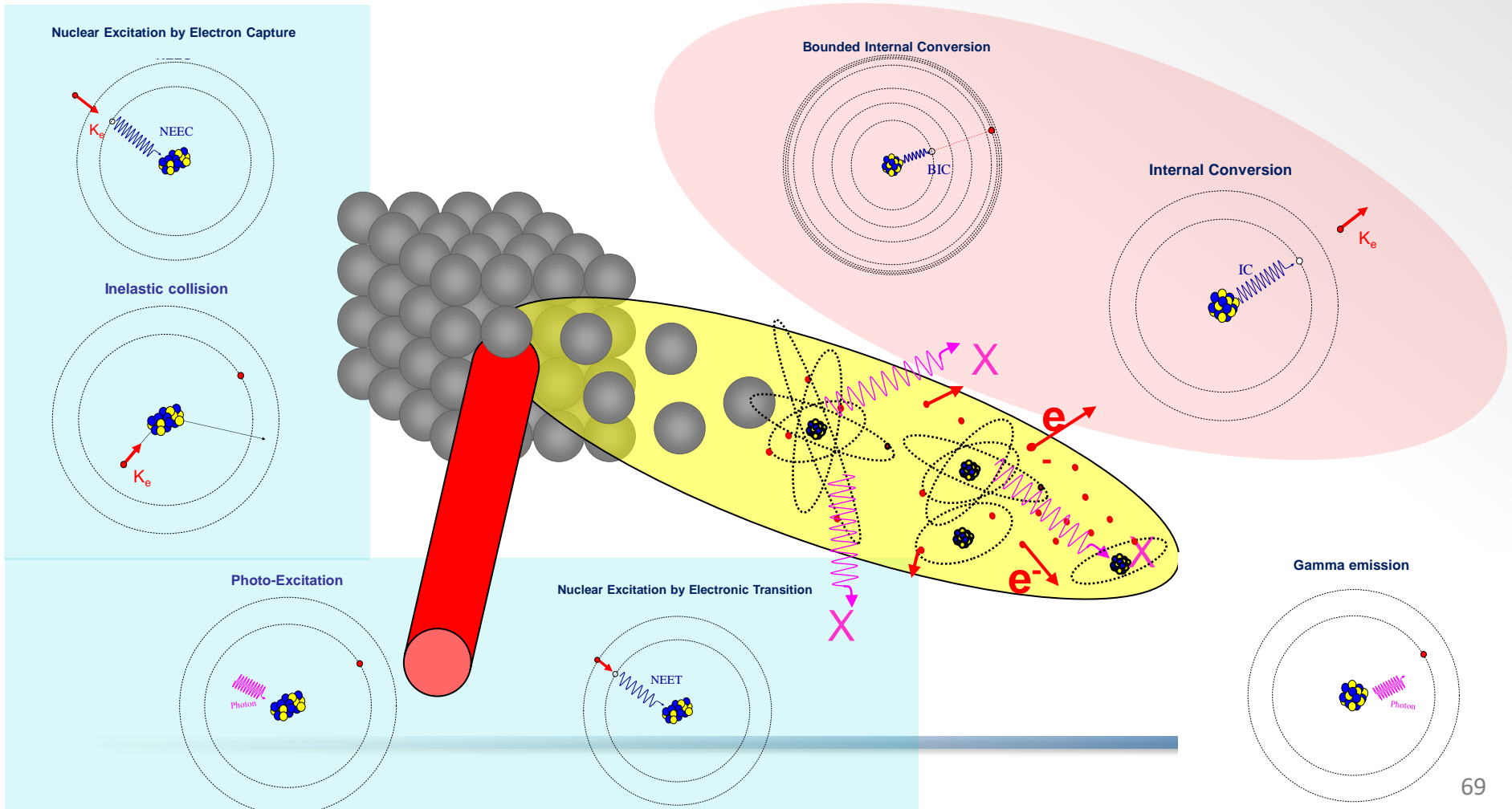
Nuclear excitation / de-excitation



Chemin, J.F. et al. - J Phys (1999) 53: 633.

Nuclear excitation / de-excitation

- Half-life modified because of **de-excitation** processes
- But also because of **excitation** processes



Nuclear excitation / de-excitation

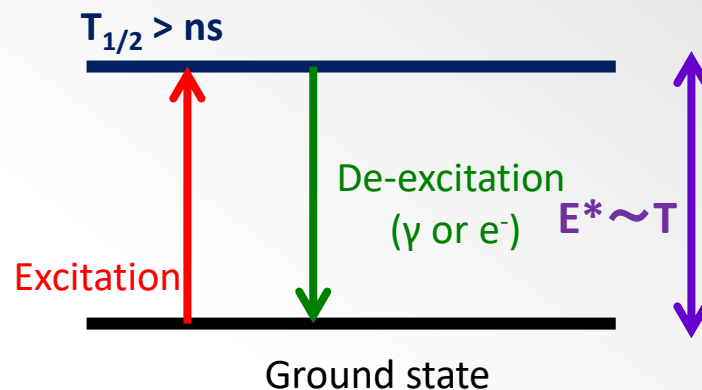
Evidence of nuclear excitation in plasmas

- Laser created plasmas:

- ▶ Temperature (LTE) range of 10 eV- few keV

- Nuclear excitations in plasmas can only be studied in specific nuclei:

- ▶ Low energy excited state: excitation energy $E^* \sim T$
- ▶ Isomeric state: lifetime longer than the plasma emission duration



➤ **Very few candidates (~ 10 stable nuclei with $E^* < 15$ keV)**

	⁴⁵ Sc	¹⁶⁹ Tm	¹⁸¹ Ta	²⁰¹ Hg	⁸³ Kr	⁷³ Ge	⁵⁷ Fe	¹⁸⁷ Os	²³⁵ U	²⁰⁵ Pb
E^* (keV)	14.2	8.4	6.2	1.55	9.4	13.3	14.4	9.7	0.077	2.3
$T_{1/2}$	318 ms	4.1 ns	6.1 μ s	81 ns	154 ns	2.9 μ s	98 ns	2.4 ns	27 m	24 μ s

C.Granja et al. Nucl. Phys. A **784**,1 (2007)


Nuclear excitation / de-excitation


Evidence of nuclear excitation in plasmas


■ Experiments are challenging: Experiments on **stable** nuclei


- ▶ Nuclear excitation cross sections are orders of magnitude smaller than atomic ones
- ▶ Low signal and high background !
- ▶ Isomeric state: weakly coupled state → excitation more difficult for this state


Nucleus	¹⁸¹ Ta	²³⁵ U	⁵⁷ Fe
E* (keV)	6.2	0.077	14.4
T _{1/2}	6.1 μs	27 m	98 ns
Process	Direct excitation	NEET	Direct excitation


 Gobet et al., J.Phys. B **41**, 145701 (2008)


 Spohr et al., Mod Opt **53** 2633 (2006)


 Andreev et al., JETP **91**, 1063 (2000)

 Claverie et al, Phys. Rev. C **70**,044303 (2004)

 Bouns et al. Phys. Rev. C, **46**,852 (1992)

 Arutyunyan et al., Sov.J.NP **53**, 23 (1991)

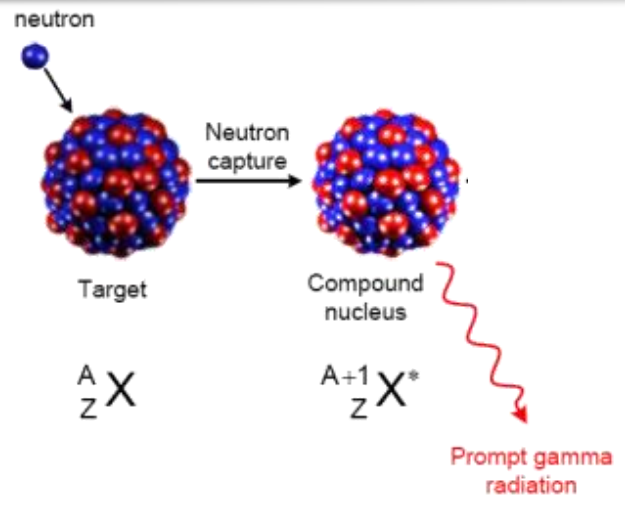
 Izawa et al, Phys. Lett. **88B**,59 (1979)

 Chefonov et al. Laser Phys. **24**, 116002 (2014)
Golovin et al. Quant. Electro. **41**, 222 (2011)

➤ Up to now: no clear evidence of nuclear excitation in plasmas

Cross section modifications on excited nucleus

What would be changed in a plasma?

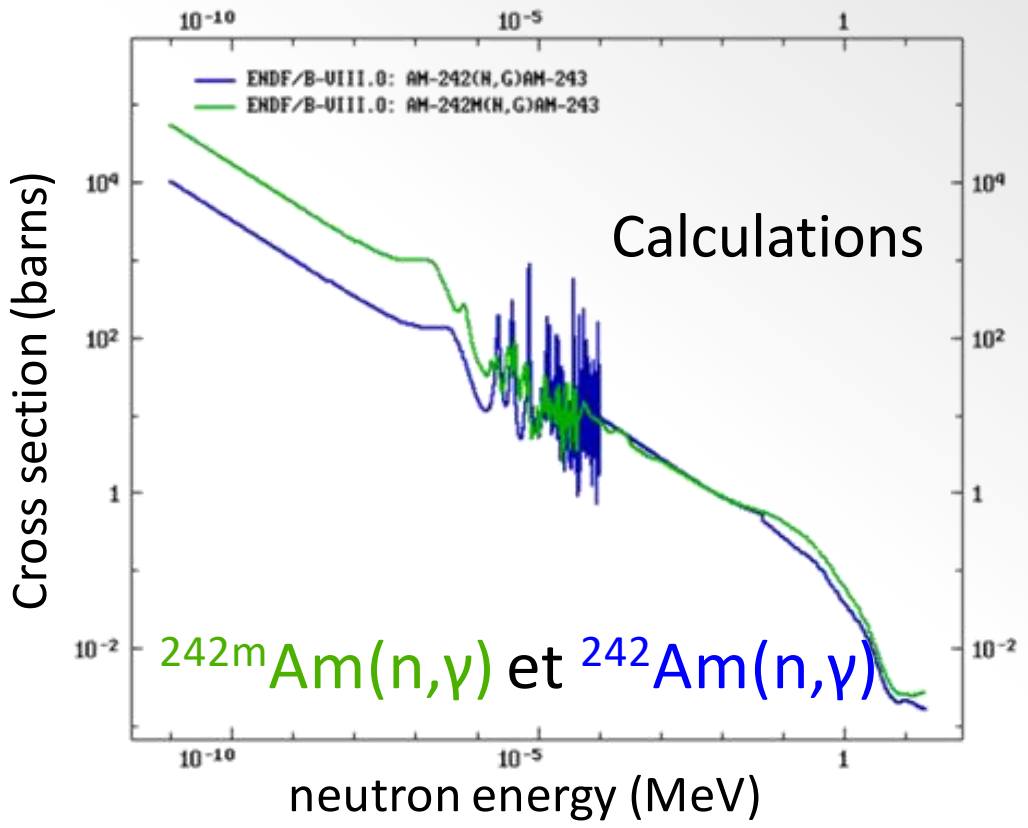


$$\begin{matrix} A \\ Z \end{matrix} X$$

$$\begin{matrix} A+1 \\ Z \end{matrix} X^*$$

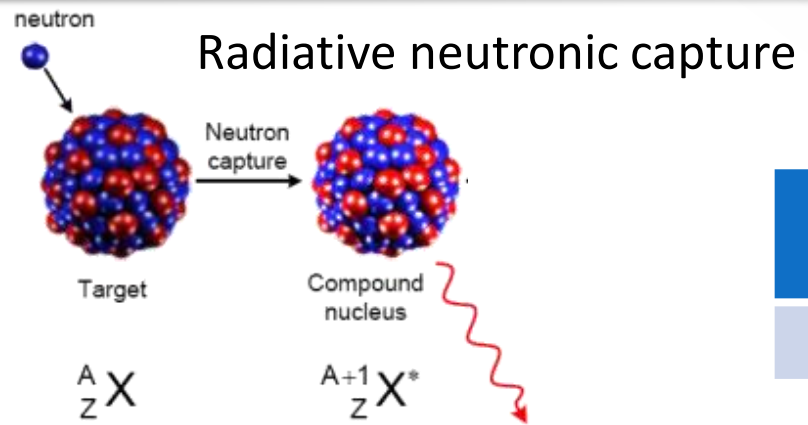
$1/2^+$ ——— $T_{1/2} = 141 \text{ a}$
 48.6 keV

1^- ——— $T_{1/2} = 16 \text{ h}$
 ^{242}Am

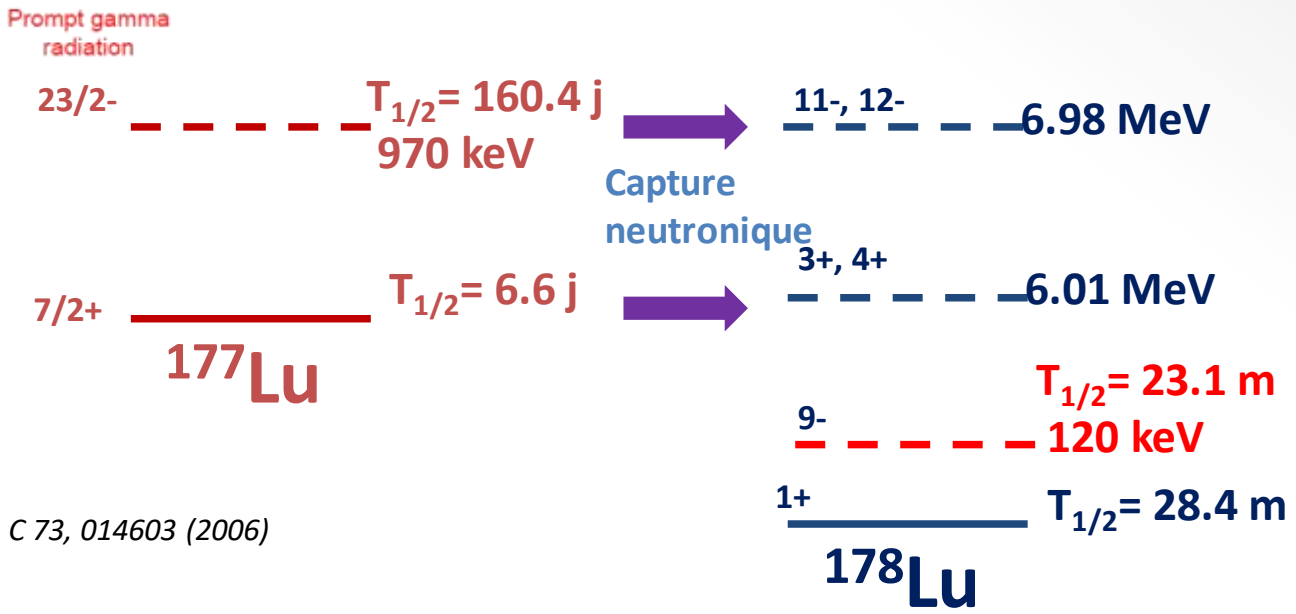


Cross section modifications on excited nucleus

What would be changed in a plasma?



Neutron energy	${}^{177m}\text{Lu}(n, \gamma){}^{178m}\text{Lu} / {}^{177}\text{Lu}(n, \gamma){}^{178m}\text{Lu}$
< 25 meV	0,47 +/- 0,07

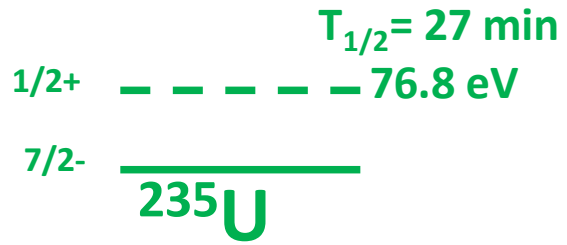
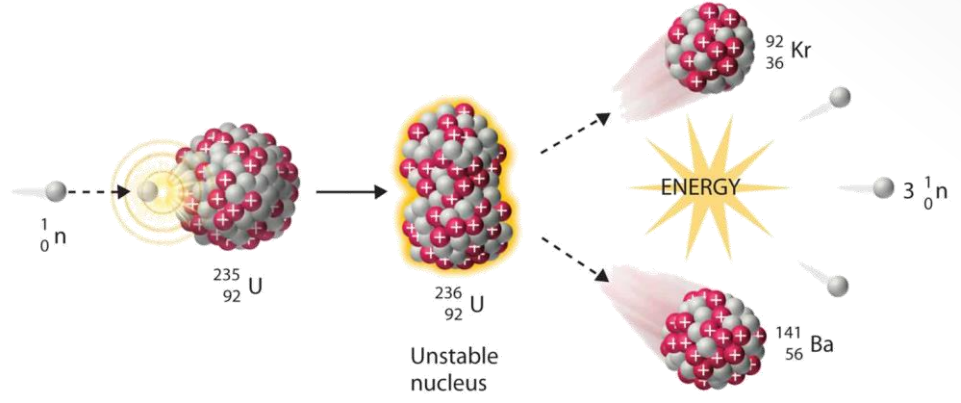


G. Bélier et al., Phys. Rev. C 73, 014603 (2006)

Cross section modifications on excited nucleus

What would be changed in a plasma?

Neutron-capture-induced Fission



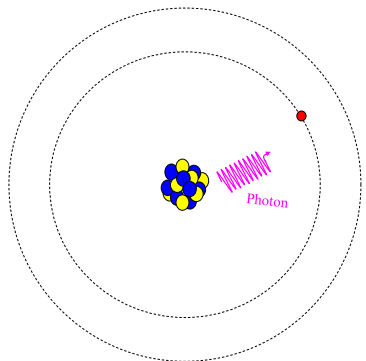
Neutron energy	${}^{235m}\text{U}(n,f) / {}^{235}\text{U}(n,f)$
< 25 meV	1,61 +/- 0,44
50 meV	2,47 +/- 0,45

A. D'Eer et al, Phys. Rev. C 38, 1270 (1988)

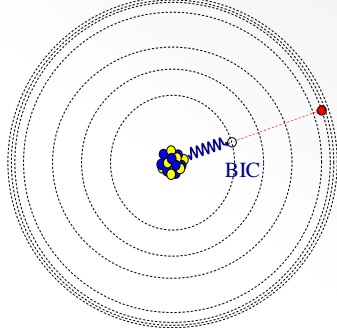
Half-life of an excited state

What would be changed in a plasma?

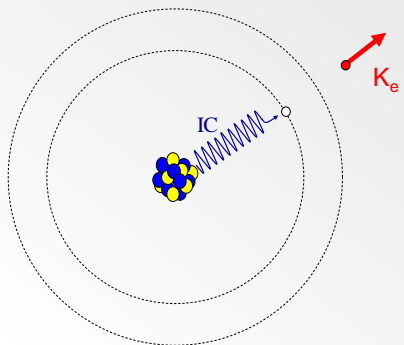
Gamma emission



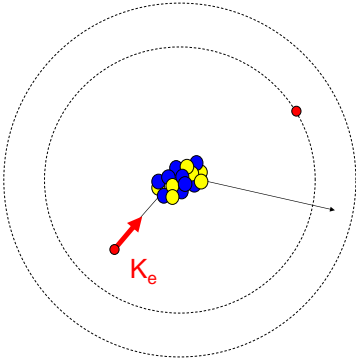
Bounded Internal Conversion



Internal Conversion



Inelastic collision



$$T_{1/2} = \frac{\ln(2)}{\lambda_\gamma + \sum_i \lambda_i}$$

Nuclear Excitation by Electron Capture

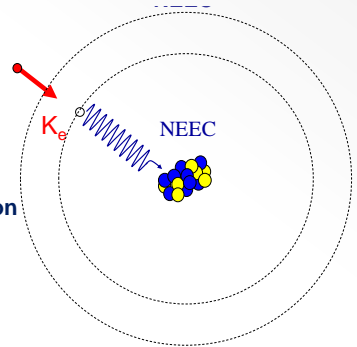
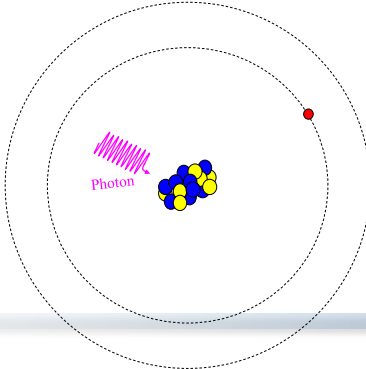
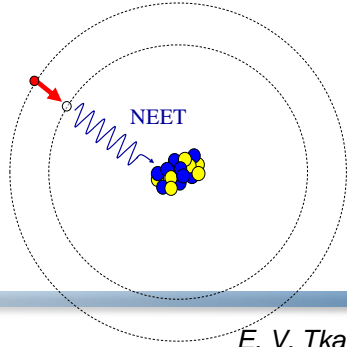


Photo-Excitation



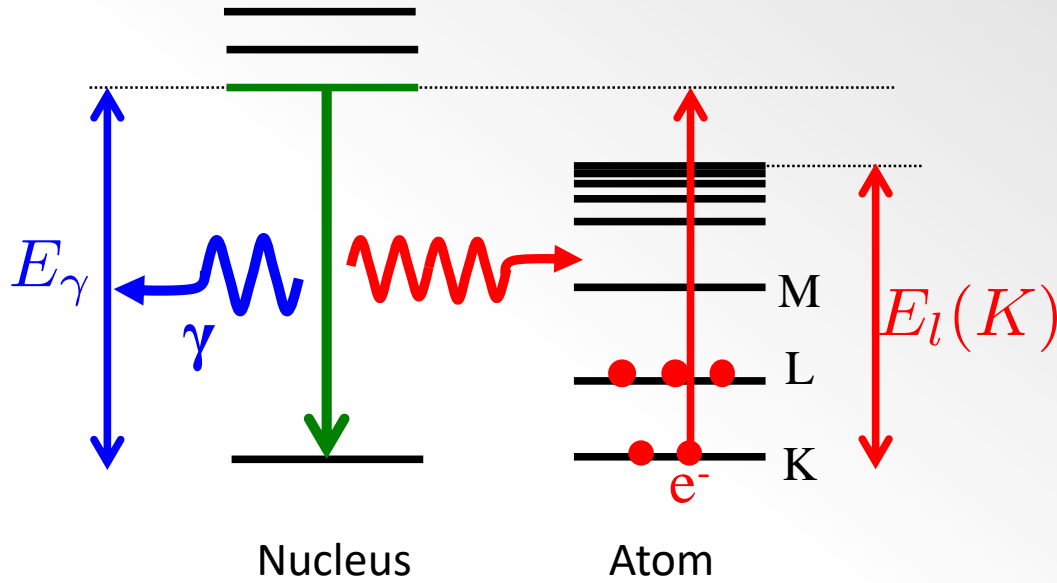
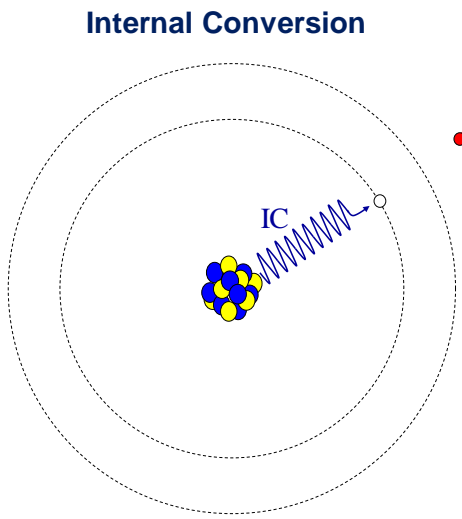
Nuclear Excitation by Electronic Transition



Half-life of an excited state

$$T_{1/2} = \frac{\ln(2)}{\lambda_\gamma + \lambda_e}$$

Ion charge effect



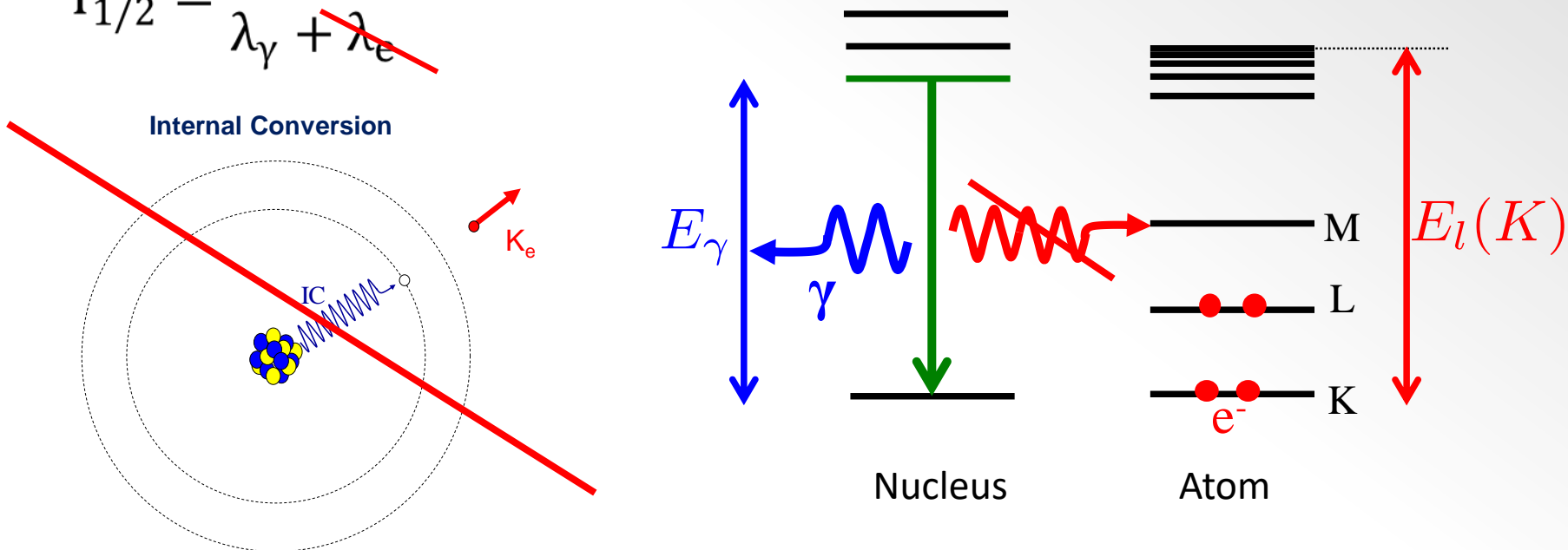
$^{125}_{52}\text{Te}$ 1^{er} état excité à 35,5 keV

Q	$T_{1/2}$ (ns)	$E_l(K)$ (keV)
0 (neutral)	1,49	31,8
48 ⁺	11 ± 2	36,6

Half-life of an excited state

$$T_{1/2} = \frac{\ln(2)}{\lambda_\gamma + \lambda_e}$$

Ion charge effect



$^{125}_{52}\text{Te}$ 1^{er} état excité à 35,5 keV

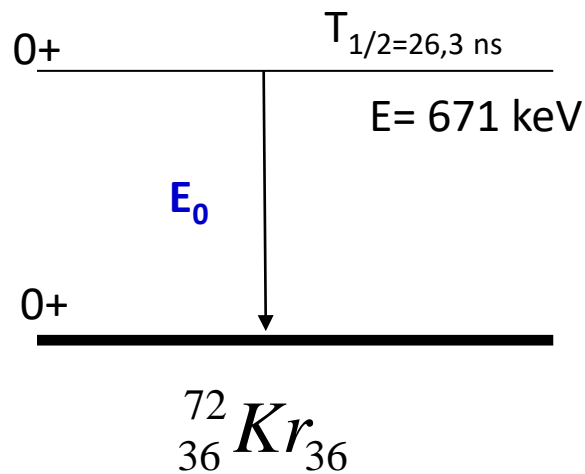
Q	$T_{1/2}$ (ns)	$E_l(K)$ (keV)
0 (neutral)	1,49	31,8
48 ⁺	11 ± 2	36,6

Charge state can modify apparent nuclear properties

Half-life of an excited state

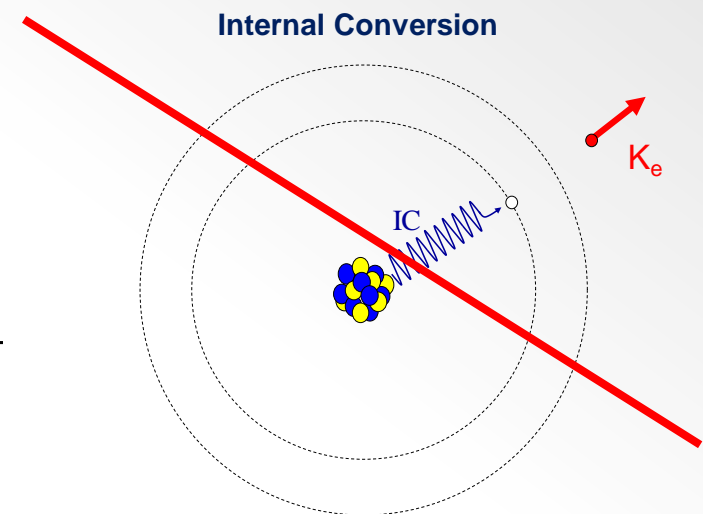
E. Bouchez et al, Phys. Rev. Lett. 90, 082502 (2003)

fully stripped ion



$0^+ - 0^+$ transition :
No γ allowed

$$T_{1/2} = \frac{\ln(2)}{\lambda_{\text{IC}}}$$



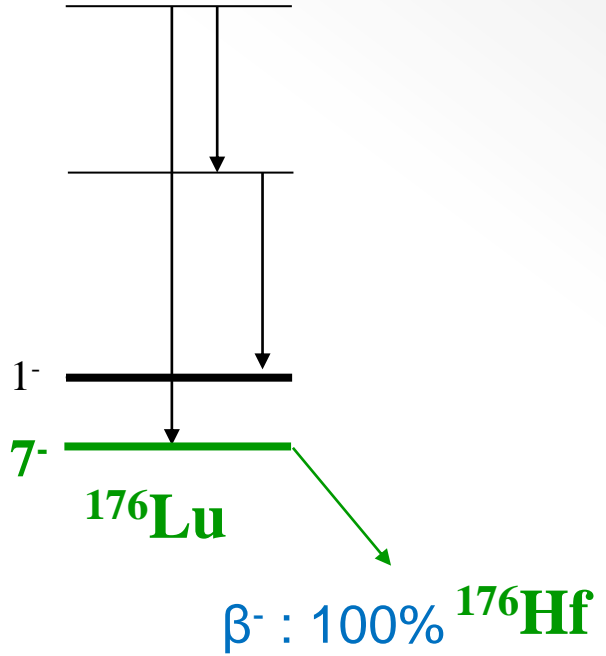
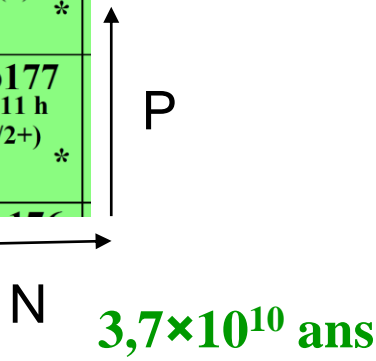
Isomer state $E = 671 \text{ keV}$

- $T_{1/2} = 26.3 \text{ ns}$ in neutral atom
- « Stable » in fully stripped ion : Kr^{36+}

Half-life of an excited state

Astrophysical consequences : S process and abundances

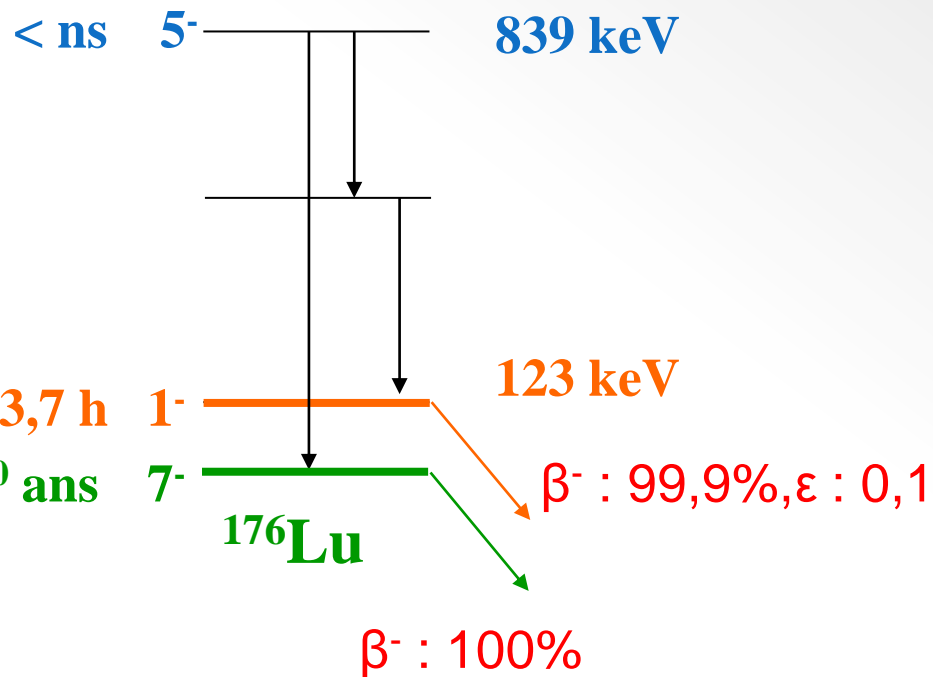
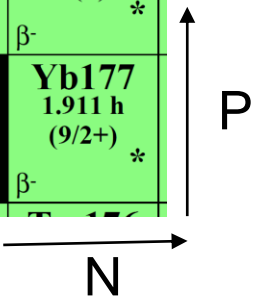
Ta176 8.09 h (1)- * EC	Ta177 56.56 h 7/2+ EC	Ta178 9.31 m 1+ * EC	Ta179 1.82 y 7/2+ * EC	Ta180 8.152 h 1+ * EC,β _{0.012} *
Hf175 70 d 5/2- * EC	Hf176 0+ 5.206	Hf177 7/2- * 18.606	Hf178 0+ * 27.297	Hf179 9/2+ * 13.629
Lu174 3.31 y (1)- * EC	Lu175 7/2+ 97.41	Lu176 3.78E10 y 7- * β ⁻ 2.59	Lu177 6.734 d 7/2+ * β ⁻	Lu178 28.4 m 1(+) * β ⁻
Yb173 5/2- 16.12	Yb174 0+ 31.8	Yb175 4.185 d 7/2- * β ⁻	Yb176 0+ * 12.7	Yb177 1.911 h (9/2+) * β ⁻



Half-life of an excited state

Astrophysical consequences : S process and abundances

Ta176 8.09 h (1)- * EC	Ta177 56.56 h 7/2+ EC	Ta178 9.31 m 1+ * EC	Ta179 1.82 y 7/2+ * EC	Ta180 8.152 h 1+ * EC,β ⁻ 0.012
Hf175 70 d 5/2- * EC	Hf176 0+ 5.206	Hf177 7/2- * 18.606	Hf178 0+ * 27.297	Hf179 9/2+ * 13.629
Lu174 3.31 y (1)- * EC	Lu175 7/2+ 97.41	Lu176 7- * 3.8E10 y β ⁻ 2.59	Lu177 7/2+ * 6.734 d β ⁻	Lu178 1(+) * 28.4 m β ⁻
Yb173 5/2- 16.12	Yb174 0+ 31.8	Yb175 7/2- * 4.185 d β ⁻	Yb176 0+ * 12.7	Yb177 9/2+ * 1.911 h β ⁻



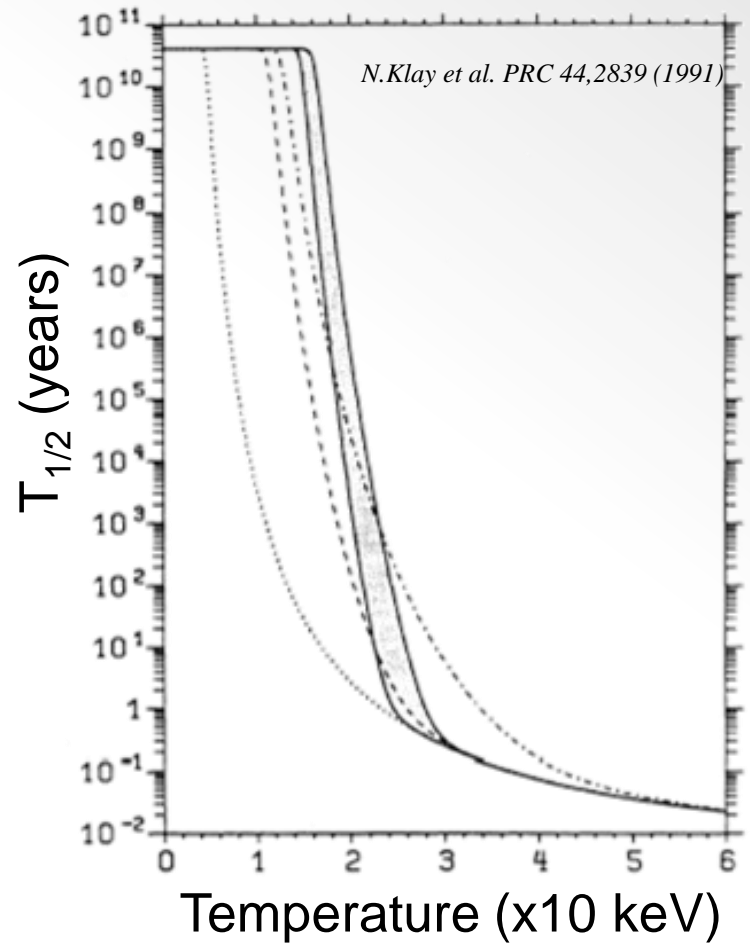
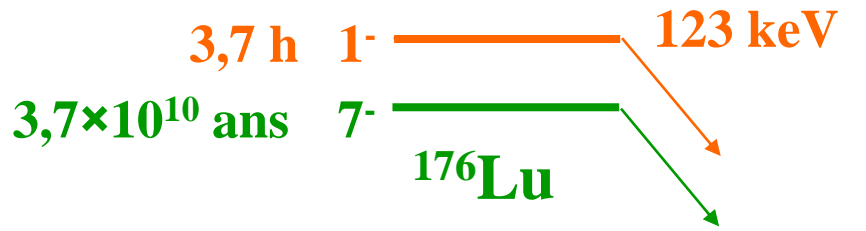
Half-life of an excited state

Astrophysical consequences : S process and abundances

Ta176 8.09 h (1)- * EC	Ta177 56.56 h 7/2+ EC	Ta178 9.31 m 1+ * EC	Ta179 1.82 y 7/2+ * EC	Ta180 8.152 h 1+ * EC,β _{0.012}
Hf175 70 d 5/2- * EC	Hf176 0+ 5.206	Hf177 7/2- * 18.606	Hf178 0+ * 27.297	Hf179 9/2+ * 13.629
Lu174 3.31 y (1)- * EC	Lu175 7/2+ 97.41	Lu176 3.8E10 y 7- * β _{2.59}	Lu177 6.734 d 7/2+ * β ₋	Lu178 28.4 m 1(+) * β ₋
Yb173 5/2- 16.12	Yb174 0+ 31.8	Yb175 4.185 d 7/2- * β ₋	Yb176 0+ * 12.7	Yb177 1.911 h (9/2+) * β ₋

N →

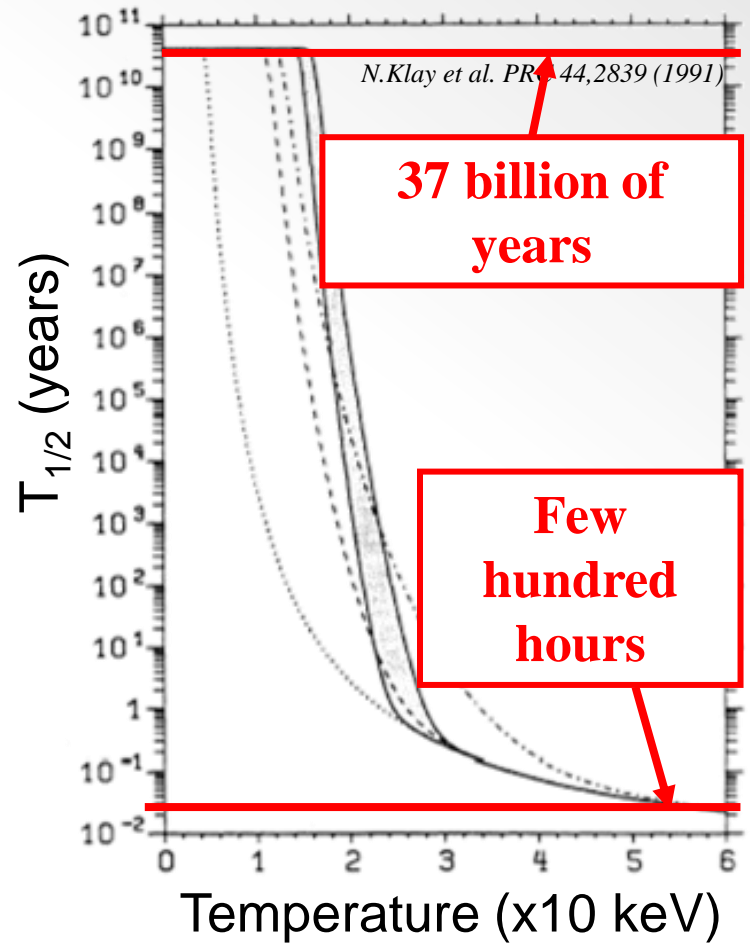
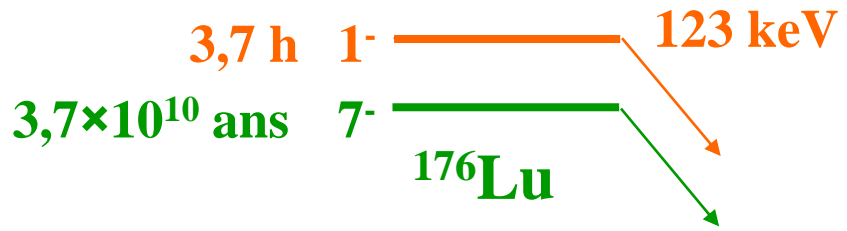
↑ P



Half-life of an excited state

Astrophysical consequences : S process and abundances

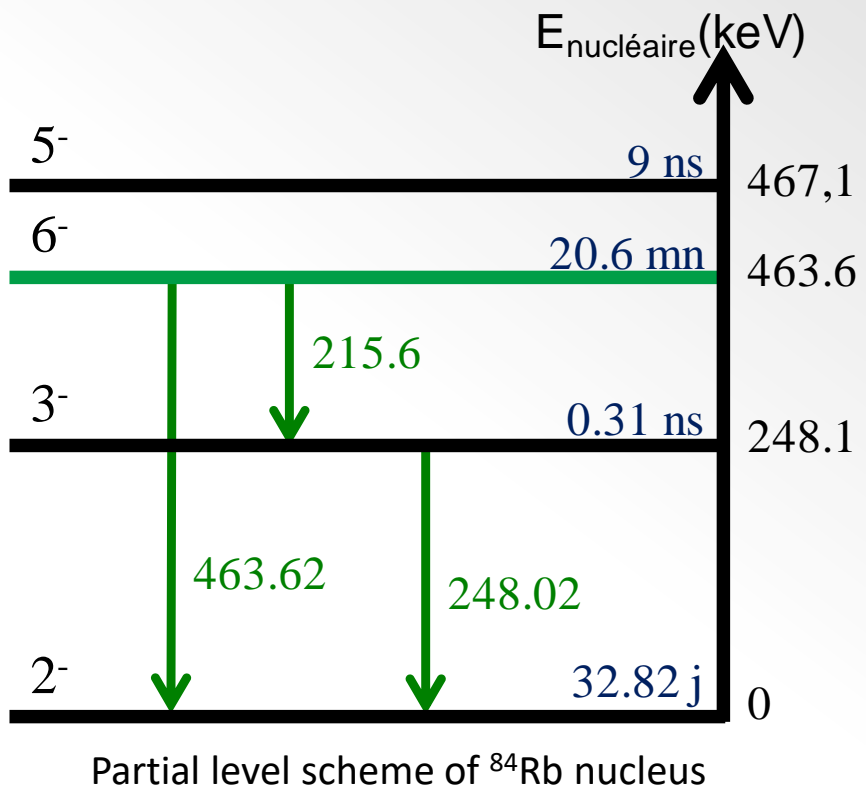
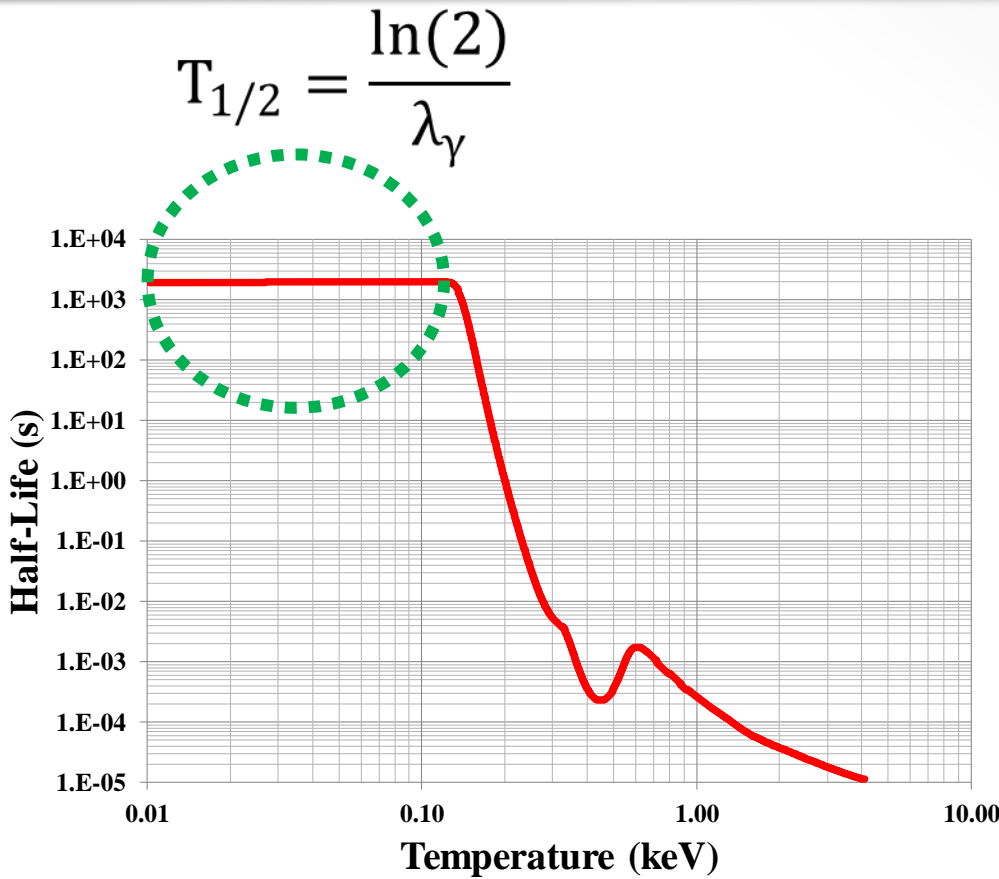
Ta176 8.09 h (1)- * EC	Ta177 56.56 h 7/2+ EC	Ta178 9.31 m 1+ * EC	Ta179 1.82 y 7/2+ * EC	Ta180 8.152 h 1+ * EC,β ⁻ _{0.012}
Hf175 70 d 5/2- * EC	Hf176 0+ 5.206	Hf177 7/2- * 18.606	Hf178 0+ * 27.297	Hf179 9/2+ * 13.629
Lu174 3.31 y (1)- * EC	Lu175 7/2+ 97.41	Lu176 3.8E10 y 7- * β ⁻ _{2.59}	Lu177 6.734 d 7/2+ * β ⁻	Lu178 28.4 m 1(+) * β ⁻
Yb173 5/2- 16.12	Yb174 0+ 31.8	Yb175 4.185 d 7/2- * β ⁻	Yb176 0+ * 12.7	Yb177 1.911 h (9/2+) * β ⁻



The abundance ratio between $^{176}\text{Hf} / ^{177}\text{Hf}$ is modified in a hot plasma

Half-life of an excited state

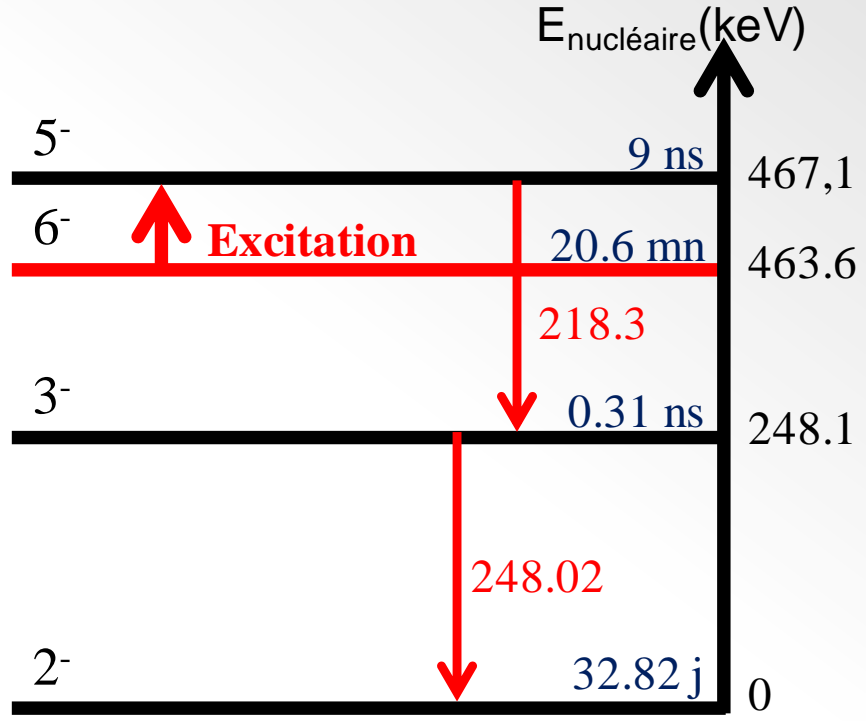
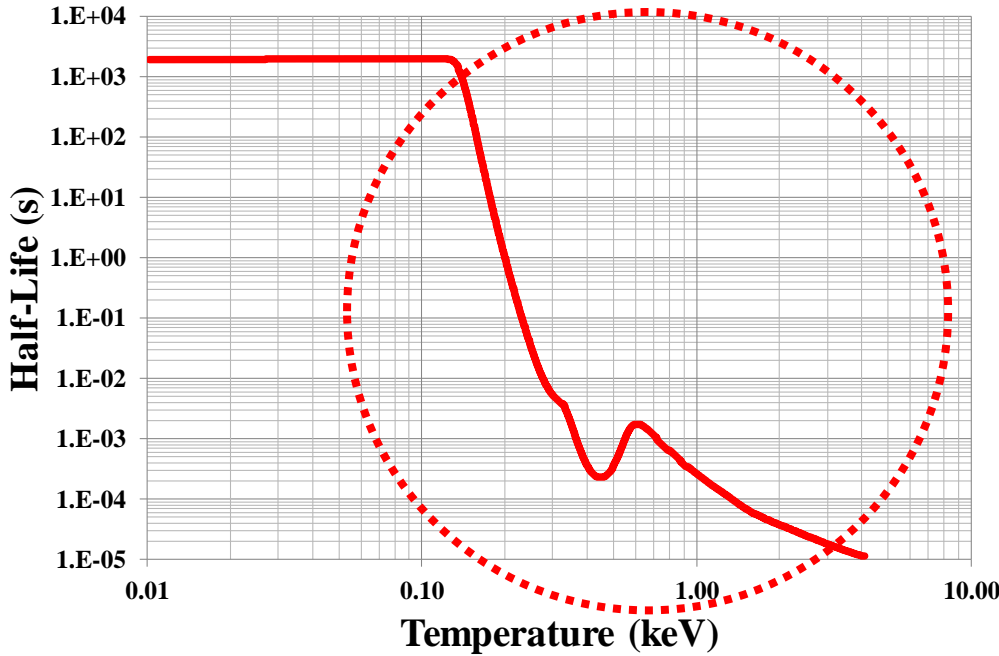
Demonstration of $T_{1/2}$ modification in a plasma : ^{84}Rb case



Half-life of an excited state

Demonstration of $T_{1/2}$ modification in a plasma : ^{84}Rb case

$$T_{1/2} = \frac{\ln(2)}{\lambda_\gamma + \lambda_{\text{excit}}}$$

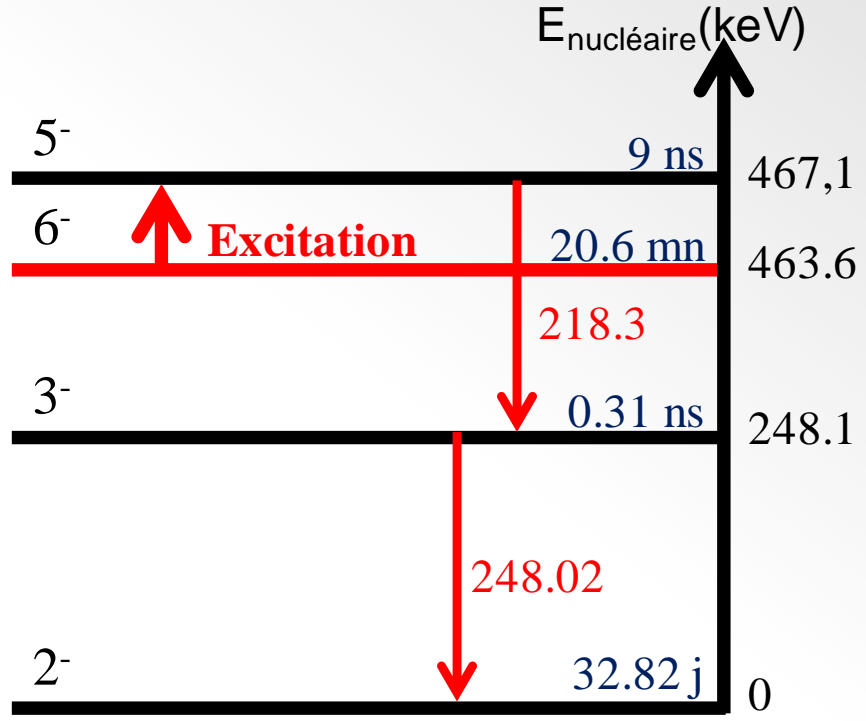
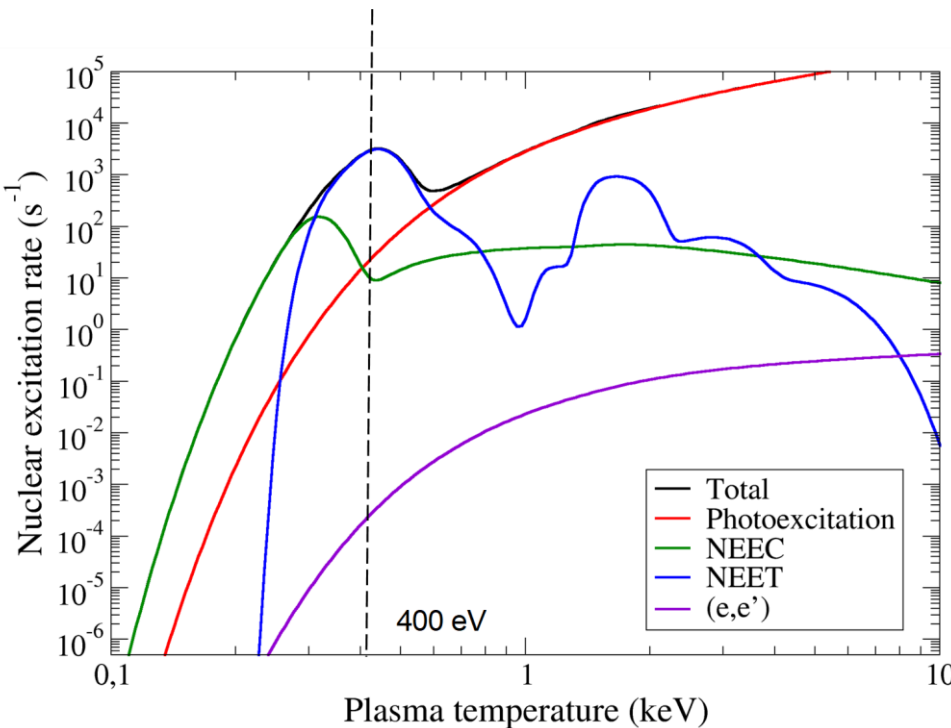


Partial level scheme of ^{84}Rb nucleus

Half-life of an excited state

Demonstration of $T_{1/2}$ modification in a plasma : ^{84}Rb case

$$T_{1/2} = \frac{\ln(2)}{\lambda_\gamma + \lambda_{\text{excit}}}$$

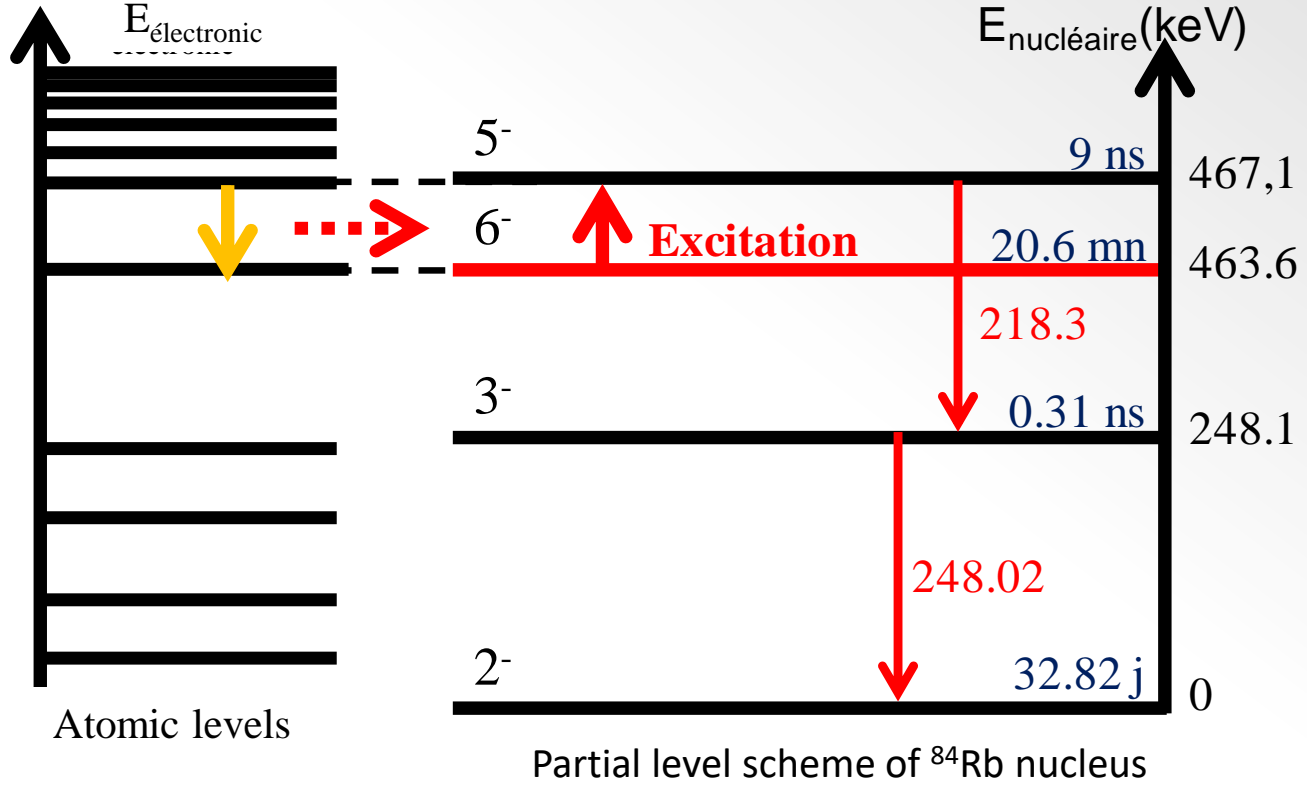


Partial level scheme of ^{84}Rb nucleus

ISOMEX code based on a **Relativistic Average Atom Model** under LTE hypothesis.
 All the ions in plasmas are described by one average ion : still valid for resonant processes in light nucleus?

Half-life of an excited state

NEET : Nuclear Excitation by Electronic Transition



- Difference with photoexcitation: takes place in a single ion
- Nuclear and atomic transitions must be resonant
- Transitions with same multipolarities
- Observed in neutral target: ¹⁹⁷Au, ¹⁸⁹Os and ¹⁹³Ir

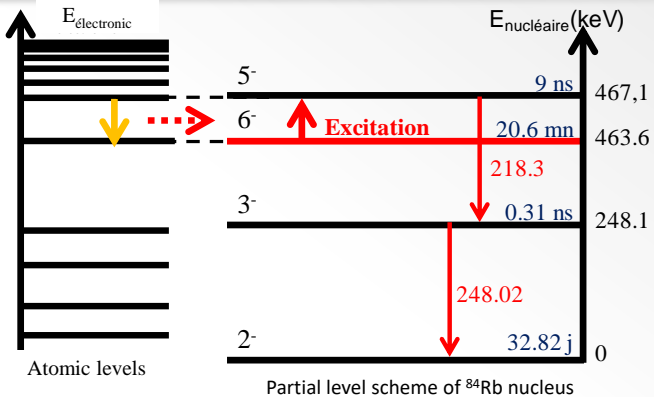
Kishimoto et al., Phys. Rev. Lett, 85, 1831 (2000)
Ahmad et al., Phys. Rev. C, 61, 051304 (2000)
Kishimoto et al., Nucl. Phys. A, 748, 3 (2005)

- Not observed in plasmas: predictions
- Morel et al., Phys. Rev. A, 69, 063414 (2004)*
Harston et al., Phys. Rev. C, 59, 2462 (1999)

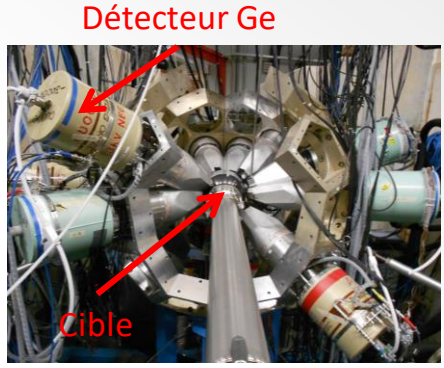
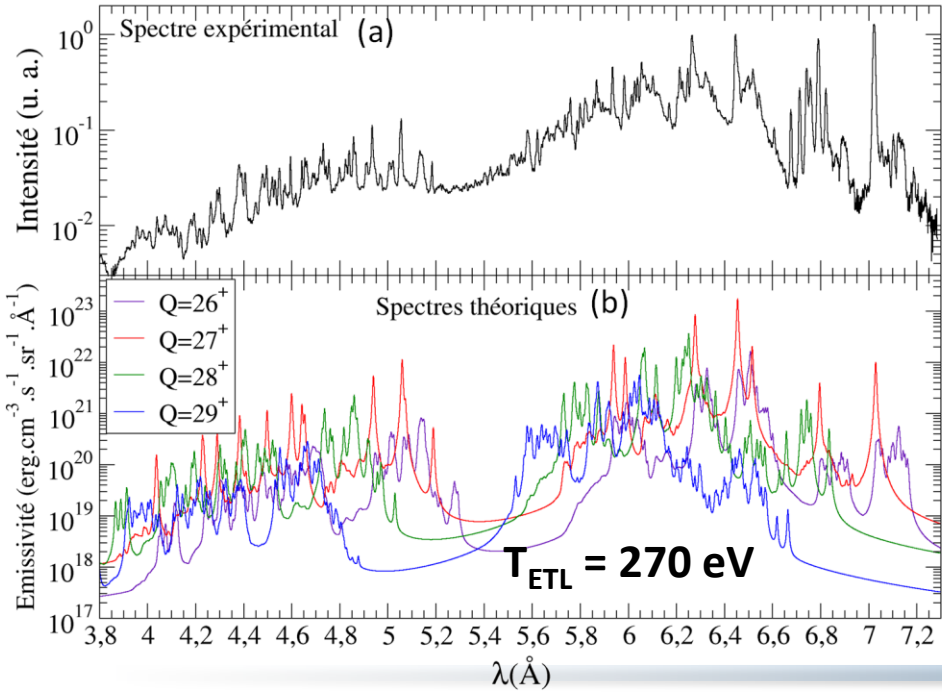
Half-life of an excited state

NEET rate estimation

Measurement of X
Spectrum from a Rb
plasma on PHELIX



Measurement of γ
transition on ALTO /
ORSAY



Transition	Nudat (keV)	Our results (keV)
5- → 6-	2,69 ± 0,23 (γ) 3,05 ± 0,18 (levels) 3,4 (suggested)	3,498 ± 0,006

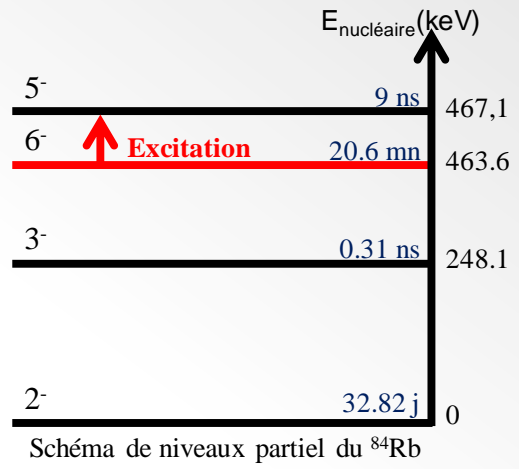
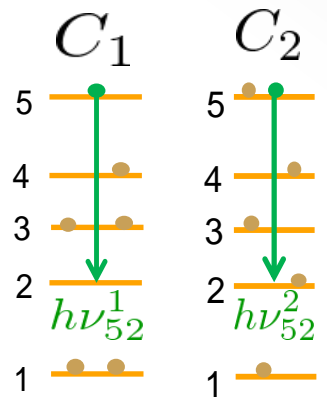
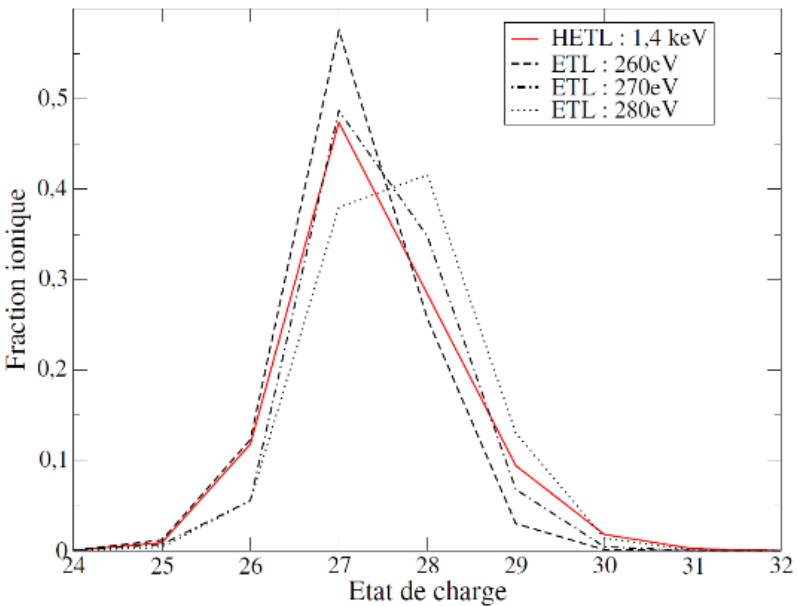
D. Denis-Petit et al., Phys. Rev. C 96, 024604 (2017)

D. Denis-Petit et al., Journal of Quantitative Spectroscopy and Radiative Transfer 148 70-89 (2014)

Half-life of an excited state

NEET rate estimation

$$\lambda_{\text{NEET}} = \boxed{\text{Plasma}} + \boxed{\text{Atom / ion}} + \boxed{\text{Nucleus}}$$



Ionization Temperature (T_Z) model

$$\overline{Q}_{\text{ETL}}(\rho, T_Z) = \overline{Q}_{\text{NETL}}(\rho, T_e)$$

→ Saha-Boltzmann distribution for electronic population

Multi-Configuration Dirac-Fock (MCDF) calculations of electronic configurations :

- atomic levels
- atomic transitions

Shell model calculations of transition probability $B_{5 \rightarrow 6^-}$ (M1)

Half-life of an excited state

NEET rate estimation

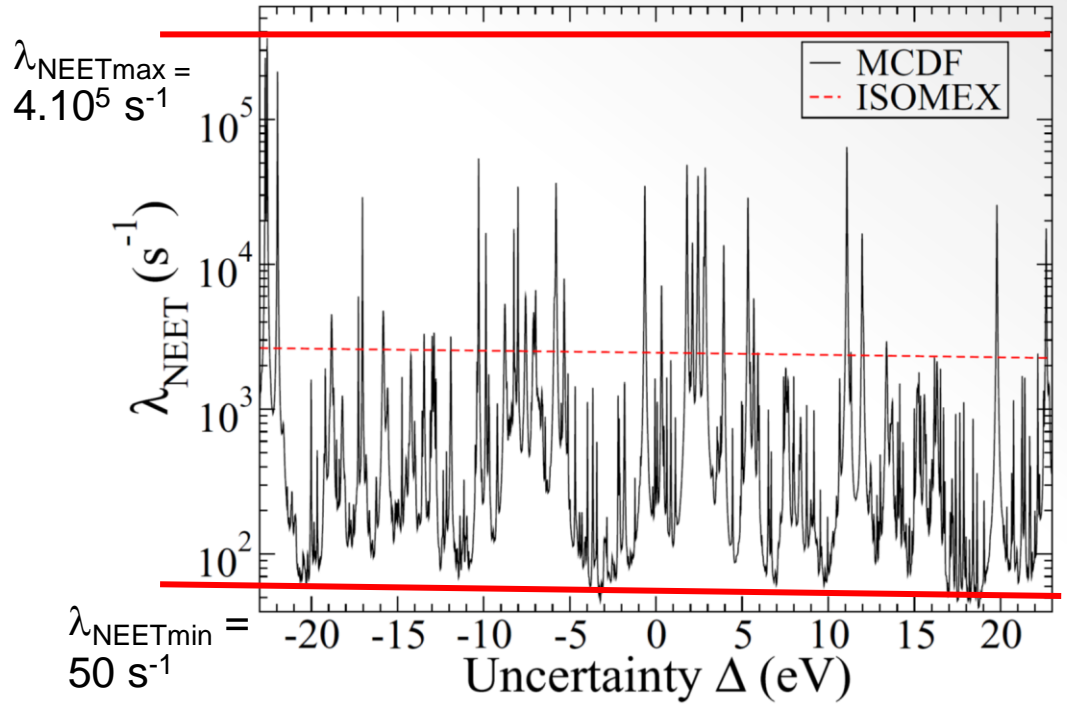
- NEET rate evolution depending of uncertainty parameter Δ :

$$\left. \begin{array}{l} E_{\text{nuclear}} \pm \Delta E_{\text{nuclear}} \\ E_{\text{electronic}} \pm \Delta E_{\text{electronic}} \end{array} \right\}$$

$$\Delta^2 = \Delta E_{\text{nuclear}}^2 + \Delta E_{\text{electronic}}^2$$

Plasma :
 Q=29+ to 34+
 $T_Z = 400 \text{ eV}$ (LTE)
 $T_e > 2 \text{ keV}$ (NLTE)
 $\rho = 10^{-2} \text{ g/cm}^3$

Quantity	Value	Uncertainty
Atomic line energy	MCDF	5 eV at 1σ
Nuclear line energy	3,498 keV	6 eV at 1σ
Atomic line width	Baranger	Factor 10
Reduced transition probability	0.08 W.u.	Factor 2



D. Denis-Petit et al. Phys. Rev. C 96, 024604 (2017)

M. Baranger, Phys. Rev. 112, 855 (1958)

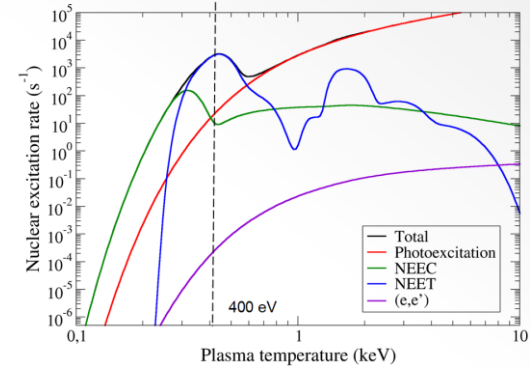
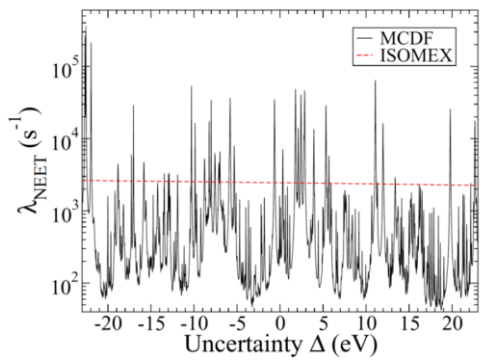
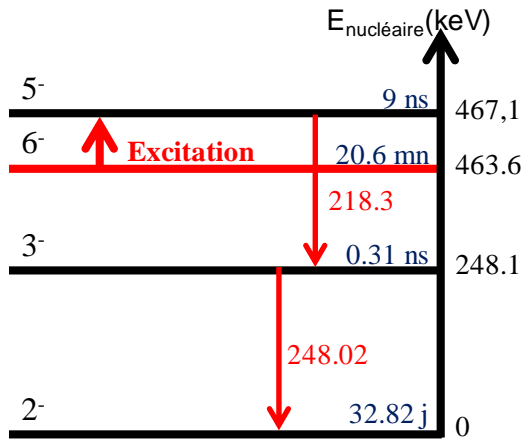
Half-life of an excited state

NEET rate estimation

- NEET rate evolution depending of uncertainty parameter Δ :

$$\left. \begin{aligned} E_{\text{nuclear}} \pm \Delta E_{\text{nuclear}} \\ E_{\text{electronic}} \pm \Delta E_{\text{electronic}} \end{aligned} \right\}$$

$$\Delta^2 = \Delta E_{\text{nuclear}}^2 + \Delta E_{\text{electronic}}^2$$



Plasma Temperature	$\lambda_{6 \rightarrow 5}$ min	$\lambda_{6 \rightarrow 5}$ max
270 eV LTE	$\sim 60 \text{ s}^{-1}$	$\sim 500 \text{ s}^{-1}$
400 eV LTE	$\sim 30 \text{ s}^{-1}$	$\sim 10^6 \text{ s}^{-1}$

Calculations not precise enough and can not deny the excitation \rightarrow experiment needed

D. Denis-Petit et al. Phys. Rev. C 96, 024604 (2017)

Half-life of an excited state

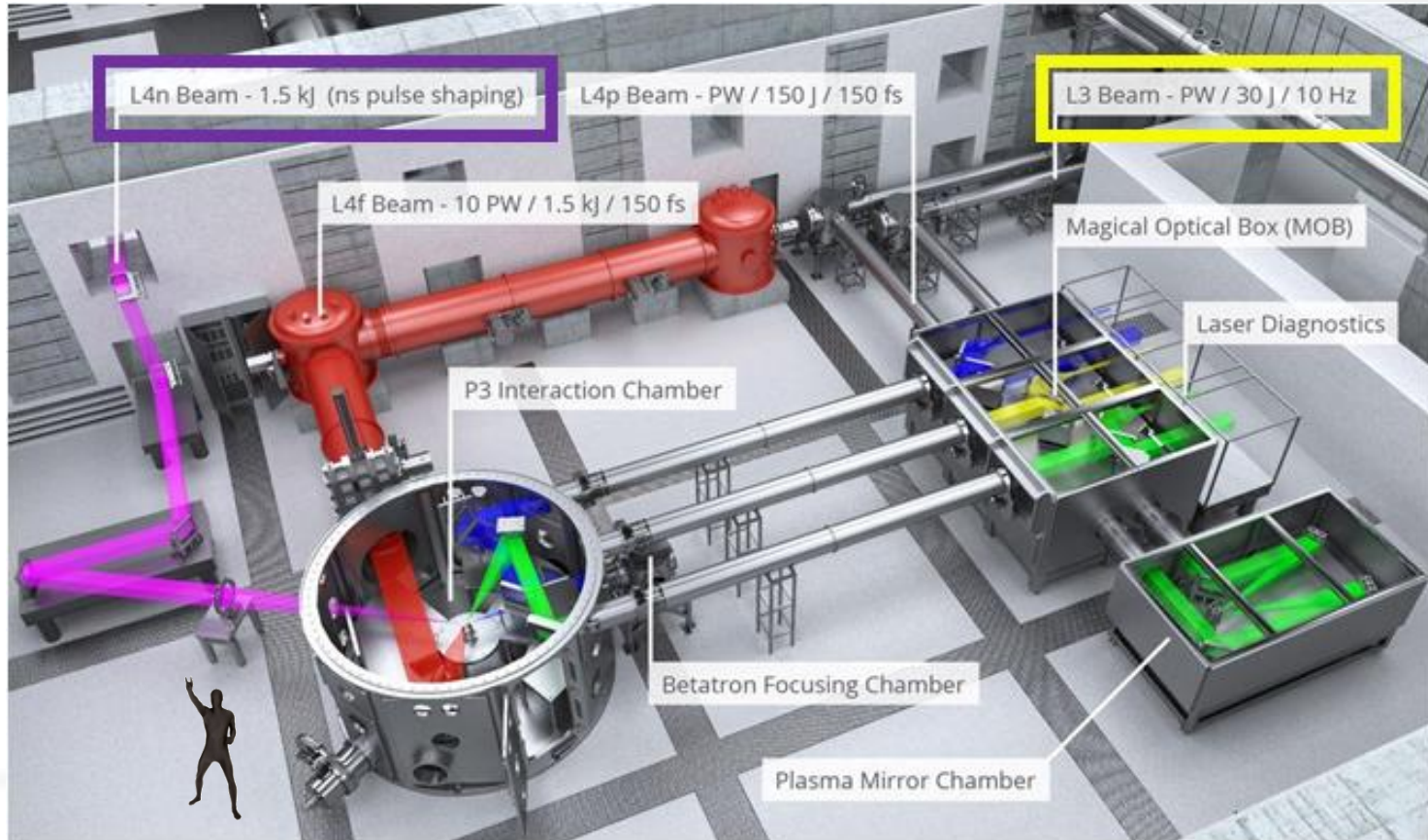
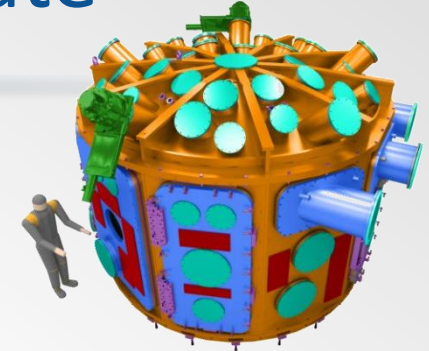
^{84}Rb experiment



ELI-Beamlines, Prague

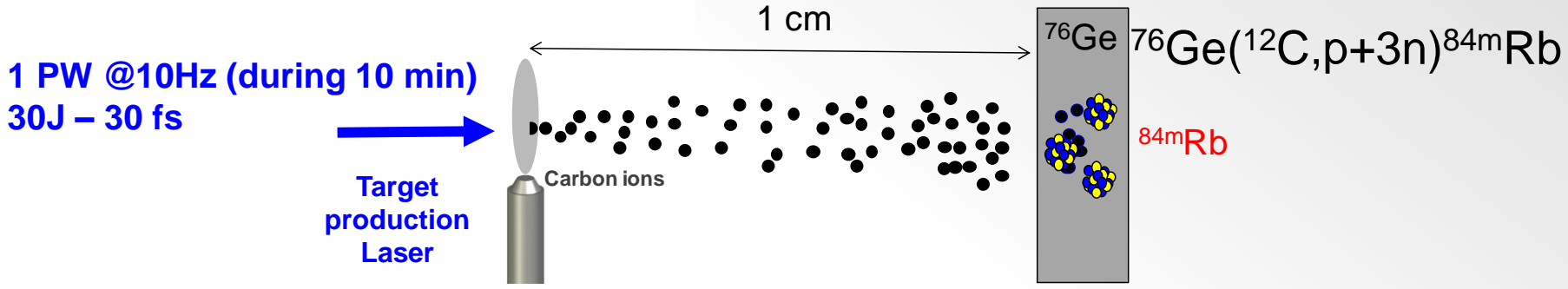
30J ; 30 fs @10Hz

1.5 kJ ; 1 ns @ 1 tir/min



Half-life of an excited state

⁸⁴Rb experiment

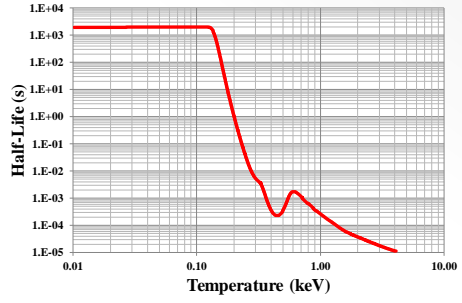
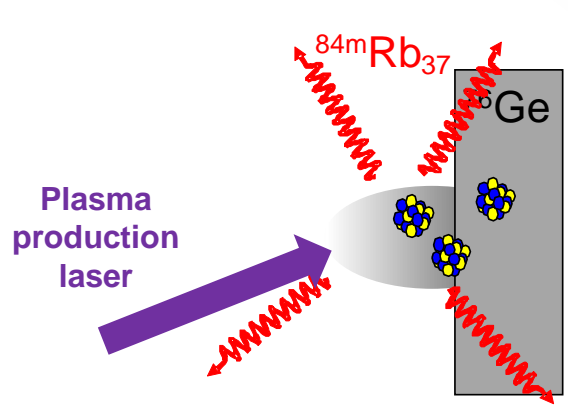


~10⁸ nuclei of ^{84m}Rb produced in a ~5 mm diameter and 3 μm thick layer

1.5 TW / 10min
1.5 kJ – 1ns

Hypothesis on plasma characteristics:

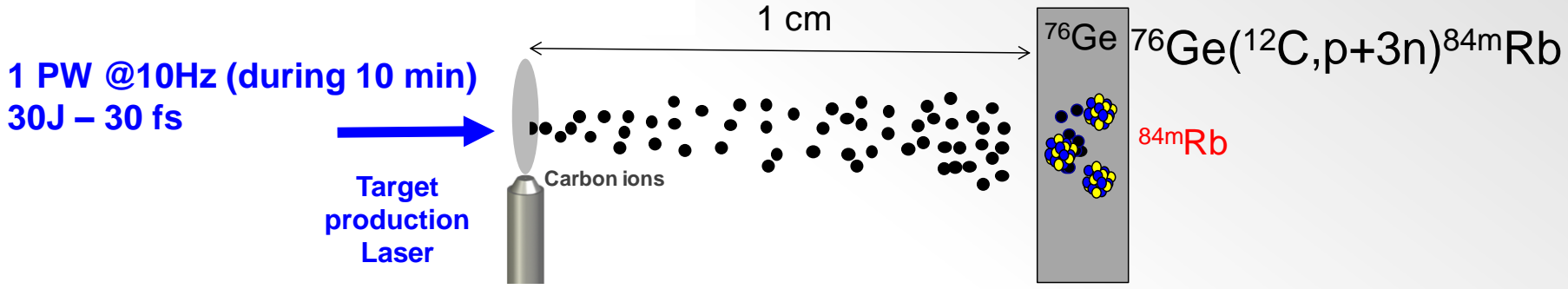
30 μm thickness
250 μm diameter
Q=29+ to 34+
ρ = 10⁻² g/cm³ } during 30 ps



Per shot : ~3 <N_{de-ex}< ~3000
Per day: ~100 <N_{de-ex}< ~400 000
with 144 cycles / day

Half-life of an excited state

⁸⁴Rb experiment

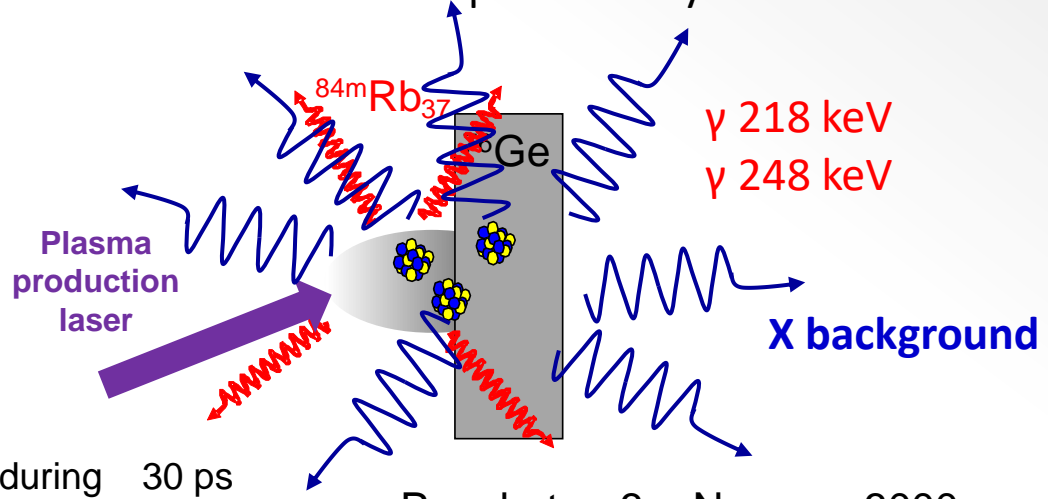


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30 μm thickness
250 μm diameter
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ρ = 10⁻² g/cm³ } during 30 ps



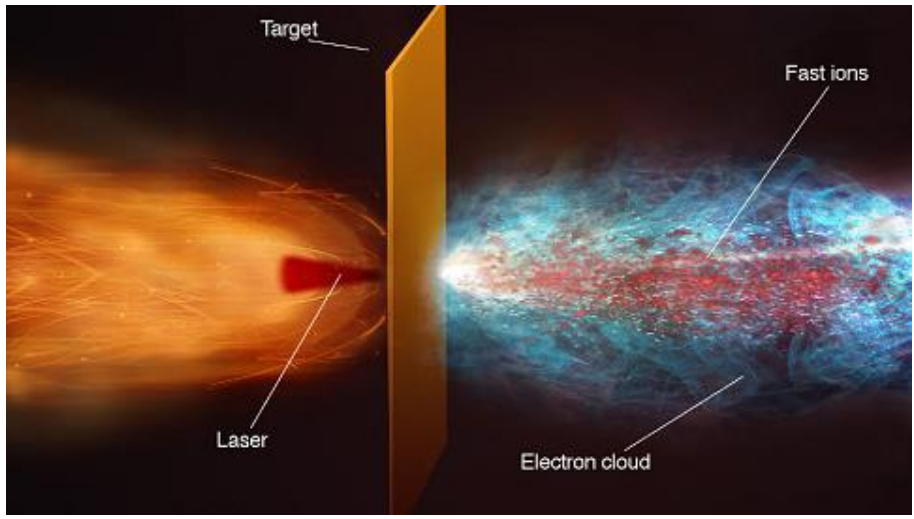
Per shot : ~3 <N_{de-ex}< ~3000
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with 144 cycles / day

Part 4

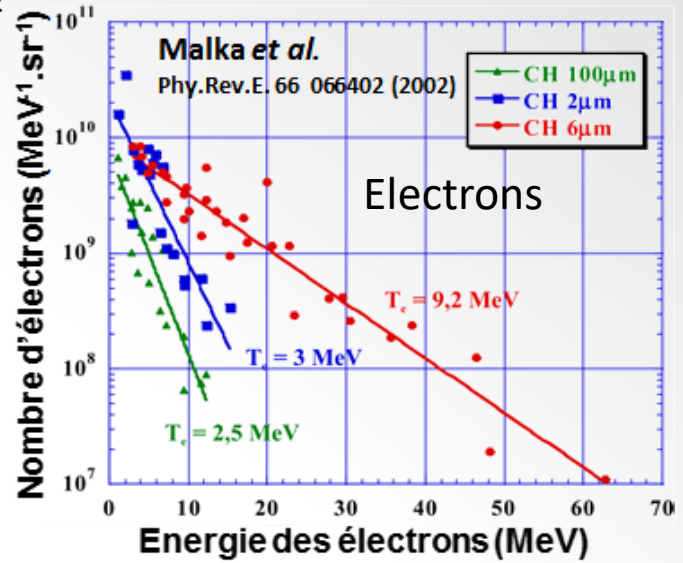
CHALLENGES TO TAKE UP

- The detection in high power laser environment
- High repetition rate lasers

The detection in high power laser environment



2.10^{19}W/cm^2



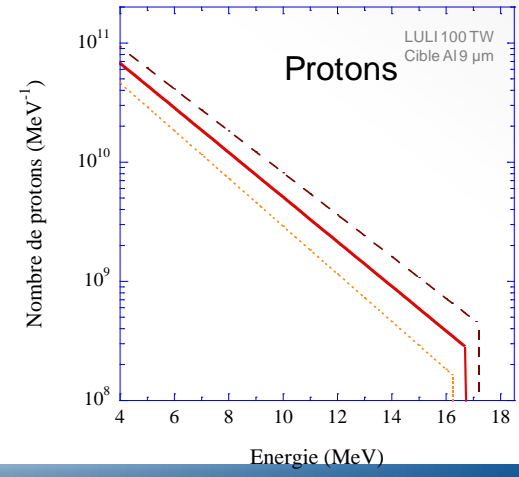
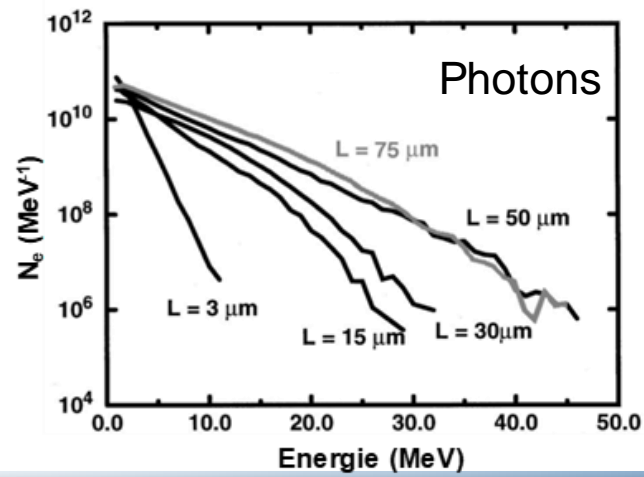
A. Macchi et al., Rev Mod Phys, vol 85, april-june 2013

✓ Physical signals in detectors :

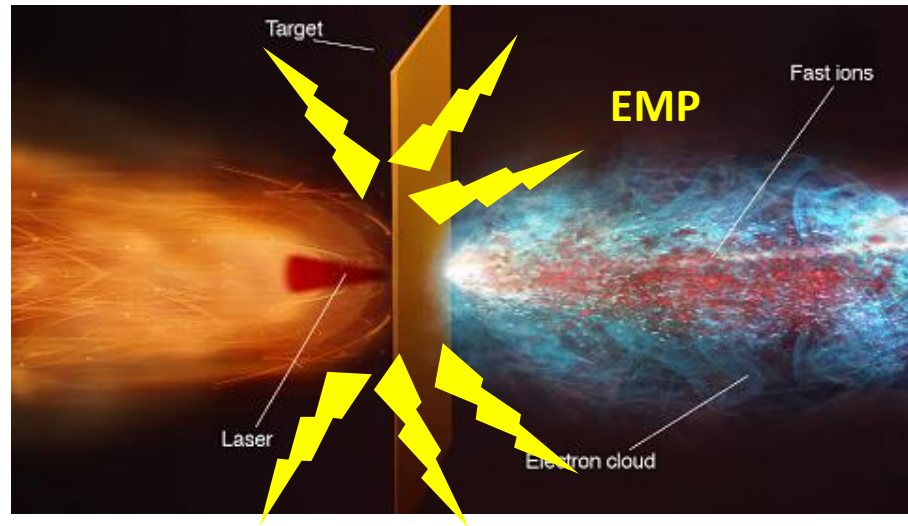
$E \sim 0.1 \text{ J } (\sim 10^{12} \text{ MeV})$ in few ns

$P \sim 100 \text{ MW}$

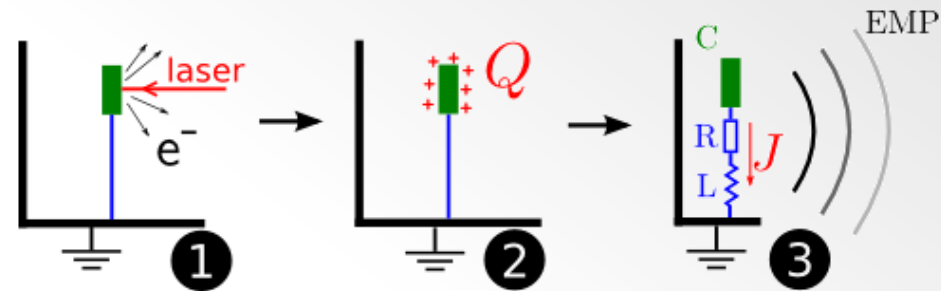
$\Phi \sim 10^8 \text{ W/cm}^2$



The detection in high power laser environment



J.-L. Dubois, et al., Phys. Rev. E 89, 013102 (2014).



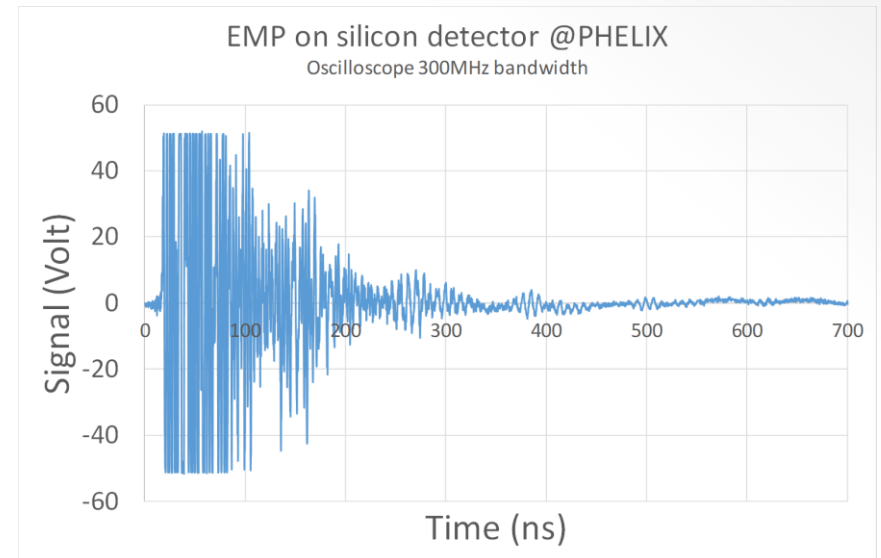
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$P \sim 100 \text{ MW}$

$\Phi \sim 10^8 \text{ W/cm}^2$

✓ Electro Magnetic Pulse (EMP) Susceptibility



The detection in high power laser environment

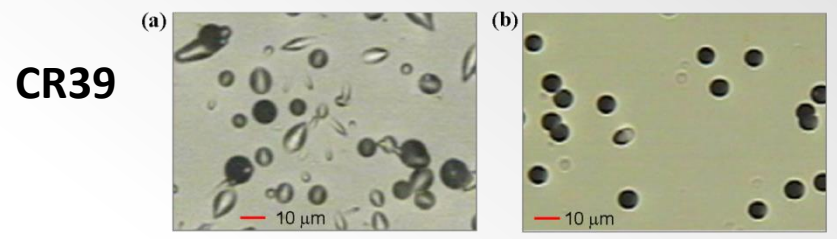
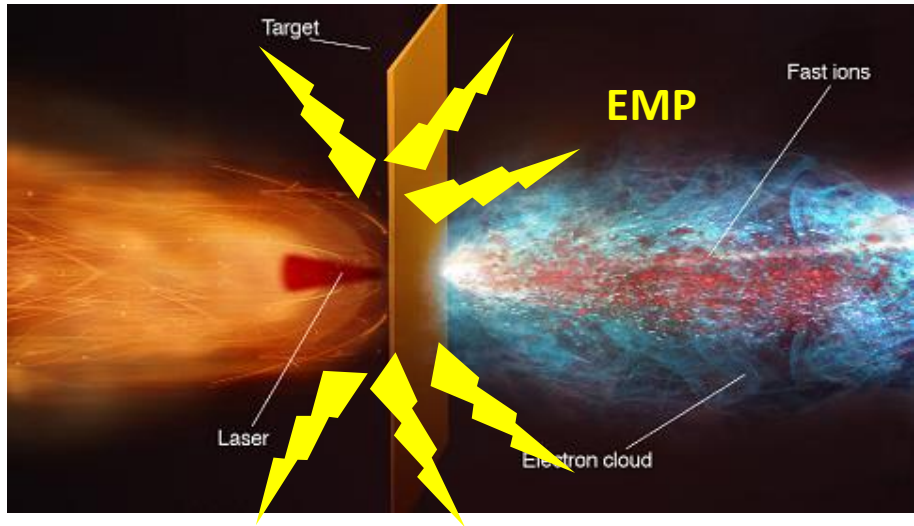


Fig. 3. Typical pictures of etched track pits caused by two different particles. (a) 5.4-MeV α particles (b) \sim 0.9-MeV protons.
 JY.LEE et al. ; Journal of the Korean Physical Society, Vol. 51, No. 1, July 2007

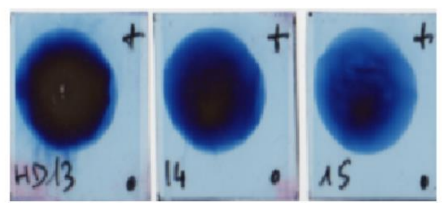
✓ Physical signals in detectors :

$E \sim 0.1 \text{ J}$ ($\sim 10^{12} \text{ MeV}$) in few ns
 $P \sim 100 \text{ MW}$
 $\Phi \sim 10^8 \text{ W/cm}^2$

➔ Use passive detectors

✓ Electro Magnetic Pulse (EMP) Susceptibility

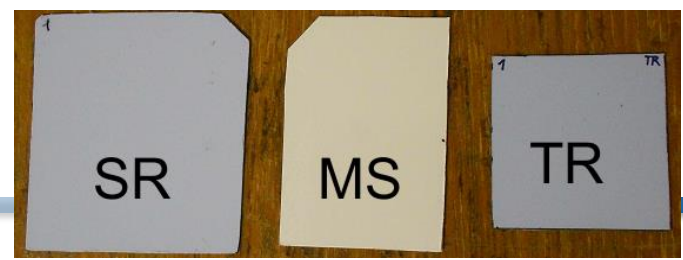
RCF



Gafchromic ©

Optical Density measurement

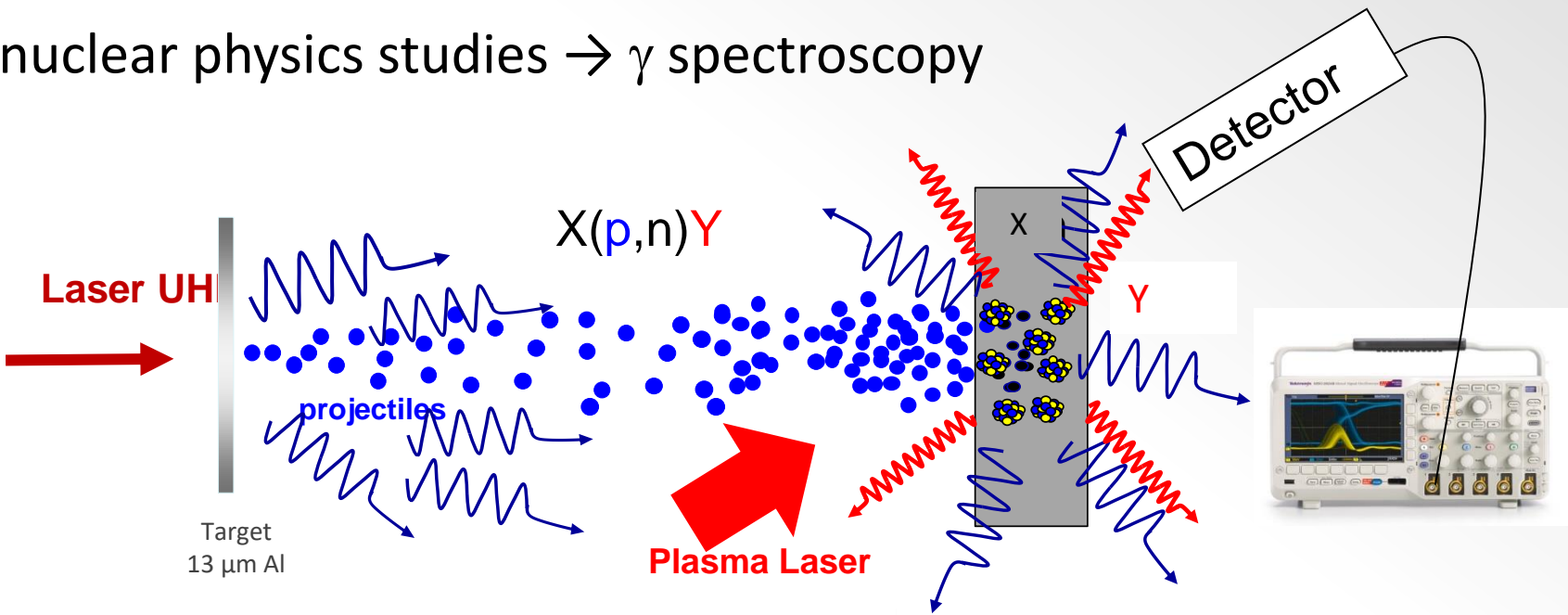
IP



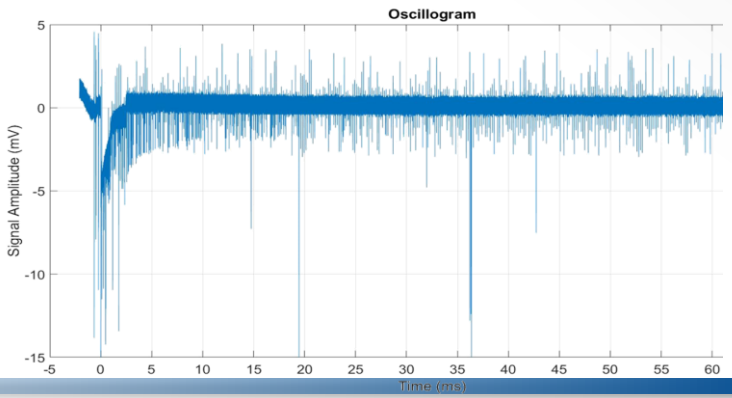
Fuji ©

The detection in high power laser environment

For nuclear physics studies → γ spectroscopy



Digitalisation of all the signals :
1 Gs/s ; 14 bits ; during ~1 s



See E. Atukpor poster

The detection in high power laser environment

For nuclear physics studies \rightarrow γ spectroscopy

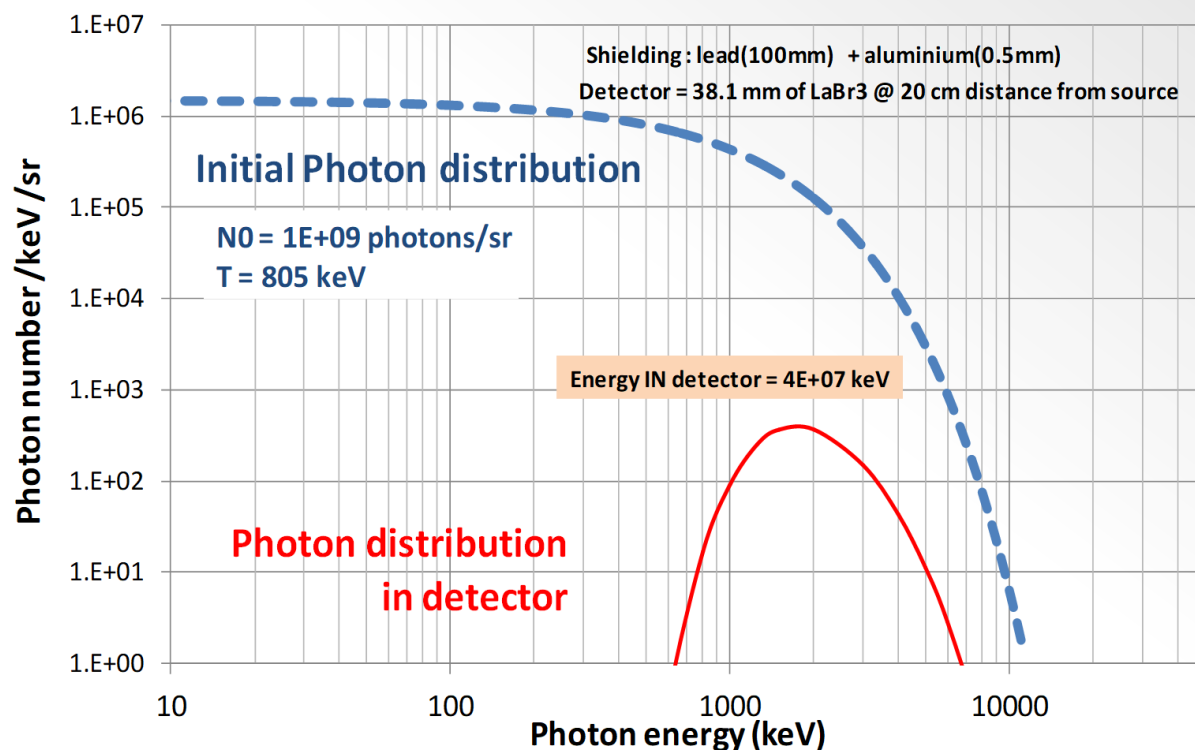
- X ray Shielding ?

- ✓ A lot of soft X rays

- ✓ Still 40 GeV energy deposition in LaBr₃ through 10 cm thick Pb shield

- ✓ A 10 μ m diameter hole @20 cm let pass through 1MeV deposit in a LaBr₃ detector

\rightarrow We need a sealed shielding , but not compatible with detection of few photon detection of \sim 100s of keV



The detection in high power laser environment

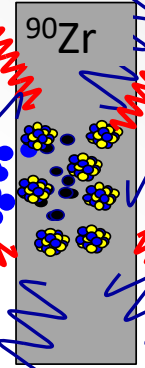
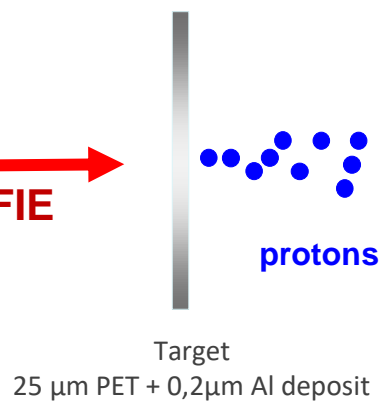
A test experiment

Laser	Energy (Joules)	Pulse duration (ps)	Power (TW)	focal (μm)	Intensity (W/cm ²)
LULI Elfie	20	350	57	25	1.E+19

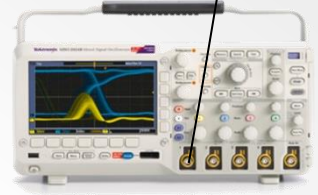


ELFIE

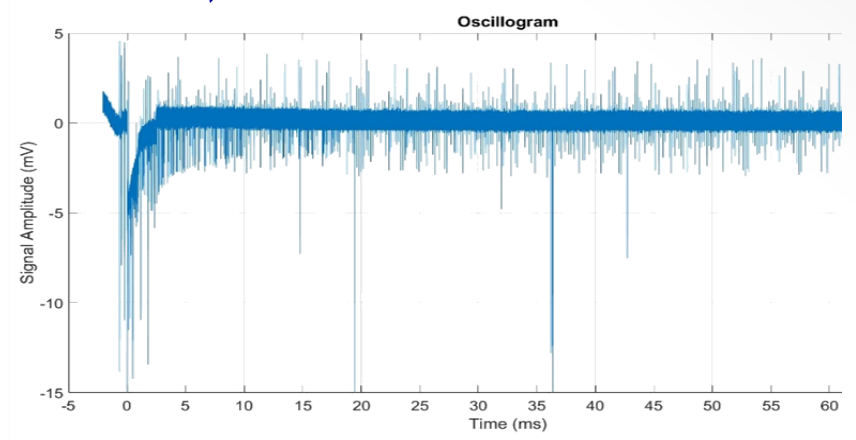
2014



LaBr₃+PMT



Level Energy and Spin	Half life	γ transition Energy
0 keV ; 8+	14,6 h	/
122.37 keV ; 6+	63 μs	122.37 keV
382.01 keV ; 1+	6.19 ms	257.34 keV



The detection in high power laser environment

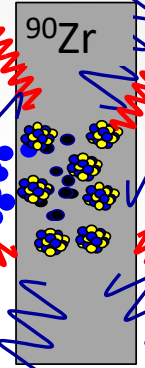
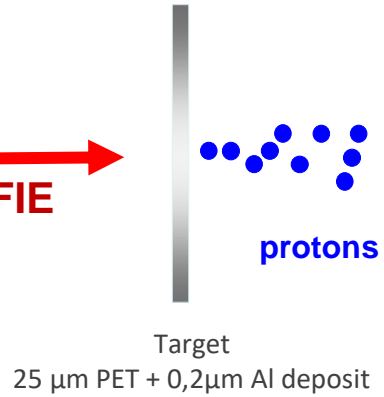
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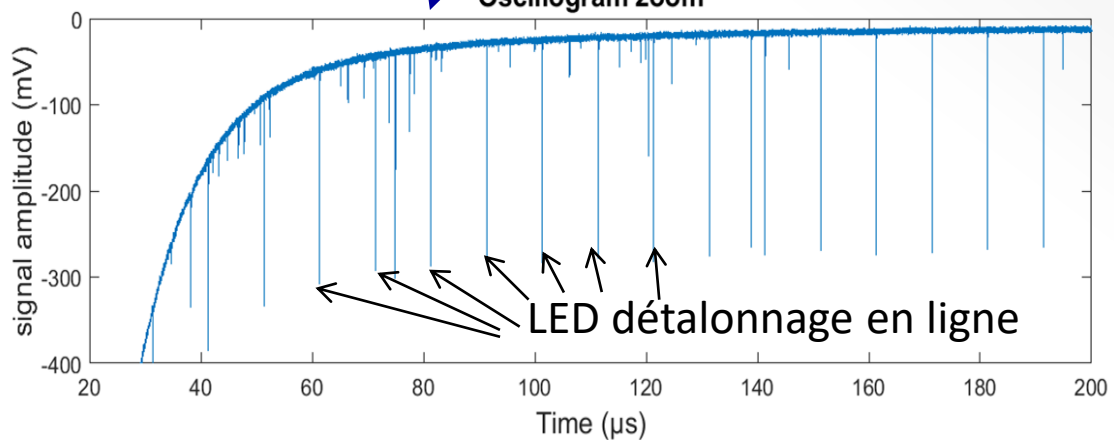
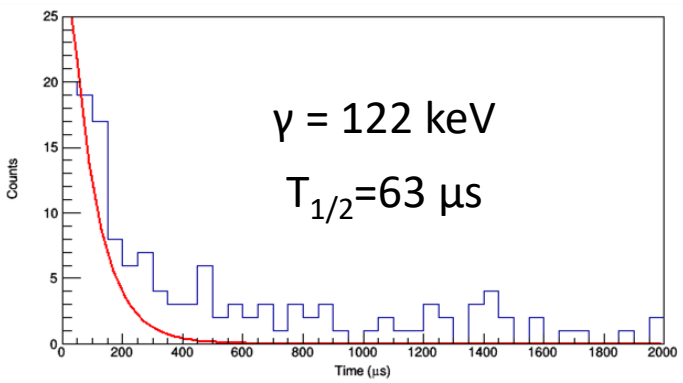
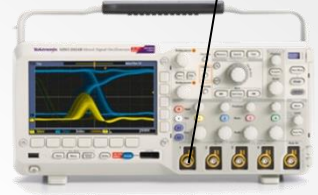


ELFIE

2016

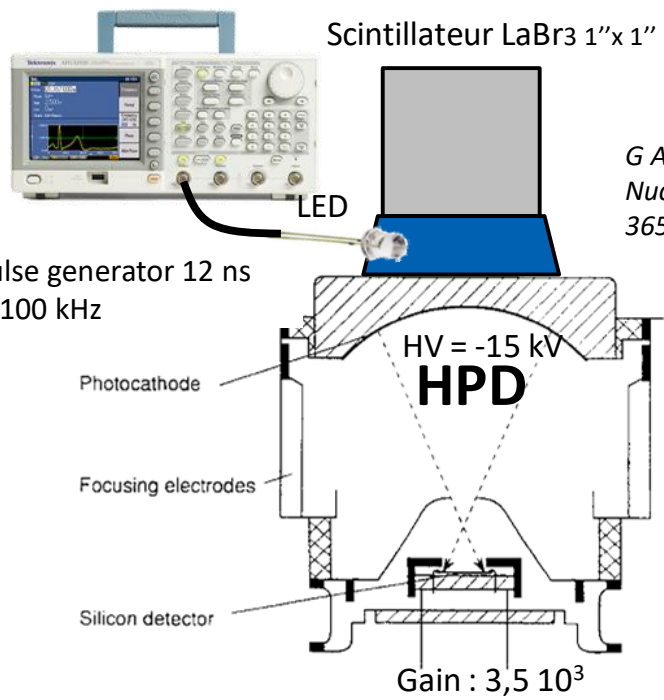


LaBr₃+HPD

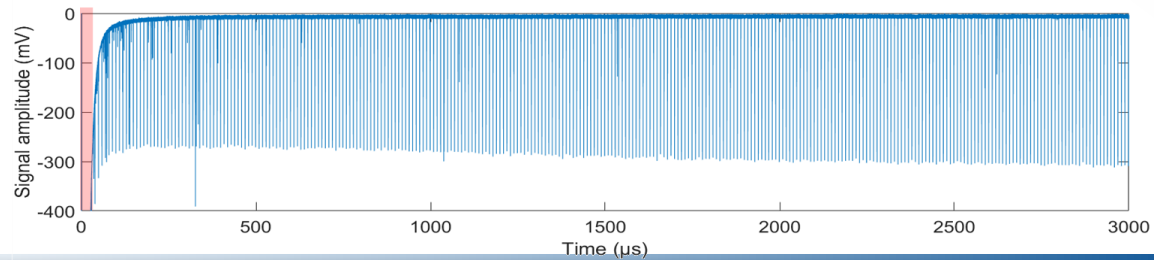
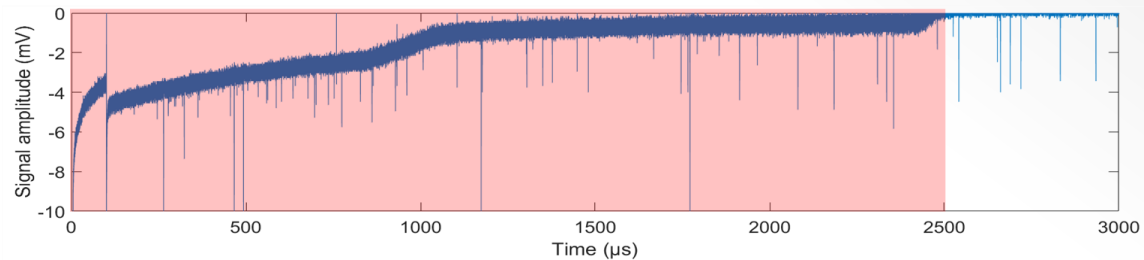
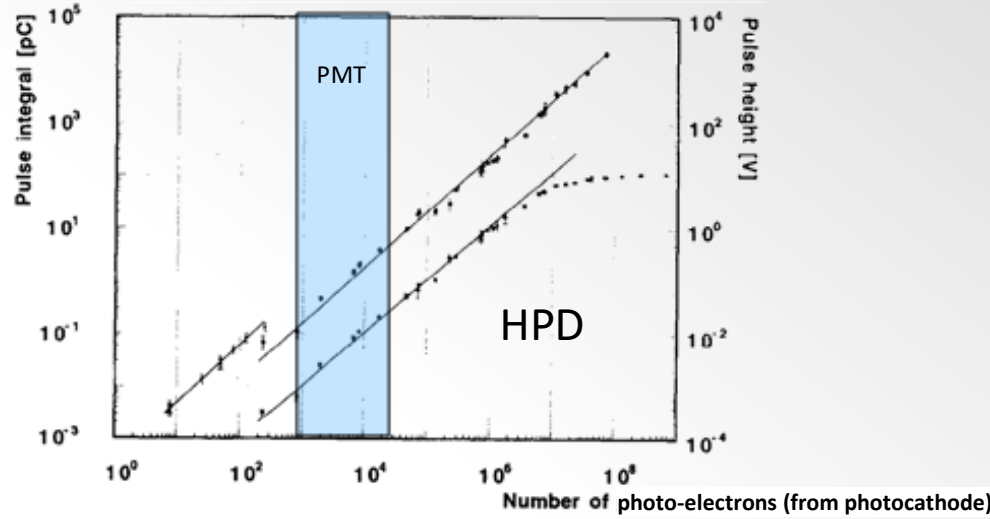


The detection in high power laser environment

A test experiment



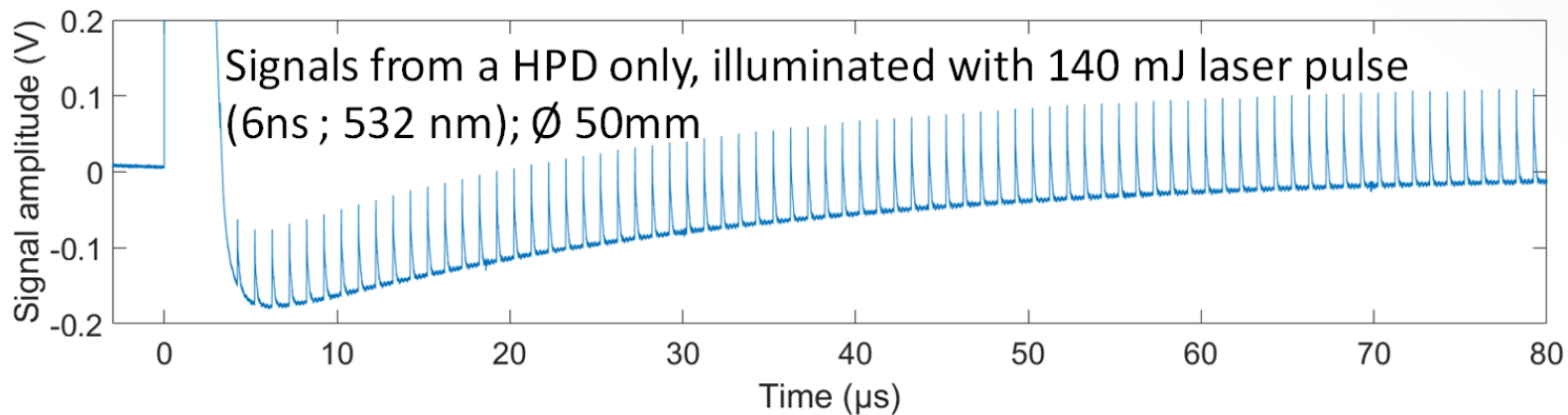
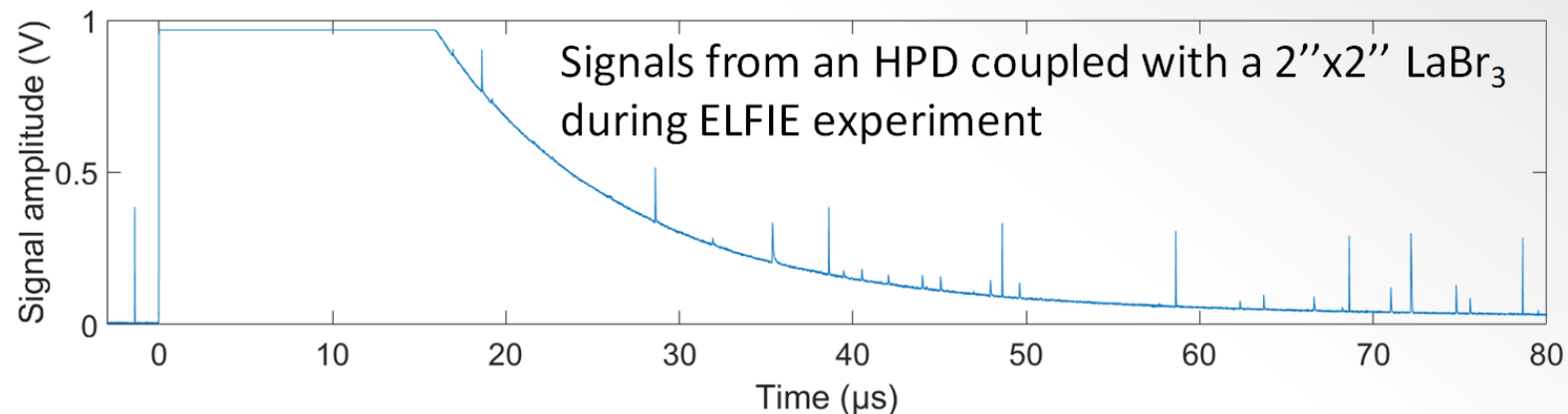
G Anzivino et al.,
Nucl.Instrum.Meth.A
365 (1995) 76-82



From 2500 μs recovery time to 50 μs

The detection in high power laser environment

Origin of long dead time?



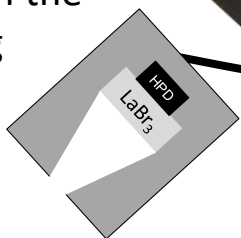
Afterglow is mainly responsible of long dead time

The detection in high power laser environment

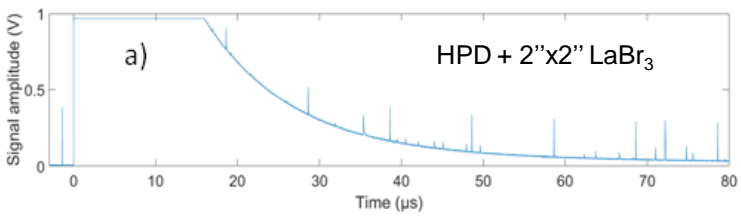
We are doing Instrumental research :
(E. Atukpor Thesis)

Scintillators

Filtering the light emission from the LaBr₃ ou CeBr₃ to remove long time component

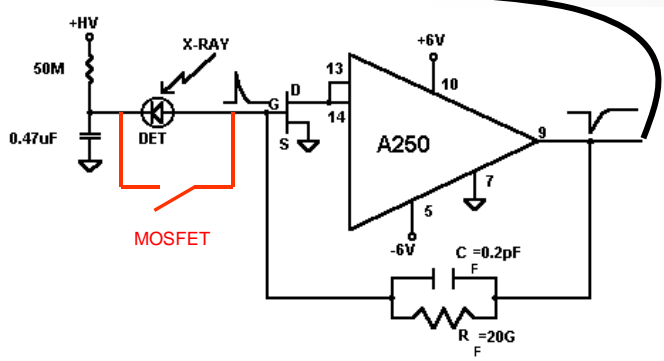


High repetition rate digitizer



Semiconductors

Shortcut the detector during the X flash to protect amplifiers



High repetition rate lasers

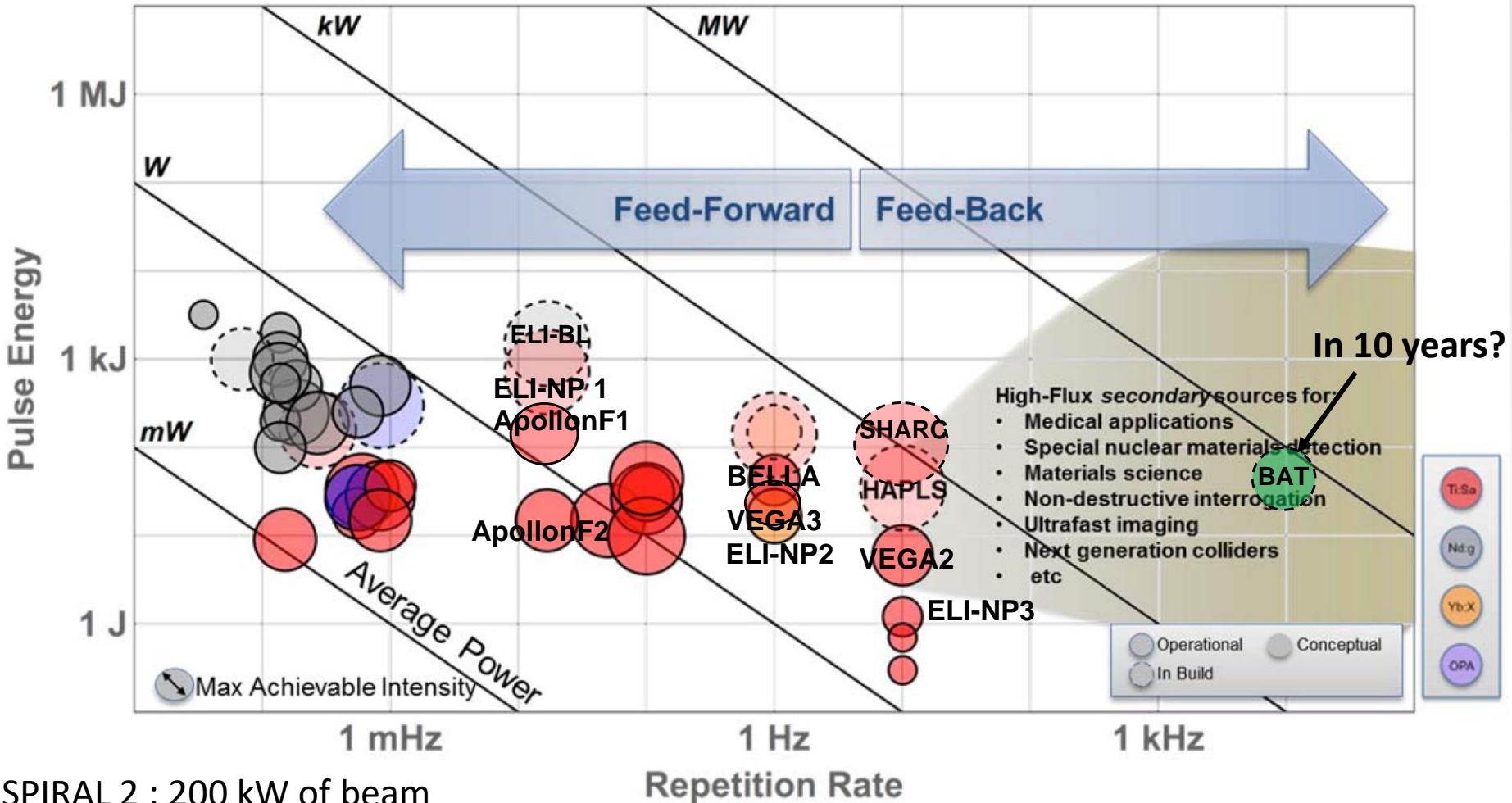
High energy and long pulse lasers for the production of hot ($T \sim 0.1-1$ keV) and dense ($\sim 10^{-2}$ g/cm³) plasmas

Laser	Localisation	Energie (Joules)	Durée du pulse (fs)	Puissance (TW)	Intensité (W/cm ²)	Cadence (Hz)	a0
CENBG	Bordeaux	1	9000000	1,E-04	1E+13	10	0,003
Apollon F3	Paris-Saclay	200	1000000	0,2	1E+15	0,02	0,0
PHLIX long	Darmstadt	1 000	5000000	0,2	1E+15	0,0002	0,0
ELI-BL ATON L4	Prague	1 500	1000000	1,5	8E+15	0,02	0,1
LULI 2000	Paris-Saclay	1600	1500000	1,1	1E+16	0,0001	0,0
MegaJoule	Bordeaux	1 300 000	3200000	406	6E+16	0,000006	0,2
Eclipse (upgrade)	Bordeaux	1,5	30	50	3,E+19	1	3,7
ELI-BL HAPLS	Prague	30	30	1000	3E+20	10	16,0
VEGA3	Salamanque	30	30	1000	1E+21	1	31,9
PICO 2000	Paris-Saclay	60	1000	60	5E+19	0,0003	4,6
PHLIX short	Darmstadt	130	500	260	4E+20	0,0002	18,1
Apollon F1	Paris-Saclay	150	15	10000	1E+22	0,02	76,7
TITAN	Livermore	210	5000	42	2E+19	0,0006	4,4
VULCAN	Oxford	500	500	1000	1E+21	0,0002	25,5

Intense lasers for particle acceleration :

- Electrons (from the 90s); $E_{\max} e^- \sim \text{GeV}$ currently
- Ions (from the 2000's); $E_{\max} \text{ protons} \sim 100 \text{ MeV}$ currently

High repetition rate lasers



SPIRAL 2 : 200 kW of beam

Conclusions

- Lasers: unique tools to study nuclear properties in extreme condition: it creates both targets and projectiles
- Lots of teams working on Laser-plasma acceleration ; very few from accelerators and nuclear physics community. Accessible application is radio-isotope production
- Evaluation of nuclear excitations rates in plasma is complex
 - ▶ Description of the plasma
 - ▶ Accurate characterization of the nuclear and atomic transitions required
- $^{84\text{m}}\text{Rb}$: a good candidate to evidence nuclear excitations in plasma
 - ▶ Nuclear transition characterised (2 experiments and shell model calculations)
 - ▶ Atomic physics in plasma well described and experimentally validated
 - ▶ Estimation uncertainties too high, need experimental evidence : ELI-BL with two laser beams
- The main issue remains the detection in such perturbed environment